# ON THE STABILITY OF RADIAL SOLUTIONS TO AN ANISOTROPIC GINZBURG-LANDAU EQUATION 

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#### Abstract

We study the linear stability of entire radial solutions $u\left(r e^{i \theta}\right)=f(r) e^{i \theta}$, with positive increasing profile $f(r)$, to the anisotropic Ginzburg-Landau equation $$
-\Delta u-\delta\left(\partial_{x}+i \partial_{y}\right)^{2} \bar{u}=\left(1-|u|^{2}\right) u, \quad-1<\delta<1,
$$ which arises in various liquid crystal models. In the isotropic case $\delta=0$, Mironescu showed that such solution is nondegenerately stable. We prove stability of this radial solution in the range $\delta \in\left(\delta_{1}, 0\right]$ for some $-1<\delta_{1}<0$, and instability outside this range. In strong contrast with the isotropic case, stability with respect to higher Fourier modes is not a direct consequence of stability with respect to lower Fourier modes. In particular, in the case where $\delta \approx-1$, lower modes are stable and yet higher modes are unstable.


## 1. Introduction

Given $\delta \in(-1,1)$ and $u: \mathbb{R}^{2} \rightarrow \mathbb{C}$, we consider the anisotropic energy

$$
\begin{equation*}
\mathfrak{E}[u]=\int_{\mathbb{R}^{2}} \frac{1}{2}|\nabla u|^{2}+\frac{\delta}{2} \operatorname{Re}\left\{\left(\partial_{\eta} \bar{u}\right)^{2}\right\}+\frac{1}{4}\left(1-|u|^{2}\right)^{2} \mathrm{~d} x, \quad \text { where } \partial_{\eta}=\partial_{x}+i \partial_{y} . \tag{1}
\end{equation*}
$$

Minimizers and stable critical points of $\mathfrak{E}$ are relevant in describing 2D point defects (or 3D straight-line defects) in some liquid crystal configurations (e.g. smectic- $C^{*}$ thin films (4) and nematics close to the Fréedericksz transition [2]). This energy can also be viewed as a toy model to understand intricate phenomena triggered by elastic anisotropy in the more complex Landau-de Gennes energy [11.

Remark 1.1. The anisotropic term $\operatorname{Re}\left\{\left(\partial_{\eta} \bar{u}\right)^{2}\right\}$ can be rewritten as

$$
\operatorname{Re}\left\{\left(\partial_{\eta} \bar{u}\right)^{2}\right\}=(\nabla \cdot u)^{2}-(\nabla \times u)^{2},
$$

so that, in view of the identity $|\nabla u|^{2}=(\nabla \cdot u)^{2}+(\nabla \times u)^{2}-2 \operatorname{det}(\nabla u)$, energy (1) differs from

$$
\widetilde{\mathfrak{E}}[u]=\int \frac{k_{s}}{2}(\nabla \cdot u)^{2}+\frac{k_{b}}{2}(\nabla \times u)^{2}+\frac{1}{4}\left(1-|u|^{2}\right)^{2}, \quad k_{s}=1+\delta, k_{b}=1-\delta,
$$

only by the integral of the null Lagrangian $\operatorname{det}(\nabla u)$. This is precisely the form that appears in [4] where minimizers of

$$
\begin{equation*}
\widetilde{\mathfrak{E}}_{\varepsilon}[u]=\int_{\Omega} \frac{k_{s}}{2}(\nabla \cdot u)^{2}+\frac{k_{b}}{2}(\nabla \times u)^{2}+\frac{1}{4 \varepsilon^{2}}\left(1-|u|^{2}\right)^{2} \tag{2}
\end{equation*}
$$

are investigated in the limit as $\varepsilon \rightarrow 0^{+}$in a bounded planar domain $\Omega$.

Critical points of $\mathfrak{E}$ are solutions of the Euler-Lagrange equation

$$
\begin{align*}
& \mathfrak{L}_{\delta} u=\left(|u|^{2}-1\right) u \quad \text { in } \mathbb{R}^{2} \\
& \mathfrak{L}_{\delta} u:=\Delta u+\delta \partial_{\eta \eta} \bar{u} . \tag{3}
\end{align*}
$$

We are interested in symmetric solutions of the form

$$
\begin{equation*}
u\left(r e^{i \theta}\right)=f(r) e^{i \alpha} e^{i \theta} \quad \text { for some } \alpha \in \mathbb{R} \tag{4}
\end{equation*}
$$

with a radial profile $f(r)$ satisfying

$$
\begin{equation*}
f(0)=0, \quad \lim _{r \rightarrow+\infty} f(r)=1, \quad|f(r)|>0 \quad \forall r \in(0, \infty) \tag{5}
\end{equation*}
$$

Formally, one can always look for solutions of (3) in the form (4) (as a consequence of the $O(2)$-invariance of $\mathfrak{E})$, and $f$ must solve

$$
T f+\delta e^{-2 i \alpha} T \bar{f}=\left(|f|^{2}-1\right) f, \quad T=\frac{d^{2}}{d r^{2}}+\frac{1}{r} \frac{d}{d r}-\frac{1}{r^{2}} .
$$

At this point we see a fundamental difference with respect to the isotropic case $\delta=0$. If $\delta=0$, one can find solutions as above for a real-valued function $f$, which moreover does not depend on $\alpha$. In the anisotropic case $\delta \neq 0$, as remarked in [2], the function $f$ can be real-valued only if $\alpha \equiv 0$ modulo $\pi / 2$. In that case, the existence and uniqueness of a solution satisfying (5) follows from the case $\delta=0$ (see [1; 6]). Otherwise, the function $f$ must be complex valued.

Remark 1.2. Another difference with respect to the isotropic case is that for $\delta \neq 0$ the Ansatz $u\left(r e^{i \theta}\right)=f(r) e^{i m \theta}$ cannot provide a solution when the winding number $m$ is $\neq 1$.

In [2], the core energies of the two symmetric solutions corresponding to $\alpha=0, \pi / 2$ are compared, to find that the lowest energy corresponds to $\alpha=0$ for $\delta<0$ and $\alpha=\pi / 2$ for $\delta>0$. In view of Remark 1.1 this is consistent with the fact that $\nabla \times e^{i \theta}=0$, while $\nabla \cdot i e^{i \theta}=0$; indeed, for $\delta<0$ the energy $\widetilde{\mathfrak{E}}[u]$ in Remark 1.1 penalizes more strongly the term $(\nabla \times u)^{2}$ than the term $(\nabla \cdot u)^{2}$, since in this case $k_{b}=1-\delta>k_{s}=1+\delta$. In [4, Proposition 3.1] the authors use this to show that minimizers of (2) behave like $e^{i \alpha} e^{i \theta}$ around point defects, with $\alpha \equiv 0$ (resp. $\pi / 2$ ) modulo $\pi$ if $\delta<0$ (resp. $\delta>0$ ). These results tell us, for $\delta \neq 0$, which one is the minimizing behavior at infinity.

Here, in contrast, we fix the far-field behavior and investigate the local stability of radial solutions with respect to compactly supported perturbations. For the isotropic case $\delta=0$, this study has been performed in [12] (see also [5), and the radial solution is stable. In the anisotropic situation $\delta \neq 0$ we find that the corresponding symmetric solution stays stable for negative $\delta$ close to zero and it loses stability for $\delta$ either positive or close to minus one (see Theorem 1.3 for precise statements).

It can be readily seen that the case $\alpha=\pi / 2$ corresponds to $\alpha=0$, after changing the sign of $\delta$. Accordingly, we only treat the case where $\alpha=0$. That is, we investigate the linear stability of solutions $u$ of the form

$$
\begin{equation*}
u_{\mathrm{rad}}^{\delta}(r, \theta)=f(r) e^{i \theta}, \quad f:(0,+\infty) \rightarrow(0,+\infty) \quad \text { with } \quad f(0)=0, \quad \lim _{r \rightarrow+\infty} f(r)=1 \tag{6}
\end{equation*}
$$

Let us note that the equation satisfied by $u_{\text {rad }}^{\delta}$, (3), reduces to the following ODE for $f$

$$
\begin{equation*}
(1+\delta) T f=\left(f^{2}-1\right) f, \quad T=\frac{d^{2}}{d r^{2}}+\frac{1}{r} \frac{d}{d r}-\frac{1}{r^{2}} \tag{7}
\end{equation*}
$$

As pointed out in [2], the rescaling of the variable by $(1+\delta)^{\frac{1}{2}}$ simplifies (7) to the standard ODE corresponding to the isotropic case $\delta=0$. Whence, existence and uniqueness of $f$ follow from [1,6]. Moreover, it is known that $f$ takes values in $(0,1)$ and is strictly increasing.

The second variation of the energy $\mathfrak{E}$ around $u_{\mathrm{rad}}^{\delta}$ is the quadratic form

$$
\begin{align*}
\mathfrak{Q}_{\mathrm{rad}}^{\delta}[v] & =\int_{\mathbb{R}^{2}}|\nabla v|^{2}+\delta \operatorname{Re}\left\{\left(\partial_{\eta} \bar{v}\right)^{2}\right\}-\left(1-\left|u_{\mathrm{rad}}^{\delta}\right|^{2}\right)|v|^{2}+2\left(u_{\mathrm{rad}}^{\delta} \cdot v\right)^{2} \mathrm{~d} x \\
& =\int_{\mathbb{R}^{2}}|\nabla v|^{2}+\delta \operatorname{Re}\left\{\left(\partial_{\eta} \bar{v}\right)^{2}\right\}-\left(1-f^{2}\right)|v|^{2}+2 f^{2}\left(e^{i \theta} \cdot v\right)^{2} \mathrm{~d} x \tag{8}
\end{align*}
$$

associated to the linear operator obtained by linearizing (3) around $u_{\text {rad }}^{\delta}$ :

$$
\mathcal{L}\left(u_{\mathrm{rad}}^{\delta}\right)[v]=-\mathfrak{L}_{\delta} v-\left(1-\left|u_{\mathrm{rad}}^{\delta}\right|^{2}\right) v+2\left(u_{\mathrm{rad}}^{\delta} \cdot v\right) u_{\mathrm{rad}}^{\delta}
$$

where $u \cdot v:=\operatorname{Re}\{u \bar{v}\}$ denotes the standard inner product of complex-valued functions.
Taking into account the asymptotic expansion $f(r)=1+O\left(r^{-2}\right)$ as $r \rightarrow \infty$ (see [1,6), it follows that the energy space of $\mathfrak{Q}_{\text {rad }}^{\delta}$ naturally corresponds to

$$
\mathcal{H}:=\left\{v \in H_{l o c}^{1}\left(\mathbb{R}^{2}\right): \int_{\mathbb{R}^{2}}|\nabla v|^{2}+\frac{1}{r^{2}}|v|^{2}+\left(e^{i \theta} \cdot v\right)^{2} \mathrm{~d} x<+\infty\right\}
$$

Also, the translational invariance of $\mathfrak{E}$ readily provides two elements of $\mathcal{H}$ at which $\mathfrak{Q}_{\text {rad }}^{\delta}$ vanishes, namely

$$
\partial_{x} u_{\mathrm{rad}}^{\delta}=e^{i \theta}\left(f^{\prime} \cos \theta-i \frac{f}{r} \sin \theta\right), \quad \quad \partial_{y} u_{\mathrm{rad}}^{\delta}=e^{i \theta}\left(f^{\prime} \sin \theta+i \frac{f}{r} \cos \theta\right)
$$

and the linear space they generate is denoted by

$$
K_{0}=\operatorname{span}\left\{\partial_{x} u_{\mathrm{rad}}^{\delta}, \partial_{y} u_{\mathrm{rad}}^{\delta}\right\}
$$

Our main result shows that the symmetric solution $u_{\mathrm{rad}}^{\delta}$ is stable when $\delta \leq 0$ is small, and unstable otherwise:

Theorem 1.3. Let $u_{\mathrm{rad}}^{\delta}$ denote the radial solution (6) of the anisotropic Ginzburg-Landau equation (3), and let $\mathfrak{Q}_{\text {rad }}^{\delta}$ denote the quadratic form (8) associated to the energy $\mathfrak{E}$ around $u_{\mathrm{rad}}^{\delta}$. Then, there exists a unique number $\delta_{1} \in(-1,0)$ such that

- for every $\delta \in\left(\delta_{1}, 0\right], u_{\mathrm{rad}}^{\delta}$ is nondegenerately stable: namely,

$$
\mathfrak{Q}_{\mathrm{rad}}^{\delta}[v]>0 \quad \text { for all } v \in H \backslash K_{0}
$$

- for every $\delta \in\left(-1, \delta_{1}\right) \cup(0,1)$, $u_{\mathrm{rad}}^{\delta}$ is linearly unstable: namely,

$$
\mathfrak{Q}_{\mathrm{rad}}^{\delta}[v]<0 \quad \text { for some } v \in H
$$

Remark 1.4. The most relevant range from the stand point of physics is $\delta \in(-1,0]$ since for $\delta>0$ the far-field behavior corresponding to $\alpha=0$ is non-minimizing, and this translates here into instability of the radial solution.

Remark 1.5. In the stability range $\delta \in\left(\delta_{1}, 0\right.$ ], a contradiction argument as in [5, Lemma 3.1] provides a coercivity estimate of the form

$$
\mathfrak{Q}_{\mathrm{rad}}^{\delta}[v] \geq C(\delta) \int_{\mathbb{R}^{2}}|\nabla v|^{2} \mathrm{~d} x \quad \forall v \in K_{0}^{\perp}: \int_{\mathbb{S}^{1}}\left(i e^{i \theta}\right) \cdot v\left(r e^{i \theta}\right) d \theta=0 \quad \forall r>0
$$

where $\perp$ denotes orthogonality in $\mathcal{H}$. Using this coercivity for $\delta=0$, one can deduce stability for small negative $\delta$ via a relatively simple perturbation argument, combined with properties of the lower modes in §3. Instead, we will give a more quantitative proof, which provides an explicit range for stability: we deduce that $\delta_{1} \leq-1 / \sqrt{5}$.

Our proof of Theorem 1.3 follows the general strategy of [12]: we decompose $v$ into Fourier modes

$$
v=e^{i \theta} \sum_{n \in \mathbb{Z}} w_{n}(r) e^{i n \theta} .
$$

and we are led to studying the sign of $\mathfrak{Q}_{\text {rad }}^{\delta}$, separately, for each mode

$$
e^{i \theta}\left(w_{n}(r) e^{i n \theta}+w_{-n}(r) e^{-i n \theta}\right)
$$

As in [12, the lower modes $n=0$ and $n=1$ play a special role. They can be studied via an appropriate decomposition already used in [12 (see also [5]). For any $\delta \in(-1,0]$ we find that these lower modes are stable, while for $\delta>0$ the mode corresponding to $n=0$ is unstable.

A major difference of the present work compared to [12] (or similar results in [8-10]) pertains to the higher modes $n \geq 2$. In contrast with the cited works, stability for the higher modes is not an obvious consequence of stability for the lower modes. More precisely in the isotropic case we have

$$
\mathfrak{Q}_{\mathrm{rad}}^{0}\left[e^{i \theta}\left(w_{+}(r) e^{i n \theta}+w_{-}(r) e^{-i n \theta}\right)\right] \geq \mathfrak{Q}_{\mathrm{rad}}^{0}\left[e^{i \theta}\left(w_{+}(r) e^{i \theta}+w_{-}(r) e^{-i \theta}\right)\right] \quad \forall n \geq 1,
$$

but for $\delta \neq 0$ this is not valid anymore, see (16). This feature is new and specific to the anisotropic case $\delta \neq 0$. Our strategy to study the sign of these higher modes is based on the same decomposition used for $n=1$, and a careful balance of the contributions of additional terms, which end up causing instability for $\delta \approx-1$.

The article is organized as follows. In Section 2 we recall the splitting property of the quadratic form $\mathfrak{Q}_{\text {rad }}^{\delta}$ with respect to Fourier expansion. In Section 3 we study the stability of lower modes, and in Section 4 the instability of higher modes. In Section 5 we give the proof of Theorem 1.3. In addition, we included Appendix A to recall the details of the decomposition used to study the lower modes, adapted to our notations.

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## 2. Fourier splitting

Recall that $f(r)=f_{0}\left((1+\delta)^{-\frac{1}{2}} r\right)$ where $f_{0}$ is the classical Ginzburg-Landau vortex profile corresponding to the case $\delta=0$. That is, the unique solution of

$$
\begin{equation*}
f_{0}^{\prime \prime}+\frac{1}{r} f_{0}^{\prime}-\frac{1}{r^{2}} f_{0}=-\left(1-f_{0}^{2}\right) f_{0}, \quad f_{0}>0 \text { on }(0,+\infty), \quad f_{0}(0)=0, \quad \lim _{r \rightarrow+\infty} f_{0}(r)=1 \tag{9}
\end{equation*}
$$

We rescale variables and consider $\mathcal{Q}^{\delta}[v]=\mathfrak{Q}_{\text {rad }}^{\delta}[\tilde{v}]$ where $\tilde{v}(\tilde{x})=v\left((1+\delta)^{-\frac{1}{2}} \tilde{x}\right)$, so that

$$
\begin{equation*}
\mathcal{Q}^{\delta}[v]=\int_{\mathbb{R}^{2}}|\nabla v|^{2}+\delta \operatorname{Re}\left\{\left(\partial_{\eta} \bar{v}\right)^{2}\right\}+(1+\delta)\left\{2 f_{0}^{2}\left(e^{i \theta} \cdot v\right)^{2}-\left(1-f_{0}^{2}\right)|v|^{2}\right\} \mathrm{d} x \tag{10}
\end{equation*}
$$

which corresponds to the second variation of the appropriately rescaled energy around $u_{\text {rad }}^{0}$. Following 12 we decompose $v$ using Fourier series, as

$$
\begin{equation*}
v=e^{i \theta} w=e^{i \theta} \sum_{n \in \mathbb{Z}} w_{n}(r) e^{i n \theta} \tag{11}
\end{equation*}
$$

where we have conveniently shifted the index $n-1 \mapsto n$.
This decomposition provides a "diagonalization" of the linearized operator:
Lemma 2.1. The quadratic form (10) splits as

$$
\mathcal{Q}^{\delta}[v]=\mathcal{Q}^{\delta}\left[w_{0}(r) e^{i \theta}\right]+\sum_{n \geq 1} \mathcal{Q}^{\delta}\left[e^{i \theta}\left(w_{n}(r) e^{i n \theta}+w_{-n}(r) e^{-i n \theta}\right)\right]
$$

Proof of Lemma 2.1. Lemma 2.1 essentially asserts that the family of functions

$$
\begin{equation*}
w_{0}(r) e^{i \theta}, \quad\left\{e^{i \theta}\left(w_{n}(r) e^{i n \theta}+w_{-n}(r) e^{-i n \theta}\right): n \geq 1\right\} \tag{12}
\end{equation*}
$$

is orthogonal for the quadratic form $\mathcal{Q}$. This quadratic form 10 is composed of three terms. For the first term,

$$
\int_{\mathbb{R}^{2}}|\nabla v|^{2} \mathrm{~d} x
$$

the orthogonality of $\sqrt[12]{ }$ is a standard fact (recall e.g. in $\sqrt[12]{ })$. For the third term,

$$
\int_{\mathbb{R}^{2}}\left\{f_{0}^{2}\left(e^{i \theta} \cdot v\right)^{2}-\left(1-f_{0}^{2}\right)|v|^{2}\right\} \mathrm{d} x
$$

the orthogonality of $\sqrt{12})$ is proved in $\sqrt[12]]{ }$. The novelty here, with respect to $\sqrt{12]}$, concerns the anisotropic term

$$
\int_{\mathbb{R}^{2}} \operatorname{Re}\left\{\left(\partial_{\eta} \bar{v}\right)^{2}\right\} \mathrm{d} x
$$

The orthogonality of 12 for this anisotropic term, as a matter of fact, follows from the calculations in [3, §3.2]. As our notations are different, we sketch a proof here for the reader's convenience.

We compute

$$
\partial_{\eta} \bar{v}=e^{i \theta} \partial_{r} \bar{v}+\frac{i e^{i \theta}}{r} \partial_{\theta} \bar{v}=\sum_{n \in \mathbb{Z}}\left(\bar{w}_{n}^{\prime}+\frac{1+n}{r} \bar{w}_{n}\right) e^{-i n \theta}
$$

and deduce, using the orthogonality of $\left\{e^{i n \theta}\right\}$ in $L^{2}\left(\mathbb{S}^{1}\right)$,

$$
\begin{aligned}
& f_{\mathbb{S} 1} \operatorname{Re}\left\{\left(\partial_{\eta} \bar{v}\right)^{2}\right\} d \theta \\
& =\operatorname{Re}\left\{\sum_{n, m \in \mathbb{Z}}\left(\bar{w}_{n}^{\prime}+\frac{1+n}{r} \bar{w}_{n}\right)\left(\bar{w}_{m}^{\prime}+\frac{1+m}{r} \bar{w}_{m}\right) f_{\mathbb{S}^{1}} e^{-i(n+m) \theta} d \theta\right\} \\
& =\operatorname{Re}\left\{\sum_{n \in \mathbb{Z}}\left(\bar{w}_{n}^{\prime}+\frac{1+n}{r} \bar{w}_{n}\right)\left(\bar{w}_{-n}^{\prime}+\frac{1-n}{r} \bar{w}_{-n}\right)\right\} \\
& =\sum_{n \in \mathbb{Z}} \operatorname{Re}\left\{\left(\bar{w}_{n}^{\prime}+\frac{1+n}{r} \bar{w}_{n}\right)\left(\bar{w}_{-n}^{\prime}+\frac{1-n}{r} \bar{w}_{-n}\right)\right\} .
\end{aligned}
$$

This implies the announced orthogonality and completes the proof of Lemma 2.1.

According to the decomposition of Lemma (2.1), we define the quadratic forms

$$
\begin{aligned}
Q_{0}^{\delta}[\varphi] & =\frac{1}{2 \pi} \mathcal{Q}^{\delta}\left[\varphi(r) e^{i \theta}\right] & \text { for } \varphi \in \mathcal{H}_{0}, \\
Q_{n}^{\delta}[\varphi, \psi] & =\frac{1}{2 \pi} \mathcal{Q}^{\delta}\left[e^{i \theta}\left(\varphi(r) e^{i n \theta}+\psi(r) e^{-i n \theta}\right)\right] & \text { for }(\varphi, \psi) \in \mathcal{H}_{1},
\end{aligned}
$$

where $\mathcal{H}_{0}$ and $\mathcal{H}_{1}$ are the natural spaces corresponding to the conditions $\varphi(r) e^{i \theta} \in \mathcal{H}$ and $e^{i \theta}\left(\varphi(r) e^{i n \theta}+\psi(r) e^{-i n \theta}\right) \in \mathcal{H}$ for $n \geq 1$, respectively.

$$
\begin{aligned}
& \mathcal{H}_{0}=\left\{\varphi \in H_{l o c}^{1}(0, \infty): \int_{0}^{+\infty}\left(\left|\varphi^{\prime}\right|^{2}+\frac{|\varphi|^{2}}{r^{2}}+\operatorname{Re}\{\varphi\}^{2}\right) r d r<+\infty\right\} \\
& \mathcal{H}_{1}=\left\{(\varphi, \psi) \in\left(H_{l o c}^{1}(0, \infty)\right)^{2}: \int_{0}^{+\infty}\left(\left|\varphi^{\prime}\right|^{2}+\left|\psi^{\prime}\right|^{2}+\frac{|\varphi|^{2}+|\psi|^{2}}{r^{2}}+|\varphi+\bar{\psi}|^{2}\right) r d r<+\infty\right\}
\end{aligned}
$$

Remark 2.2. Using the density of smooth functions in $H_{l o c}^{1}$ and cut-off functions $\chi_{\varepsilon}$ such that $\mathbf{1}_{2 \varepsilon<r<\varepsilon^{-1}} \leq \chi_{\varepsilon}(r) \leq \mathbf{1}_{\varepsilon<r<2 \varepsilon^{-1}}$ and $\left|\chi_{\varepsilon}^{\prime}(r)\right| \leq C / r$, we see that smooth test functions with compact support in $(0, \infty)$ are dense in $\mathcal{H}_{0}$ and $\mathcal{H}_{1}$. Hence, in the sequel, we will always be able to perform calculations assuming, without loss of generality, that $\varphi$ and $\psi$ are such test functions.

The quadratic forms $Q_{0}^{\delta}$ and $Q_{n}^{\delta}$ are explicitly given by

$$
\begin{align*}
& Q_{0}^{\delta}[\varphi]= \int_{0}^{\infty}\left[\left|\varphi^{\prime}\right|^{2}+\right.  \tag{13}\\
& \frac{1}{r^{2}}|\varphi|^{2}+\delta \operatorname{Re}\left\{\left(\bar{\varphi}^{\prime}+\frac{1}{r} \bar{\varphi}\right)^{2}\right\} \\
&\left.+(1+\delta)\left\{2 f_{0}^{2}(\operatorname{Re}\{\varphi\})^{2}-\left(1-f_{0}^{2}\right)|\varphi|^{2}\right\}\right] r d r  \tag{14}\\
& Q_{n}^{\delta}[\varphi, \psi]= \int_{0}^{\infty}\left[\left|\varphi^{\prime}\right|^{2}+\right. \\
&+\left|\psi^{\prime}\right|^{2}+\frac{(1+n)^{2}}{r^{2}}|\varphi|^{2}+\frac{(1-n)^{2}}{r^{2}}|\psi|^{2} \\
&+ 2 \delta \operatorname{Re}\left\{\left(\bar{\varphi}^{\prime}+\frac{1+n}{r} \bar{\varphi}\right)\left(\bar{\psi}^{\prime}+\frac{1-n}{r} \bar{\psi}\right)\right\} \\
&\left.+(1+\delta)\left\{f_{0}^{2}|\varphi+\bar{\psi}|^{2}-\left(1-f_{0}^{2}\right)\left(|\varphi|^{2}+|\psi|^{2}\right)\right\}\right] r d r
\end{align*}
$$

Remark 2.3. For every $n \geq 1$ there is a further splitting, namely

$$
Q_{n}^{\delta}[\varphi, \psi]=Q_{n}^{\delta}[\operatorname{Re}\{\varphi\}, \operatorname{Re}\{\psi\}]+Q_{n}^{\delta}[\operatorname{Im}\{\varphi\},-\operatorname{Im}\{\psi\}]
$$

Consequently, it will be sufficient to consider real-valued test functions $\varphi, \psi$.
3. Study of the lower modes $Q_{0}^{\delta}$ and $Q_{1}^{\delta}$

We show that $Q_{0}^{\delta}$ is positive for $\delta \leq 0$, but it can become negative for $\delta>0$. In addition, we prove that $Q_{1}^{\delta}$ is nonnegative for all $\delta \in(-1,0]$.
3.1. Positivity of $Q_{0}^{\delta}$ for $\delta \in(-1,0]$. Let us recall from $(13)$ that $Q_{0}^{\delta}$ is given by

$$
\begin{aligned}
Q_{0}^{\delta}[\varphi]=\int_{0}^{\infty}\left[\left|\varphi^{\prime}\right|^{2}+\right. & \frac{1}{r^{2}}|\varphi|^{2}+\delta \operatorname{Re}\left\{\left(\bar{\varphi}^{\prime}+\frac{1}{r} \bar{\varphi}\right)^{2}\right\} \\
& \left.+(1+\delta)\left\{2 f_{0}^{2}(\operatorname{Re}\{\varphi\})^{2}-\left(1-f_{0}^{2}\right)|\varphi|^{2}\right\}\right] r d r
\end{aligned}
$$

We now introduce the quadratic form

$$
\begin{aligned}
A_{0}[\varphi] & :=Q_{0}^{0}[\varphi] \\
& =\int_{0}^{\infty}\left[\left|\varphi^{\prime}\right|^{2}+\frac{1}{r^{2}}|\varphi|^{2}+2 f_{0}^{2}(\operatorname{Re}\{\varphi\})^{2}-\left(1-f_{0}^{2}\right)|\varphi|^{2}\right] r d r
\end{aligned}
$$

It is known that $A_{0}[\varphi]>0$, unless $\varphi=0$ (see Appendix A for more details). Moreover, we have the identity

$$
\begin{aligned}
Q_{0}^{\delta}[\varphi]= & (1+\delta) A_{0}[\operatorname{Re}\{\varphi\}]+(1-\delta) A_{0}[i \operatorname{Im}\{\varphi\}]-2 \delta \int\left(1-f_{0}^{2}\right)(\operatorname{Im}\{\varphi\})^{2} r d r \\
& +\delta \int_{0}^{\infty} \frac{d}{d r}\left[(\operatorname{Re}\{\varphi\})^{2}-(\operatorname{Im}\{\varphi\})^{2}\right] d r \\
= & (1+\delta) A_{0}[\operatorname{Re}\{\varphi\}]+(1-\delta) A_{0}[i \operatorname{Im}\{\varphi\}]-2 \delta \int\left(1-f_{0}^{2}\right)(\operatorname{Im}\{\varphi\})^{2} r d r,
\end{aligned}
$$

which is valid for any $\varphi \in C_{c}^{\infty}(0, \infty)$, hence for $\varphi \in \mathcal{H}_{0}$ thanks to Remark 2.2. Since $1-f_{0}^{2} \geq 0$, we deduce the positivity of $Q_{0}^{\delta}$ for every $\delta \in(-1,0]$.
3.2. Instability for $\delta>0$. Using the formula (18) obtained for $A_{0}$ in Appendix A, we see that for any compactly supported real-valued test function $\chi$ we have

$$
Q_{0}^{\delta}\left[i f_{0} \chi\right]=(1-\delta) \int f_{0}^{2}\left(\chi^{\prime}\right)^{2} r d r-2 \delta \int\left(1-f_{0}^{2}\right) f_{0}^{2} \chi^{2} r d r
$$

Applying this to $\chi_{n}(r)=\chi_{1}(r / n)$, for a fixed test function $\chi_{1}$, and using the asymptotic expansion [1, 6:

$$
f_{0}(r)=1-\frac{1}{2 r^{2}}+O\left(r^{-4}\right) \quad \text { as } r \rightarrow \infty
$$

we see that

$$
\lim _{n \rightarrow \infty} Q_{0}^{\delta}\left[i f_{0} \chi_{n}\right]=(1-\delta) \int\left(\chi_{1}^{\prime}\right)^{2} r d r-2 \delta \int \frac{\chi_{1}^{2}}{r^{2}} r d r
$$

When $\delta>0$, this expression must be negative for some $\chi_{1}$, since Hardy's inequality is known to fail in two dimensions. Explicitly, by choosing

$$
\chi_{1}(r)=\sin (\sqrt{\lambda} \ln r) \mathbf{1}_{\left(1, e^{\pi / \sqrt{\lambda})}\right.}(r) \quad \text { for } \lambda=\frac{\delta}{1-\delta}>0
$$

we have that $\chi_{1} \in H^{1}(0, \infty)$ is compactly supported, and

$$
\lim _{n \rightarrow \infty} Q_{0}^{\delta}\left[i f_{0} \chi_{n}\right]=-\delta \int \frac{\chi_{1}^{2}}{r^{2}} r d r<0
$$

Whence, for $\delta>0$, the mode of order 0 already brings instability. This comes as no surprise as this mode corresponds to infinitesimal rotations (see Appendix A), and we know that the far-field behavior $e^{i \theta}$ is unstable: rotating this far-field behavior decreases the energy.
3.3. Positivity of $Q_{1}^{\delta}$ for $\delta \leq 0$. Recall, according to (14), that $Q_{1}^{\delta}$ is given by

$$
\begin{aligned}
Q_{1}^{\delta}[\varphi, \psi]=\int_{0}^{\infty}\left[\left|\varphi^{\prime}\right|^{2}\right. & +\left|\psi^{\prime}\right|^{2}+\frac{4}{r^{2}}|\varphi|^{2} \\
+ & 2 \delta \operatorname{Re}\left\{\left(\bar{\varphi}^{\prime}+\frac{2}{r} \bar{\varphi}\right) \bar{\psi}^{\prime}\right\} \\
& \left.+(1+\delta)\left\{f_{0}^{2}|\varphi+\bar{\psi}|^{2}-\left(1-f_{0}^{2}\right)\left(|\varphi|^{2}+|\psi|^{2}\right)\right\}\right] r d r
\end{aligned}
$$

We introduce the quadratic form $A_{1}:=Q_{1}^{0}$, namely

$$
\begin{aligned}
A_{1}[\varphi, \psi]=\int_{0}^{\infty}\left[\left|\varphi^{\prime}\right|^{2}\right. & +\left|\psi^{\prime}\right|^{2}+\frac{4}{r^{2}}|\varphi|^{2} \\
& \left.\quad+f_{0}^{2}|\varphi+\bar{\psi}|^{2}-\left(1-f_{0}^{2}\right)\left(|\varphi|^{2}+|\psi|^{2}\right)\right] r d r
\end{aligned}
$$

It is a known fact that $A_{1}$ is nonnegative on $\mathcal{H}_{1}$, and vanishes exactly at pairs $(\varphi, \psi)$ corresponding to maps $v$ which are linear combinations of $\partial_{x} u_{\mathrm{rad}}^{0}$ and $\partial_{y} u_{\mathrm{rad}}^{0}$ (see Appendix A for more details). Moreover, we have

$$
\begin{align*}
Q_{1}^{\delta}[\varphi, \psi]-(1+\delta) A_{1}[\varphi, \psi]= & -\delta \int_{0}^{\infty}\left[\left|\varphi^{\prime}\right|^{2}+\left|\psi^{\prime}\right|^{2}+\frac{4}{r^{2}}|\varphi|^{2}\right] r d r  \tag{15}\\
& +2 \delta \int_{0}^{\infty} \operatorname{Re}\left\{\left(\bar{\varphi}^{\prime}+\frac{2}{r} \bar{\varphi}\right) \bar{\psi}^{\prime}\right\} r d r \\
= & -\delta \int_{0}^{\infty}\left|\varphi^{\prime}+\frac{2}{r} \varphi-\bar{\psi}^{\prime}\right|^{2} r d r-2 \delta \int_{0}^{\infty} \frac{d}{d r}\left[|\varphi|^{2}\right] d r \\
= & -\delta \int_{0}^{\infty}\left|\varphi^{\prime}+\frac{2}{r} \varphi-\bar{\psi}^{\prime}\right|^{2} r d r,
\end{align*}
$$

for $(\varphi, \psi) \in\left(C_{c}^{\infty}(0, \infty)\right)^{2}$, hence for all $(\varphi, \psi) \in \mathcal{H}_{1}$. From this identity we infer that $Q_{1}^{\delta} \geq 0$ for every $\delta \in(-1,0]$, and equality can only occur when $v$ is a linear combination of $\partial_{x} u_{\mathrm{rad}}^{0}$ and $\partial_{y} u_{\mathrm{rad}}^{0}$.

## 4. Study of the higher modes $Q_{n}^{\delta}$ for $n \geq 2$

4.1. Positivity of $Q_{n}^{\delta}$ for $n \geq 2$ and $\delta \in[-1 / \sqrt{5}, 0]$. Let us recall: in the isotropic case, the positivity of $Q_{n}^{\delta}$ (any $n \geq 2$ ) is a consequence of the fact that $Q_{n}^{0} \geq Q_{1}^{0}$. Here, from the
definition (14) of $Q_{n}^{\delta}$, we have

$$
\begin{align*}
& Q_{n}^{\delta}[\varphi, \psi]-Q_{1}^{\delta}[\varphi, \psi]  \tag{16}\\
& \begin{aligned}
(n-1) \int_{0}^{\infty}\left[\frac{n+3}{r^{2}}|\varphi|^{2}+\frac{n-1}{r^{2}}|\psi|^{2}\right. & -2 \delta \frac{n+1}{r^{2}} \operatorname{Re}\{\bar{\varphi} \bar{\psi}\} \\
& \left.+2 \frac{\delta}{r} \operatorname{Re}\left\{\bar{\varphi} \bar{\psi}^{\prime}-\bar{\varphi}^{\prime} \bar{\psi}\right\}\right] r d r .
\end{aligned}
\end{align*}
$$

Unlike what happens in the isotropic case, this does not obviously have a sign (because of the last term which contains derivatives).

It seems reasonable to use a decomposition for $\varphi, \psi$ adapted to $Q_{1}^{\delta}$, as in Appendix A. Accordingly, we define for any real-valued test functions $\zeta, \eta$, the adapted quadratic form

$$
B_{n}^{\delta}[\zeta, \eta]=\frac{1}{2} Q_{n}^{\delta}\left[f_{0}^{\prime} \zeta-r^{-1} f_{0} \eta, f_{0}^{\prime} \zeta+r^{-1} f_{0} \eta\right]
$$

Decomposing

$$
Q_{n}^{\delta}=(1+\delta) A_{1}+Q_{1}^{\delta}-(1+\delta) A_{1}+Q_{n}^{\delta}-Q_{1}^{\delta}
$$

and using the above expressions of $Q_{n}^{\delta}-Q_{1}^{\delta}(16)$ and $Q_{1}^{\delta}-(1+\delta) A_{1}$ (15), we have, for real-valued $(\varphi, \psi) \in \mathcal{H}_{1}:$

$$
\begin{aligned}
Q_{n}^{\delta}[\varphi, \psi]= & (1+\delta) A_{1}[\varphi, \psi] \\
& -\delta \int_{0}^{\infty}\left(\varphi^{\prime}+\frac{2}{r} \varphi-\psi^{\prime}\right)^{2} r d r \\
& +(n-1) \int_{0}^{\infty}\left[\frac{n+3}{r^{2}} \varphi^{2}+\frac{n-1}{r^{2}} \psi^{2}-2 \delta \frac{n+1}{r^{2}} \varphi \psi\right] \\
& +2 \delta(n-1) \int_{0}^{\infty} \frac{1}{r}\left(\varphi \psi^{\prime}-\varphi^{\prime} \psi\right) r d r .
\end{aligned}
$$

When plugging in $\varphi=f_{0}^{\prime} \zeta-r^{-1} f_{0} \eta, \psi=f_{0}^{\prime} \zeta+r^{-1} f_{0} \eta$, the first term significantly simplifies thanks to the formula (19) for $A_{1}$ in Appendix A. For the other terms we directly expand

$$
\begin{aligned}
& \varphi^{\prime}+\frac{2}{r} \varphi-\psi^{\prime}=2 f_{0}^{\prime} \frac{\zeta-\eta}{r}-2 \frac{f_{0}}{r} \eta^{\prime} \\
& \frac{n+3}{r^{2}} \varphi^{2}+\frac{n-1}{r^{2}} \psi^{2}-2 \delta \frac{n+1}{r^{2}} \varphi \psi \\
& =2(1-\delta) \frac{n+1}{r^{2}}\left(f_{0}^{\prime} \zeta\right)^{2}+2(1+\delta) \frac{n+1}{r^{2}}\left(\frac{f_{0}}{r} \eta\right)^{2}-\frac{8}{r^{2}} f_{0}^{\prime} \zeta \frac{f_{0}}{r} \eta \\
& \varphi \psi^{\prime}-\varphi^{\prime} \psi=2\left(\frac{f_{0}}{r} \eta\right)^{\prime} f_{0}^{\prime} \zeta-2\left(f_{0}^{\prime} \zeta\right)^{\prime} \frac{f_{0}}{r} \eta
\end{aligned}
$$

from which it follows that $B_{n}^{\delta}[\zeta, \eta]=(1 / 2) Q_{n}^{\delta}\left[f_{0}^{\prime} \zeta-r^{-1} f_{0} \eta, f_{0}^{\prime} \zeta+r^{-1} f_{0} \eta\right]$ can be rewritten as

$$
\begin{align*}
B_{n}^{\delta}[\zeta, \eta]= & (1+\delta) \int_{0}^{\infty}\left[\frac{f_{0}^{2}}{r^{2}}\left(\eta^{\prime}\right)^{2}+\left(f_{0}^{\prime}\right)^{2}\left(\zeta^{\prime}\right)^{2}+\frac{2}{r^{3}} f_{0} f_{0}^{\prime}(\eta-\zeta)^{2}\right] r d r  \tag{17}\\
& -2 \delta \int_{0}^{\infty}\left[\frac{f_{0}^{\prime}}{r}(\eta-\zeta)+\frac{f_{0}}{r} \eta^{\prime}\right]^{2} r d r \\
& +(n-1) \int_{0}^{\infty}\left[(1-\delta) \frac{n+1}{r^{2}}\left(f_{0}^{\prime} \zeta\right)^{2}+(1+\delta) \frac{n+1}{r^{2}}\left(\frac{f_{0}}{r} \eta\right)^{2}-\frac{4}{r^{2}}\left(f_{0}^{\prime} \zeta\right)\left(\frac{f_{0}}{r} \eta\right)\right] r d r \\
& +2 \delta(n-1) \int_{0}^{\infty} \frac{1}{r}\left[\left(\frac{f_{0}}{r} \eta\right)^{\prime} f_{0}^{\prime} \zeta-\left(f_{0}^{\prime} \zeta\right)^{\prime} \frac{f_{0}}{r} \eta\right] r d r
\end{align*}
$$

Integrating by parts, the last integral becomes

$$
\begin{aligned}
\int_{0}^{\infty} \frac{1}{r}\left[\left(\frac{f_{0}}{r} \eta\right)^{\prime} f_{0}^{\prime} \zeta-\left(f_{0}^{\prime} \zeta\right)^{\prime} \frac{f_{0}}{r} \eta\right] r d r & =2 \int_{0}^{\infty}\left(\frac{f_{0}}{r} \eta\right)^{\prime} f_{0}^{\prime} \frac{\zeta}{r} r d r \\
& =2 \int_{0}^{\infty}\left[\left(f_{0}^{\prime}-\frac{f_{0}}{r}\right) f_{0}^{\prime} \frac{\eta}{r} \frac{\zeta}{r}+\frac{f_{0}}{r} \eta^{\prime} f_{0}^{\prime} \frac{\zeta}{r}\right] r d r
\end{aligned}
$$

We use the first positive term in 17 in order to absorb this latter term: thanks to the identity

$$
\begin{aligned}
(1+\delta) \frac{f_{0}^{2}}{r^{2}}\left(\eta^{\prime}\right)^{2}+4 \delta(n-1) \frac{f_{0}}{r} \eta^{\prime} f_{0}^{\prime} \frac{\zeta}{r}= & (1+\delta)\left(\frac{f_{0}}{r} \eta^{\prime}+\frac{2 \delta}{1+\delta}(n-1) f_{0}^{\prime} \frac{\zeta}{r}\right)^{2} \\
& -4 \frac{\delta^{2}}{1+\delta}(n-1)^{2}\left(f_{0}^{\prime}\right)^{2}\left(\frac{\zeta}{r}\right)^{2}
\end{aligned}
$$

we rewrite (17) as

$$
\begin{aligned}
B_{n}^{\delta}[\zeta, \eta]= & B_{n}^{\delta, 1}[\zeta, \eta]+(n-1) B_{n}^{\delta, 2}[\zeta, \eta] \\
B_{n}^{\delta, 1}[\zeta, \eta]= & (1+\delta) \int_{0}^{\infty}\left[\left(\frac{f_{0}}{r} \eta^{\prime}+\frac{2 \delta}{1+\delta}(n-1) f_{0}^{\prime} \frac{\zeta}{r}\right)^{2}+\left(f_{0}^{\prime}\right)^{2}\left(\zeta^{\prime}\right)^{2}\right] r d r \\
& +2 \int_{0}^{\infty}\left\{(1+\delta) f_{0}^{\prime} \frac{f_{0}}{r} \frac{(\eta-\zeta)^{2}}{r^{2}}-\delta\left[\frac{f_{0}^{\prime}}{r}(\eta-\zeta)+\frac{f_{0}}{r} \eta^{\prime}\right]^{2}\right\} r d r \\
B_{n}^{\delta, 2}[\zeta, \eta]= & \int_{0}^{\infty} q_{n}^{\delta}(r)\left[f_{0}^{\prime} \frac{\zeta}{r}, \frac{f_{0}}{r} \frac{\eta}{r}\right] r d r
\end{aligned}
$$

and $q_{n}^{\delta}(r)$ is the quadratic form on $\mathbb{R}^{2}$ given by

$$
\begin{aligned}
q_{n}^{\delta}(r)[X, Y] & =a_{n} X^{2}+b_{n} Y^{2}+2 c(r) X Y \\
a_{n} & =(1-\delta)(n+1)-4 \frac{\delta^{2}}{1+\delta}(n-1) \\
b_{n} & =(1+\delta)(n+1) \\
c(r) & =-2-2 \delta\left(1-r \frac{f_{0}^{\prime}}{f_{0}}\right)
\end{aligned}
$$

We readily see that $B_{n}^{\delta, 1}$ is nonnegative for $\delta \leq 0$. Moreover, since $1>r f_{0}^{\prime} / f_{0}>07$, Proposition 2.2], for $\delta \leq 0$, it follows that

$$
|c(r)| \leq 2
$$

As $b_{n}>0$, a sufficient condition for $q_{n}^{\delta}(r)$ to be positive definite for all $r>0$ is

$$
4<a_{n} b_{n}=\left(1-\delta^{2}\right)(n+1)^{2}-4 \delta^{2}\left(n^{2}-1\right)
$$

This amounts to the condition

$$
0<\alpha(\delta) n^{2}+\beta(\delta) n+\gamma(\delta)
$$

where

$$
\begin{aligned}
& \alpha(\delta)=1-5 \delta^{2} \\
& \beta(\delta)=2\left(1-\delta^{2}\right) \\
& \gamma(\delta)=-3\left(1-\delta^{2}\right) .
\end{aligned}
$$

For $\delta \in[-1 / \sqrt{5}, 0]$ we have $\alpha(\delta), \beta(\delta) \geq 0$ so that the above polynomial in $n$ is nondecreasing on $[0,+\infty)$. Hence, it is positive for all values of $n \geq 2$ if and only if it is positive for $n=2$. That is,

$$
0<4 \alpha(\delta)+2 \beta(\delta)+\gamma(\delta)=5-21 \delta^{2}
$$

We deduce that $q_{n}^{\delta}$ is a positive definite quadratic form for all $n \geq 2$ whenever $\delta \in[-1 / \sqrt{5}, 0]$. In particular, $B_{n}^{\delta, 2} \geq 0$ and therefore $Q_{n}^{\delta} \geq 0$ for $\delta \in[-1 / \sqrt{5}, 0]$, with equality only at $(0,0)$.
4.2. Instability for $\delta \approx-1$. In this section we show that $Q_{n}^{\delta}$ can take negative values for $\delta \approx-1$ and $n \geq 1$ large enough. To this end, we choose $\eta=\zeta$ in (17), to obtain

$$
\begin{aligned}
\hat{B}_{n}^{\delta}[\zeta] & =B_{n}^{\delta}[\zeta, \zeta] \\
& =(1-\delta) \int_{0}^{\infty} \frac{f_{0}^{2}}{r^{2}}\left(\zeta^{\prime}\right)^{2} r d r+(1+\delta) \int_{0}^{\infty}\left(f_{0}^{\prime}\right)^{2}\left(\zeta^{\prime}\right)^{2} r d r+(n-1) \int_{0}^{\infty} \frac{\zeta^{2}}{r^{2}} \alpha_{n}^{\delta}(r) r d r \\
\alpha_{n}^{\delta}(r) & =(1-\delta)(n+1)\left(f_{0}^{\prime}\right)^{2}+(1+\delta)(n+1)\left(\frac{f_{0}}{r}\right)^{2}-2(2+\delta) f_{0}^{\prime} \frac{f_{0}}{r}+2 \delta\left(f_{0}^{\prime}\right)^{2}-2 \delta f_{0} f_{0}^{\prime \prime} .
\end{aligned}
$$

Using the asymptotics of $f_{0}([1,6])$

$$
f_{0}(r)=1-\frac{1}{2} r^{-2}+O\left(r^{-4}\right), \quad f_{0}^{\prime}(r)=r^{-3}+O\left(r^{-5}\right), \quad f_{0}^{\prime \prime}(r)=-3 r^{-4}+O\left(r^{-6}\right)
$$

we find, for $r \rightarrow+\infty$,

$$
\alpha_{n}^{\delta}(r)=\frac{(1+\delta)(n+1)}{r^{2}}\left(1-\frac{1}{r^{2}}\right)-4 \frac{1-\delta}{r^{4}}+O\left(r^{-6}\right) .
$$

For $\delta=-1$ the leading order is negative. Hence, there exists $\varepsilon>0$ and a compact interval $\left[r_{0}, r_{0}+1\right]$ on which $\alpha_{n}^{-1} \leq-2 \varepsilon$. Thus, we deduce that for all $n \geq 2$ there exists $\delta_{n}>-1$ such that for all $\delta \in\left(-1, \delta_{n}\right]$,

$$
-\varepsilon \geq \alpha_{n}^{\delta}(r), \quad \forall r \in\left[r_{0}, r_{0}+1\right] .
$$

Choosing a nonzero test function $\zeta_{0}$ with support in $\left[r_{0}, r_{0}+1\right]$, we obtain

$$
\hat{B}_{n}^{\delta}\left[\zeta_{0}\right] \leq C_{1}\left(\zeta_{0}\right)-(n-1) \varepsilon C_{2}\left(\zeta_{0}\right) \quad \forall \delta \in\left(-1, \delta_{n}\right]
$$

for some $C_{1}\left(\zeta_{0}\right), C_{2}\left(\zeta_{0}\right)>0$. If $n$ is large enough this becomes negative. Compared to the isotropic case this is a really new situation: lower modes are positive but higher modes can bring instability.

## 5. Proof of Theorem 1.3

In what precedes we have shown that $u_{\text {rad }}^{\delta}$ is nondegenerately stable for small $\delta \leq 0$, and unstable for $\delta>0$ and $\delta$ close to -1 . In particular, setting

$$
\delta_{1}=\sup \left\{\delta \in(-1,0): u_{\text {rad }}^{\delta} \text { is unstable }\right\},
$$

we know that $-1<\delta_{1}<0$. It remains to show that $u_{\text {rad }}^{\delta}$ is unstable for all $\delta \in\left(-1, \delta_{1}\right)$, and nondegenerately stable for $\delta \in\left(\delta_{1}, 0\right]$.

Let $\delta^{\prime} \in\left(-1, \delta_{1}\right)$ be such that $u_{\mathrm{rad}}^{\delta}$ is unstable, that is, $\mathcal{Q}^{\delta^{\prime}}[v]<0$ for some choice of $v \in H$. Given that $\delta \mapsto \mathcal{Q}^{\delta}[v]$ is an affine function which is nonnegative for $\delta=0$ and negative for $\delta=\delta^{\prime}$, we deduce that $\mathcal{Q}^{\delta}[v]<0$ for all $\delta \leq \delta^{\prime}$. Therefore, $u_{\text {rad }}^{\delta}$ is unstable for all $\delta \in\left(-1, \delta^{\prime}\right)$. By arbitrariness of $\delta^{\prime}$ we deduce that $u_{\text {rad }}^{\delta}$ is unstable for all $\delta \in\left(-1, \delta_{1}\right)$.

Let us now fix $\delta \in\left(\delta_{1}, 0\right]$. By definition of $\delta_{1}, u_{\text {rad }}^{\delta}$ is not unstable for all $\delta \in\left(\delta_{1}, 0\right]$. In other words, $\mathcal{Q}^{\delta}[v]$ is nonnegative for all $v \in \mathcal{H}$. It remains to show that, in fact, $\mathcal{Q}^{\delta}[v]>0$ for all $v \in \mathcal{H} \backslash \operatorname{span}\left(\partial_{x} u_{\mathrm{rad}}^{0}, \partial_{y} u_{\mathrm{rad}}^{0}\right)$. We observe that the function $\delta \mapsto \mathcal{Q}^{\delta}[v]$ is affine for any given $v \in \mathcal{H} \backslash \operatorname{span}\left(\partial_{x} u_{\mathrm{rad}}^{0}, \partial_{y} u_{\mathrm{rad}}^{0}\right)$; it is positive for $\delta=0$ because $u_{\mathrm{rad}}^{0}$ is nondegenerately stable, and it is nonnegative for $\delta \in\left(\delta_{1}, 0\right)$. Thus, it must be strictly positive for $\delta \in\left(\delta_{1}, 0\right)$. This proves the desired nondegenerate stability in the announced range.

## Appendix A. Positivity of $A_{0}, A_{1}$

We sketch here the approach in [12], adapted to our notation (see also [5]), based on Hardytype decompositions to show positivity of the two following quadratic forms

$$
\begin{aligned}
& A_{0}[\varphi]=\int_{0}^{\infty}\left[\left|\varphi^{\prime}\right|^{2}+\frac{1}{r^{2}}|\varphi|^{2}\right. \\
& \left.+2 f_{0}^{2}(\operatorname{Re}\{\varphi\})^{2}-\left(1-f_{0}^{2}\right)|\varphi|^{2}\right] r d r, \\
& A_{1}[\varphi, \psi]=\int_{0}^{\infty}\left[\left|\varphi^{\prime}\right|^{2}+\left|\psi^{\prime}\right|^{2}+\frac{4}{r^{2}}|\varphi|^{2}\right. \\
& \left.+f_{0}^{2}|\varphi+\bar{\psi}|^{2}-\left(1-f_{0}^{2}\right)\left(|\varphi|^{2}+|\psi|^{2}\right)\right] r d r .
\end{aligned}
$$

Testing equation (9), solved by $f_{0}$, against $f_{0}|\tilde{\varphi}|^{2}$ for any smooth compactly supported $\tilde{\varphi} \in C_{c}^{\infty}(\mathbb{R} ; \mathbb{C})$, one obtains

$$
\int_{0}^{\infty}\left[\left(f_{0}^{\prime}\right)^{2}|\tilde{\varphi}|^{2}+2 f_{0} f_{0}^{\prime} \tilde{\varphi} \cdot \tilde{\varphi}^{\prime}+\frac{f_{0}^{2}}{r^{2}}|\tilde{\varphi}|^{2}-\left(1-f_{0}^{2}\right) f_{0}^{2}|\tilde{\varphi}|^{2}\right] r d r=0
$$

so that

$$
\begin{equation*}
A_{0}\left[f_{0} \tilde{\varphi}\right]=\int_{0}^{\infty}\left[f_{0}^{2}\left|\tilde{\varphi}^{\prime}\right|^{2}+2 f_{0}^{4}(\operatorname{Re}\{\tilde{\varphi}\})^{2}\right] r d r \tag{18}
\end{equation*}
$$

By density of test functions, and since $f_{0}>0$, we deduce that $A_{0}[\varphi]>0$ for any non-zero $\varphi \in \mathcal{H}_{0}$. Moreover $A_{0}[\varphi] \approx 0$ exactly when $\varphi \approx i f_{0}$. This corresponds to the fact that in the isotropic case $\delta=0$,

$$
\partial_{\alpha}\left[e^{i \alpha} u_{\mathrm{rad}}^{\delta}\right]_{\lfloor\alpha=0}=i f_{0} e^{i \theta}
$$

solves the linearized equation due to rotational invariance.
For $A_{1}$, it is convenient to start by splitting it as

$$
A_{1}[\varphi, \psi]=A_{1}[\operatorname{Re}\{\varphi\}, \operatorname{Re}\{\psi\}]+A_{1}[\operatorname{Im}\{\varphi\},-\operatorname{Im}\{\psi\}],
$$

so we may just treat the case of real-valued test functions $\varphi, \psi$. Guided by the fact that

$$
\partial_{x} u_{\mathrm{rad}}^{0}=e^{i \theta}\left(f_{0}^{\prime} \cos \theta-i \frac{f_{0}}{r} \sin \theta\right), \quad \partial_{y} u_{\mathrm{rad}}^{0}=e^{i \theta}\left(f_{0}^{\prime} \sin \theta+i \frac{f_{0}}{r} \cos \theta\right),
$$

solve the linearized equation around $u_{\mathrm{rad}}^{0}$, one uses the ansatz

$$
\varphi=f_{0}^{\prime} \zeta-\frac{f_{0}}{r} \eta, \quad \psi=f_{0}^{\prime} \zeta+\frac{f_{0}}{r} \eta
$$

for some real-valued $\eta, \zeta \in C_{c}^{\infty}(0, \infty)$. Testing equation (9), solved by $f_{0}$, against $f_{0} r^{-2} \eta^{2}$ we obtain

$$
\int_{0}^{\infty}\left[\left(\left(\frac{f_{0}}{r}\right)^{\prime}\right)^{2} \eta^{2}+2\left(\frac{f_{0}}{r}\right)^{\prime} \frac{f_{0}}{r} \eta \eta^{\prime}+\frac{2}{r^{4}} f_{0}^{2} \eta^{2}-\frac{2}{r^{3}} f_{0} f_{0}^{\prime} \eta^{2}-\left(1-f_{0}^{2}\right) \frac{f_{0}^{2}}{r^{2}} \eta^{2}\right] r d r=0
$$

and similarly testing (9) against $\left(f_{0}^{\prime} \zeta^{2}\right)^{\prime}$ we find

$$
\int_{0}^{\infty}\left[\left(f_{0}^{\prime \prime}\right)^{2} \zeta^{2}+2 f_{0}^{\prime} f_{0}^{\prime \prime} \zeta \zeta^{\prime}+\frac{2}{r^{2}}\left(f_{0}^{\prime}\right)^{2} \zeta^{2}-\frac{2}{r^{3}} f_{0} f_{0}^{\prime} \zeta^{2}+\left(3 f_{0}^{2}-1\right)\left(f_{0}^{\prime}\right)^{2} \zeta^{2}\right] r d r=0
$$

As a consequence of these two identities, we learn

$$
\begin{align*}
& A_{1}\left[f_{0}^{\prime} \zeta-r^{-1} f_{0} \eta, f_{0}^{\prime} \zeta+r^{-1} f_{0} \eta\right]  \tag{19}\\
& =2 \int_{0}^{\infty}\left[\frac{f_{0}^{2}}{r^{2}}\left(\eta^{\prime}\right)^{2}+\left(f_{0}^{\prime}\right)^{2}\left(\zeta^{\prime}\right)^{2}+\frac{2}{r^{3}} f_{0} f_{0}^{\prime}(\eta-\zeta)^{2}\right] r d r
\end{align*}
$$

Since $f_{0}, f_{0}^{\prime}>0$ one may consider the choice

$$
\zeta=\frac{1}{2 f_{0}^{\prime}}(\varphi+\psi), \quad \eta=\frac{r}{2 f_{0}}(\psi-\varphi)
$$

and deduce from the above that $A_{1}[\varphi, \psi]>0$ for all non-zero $(\varphi, \psi) \in \mathcal{H}_{1}$. Moreover $A_{1}[\varphi, \psi]=$ 0 exactly when $(\varphi, \psi)$ is in the real linear span of

$$
\left(f_{0}^{\prime}-\frac{f_{0}}{r}, f_{0}^{\prime}+\frac{f_{0}}{r}\right), \quad\left(i\left(f_{0}^{\prime}-\frac{f_{0}}{r}\right),-i\left(f_{0}^{\prime}+\frac{f_{0}}{r}\right)\right)
$$

which corresponds to the fact that $\partial_{x} u_{\mathrm{rad}}^{0}$ and $\partial_{y} u_{\mathrm{rad}}^{0}$ solve the linearized equation.

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