# Localized radial bumps of a neural field equation on the Euclidean plane and the Poincaré disk

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June 28, 2012

#### Abstract

We analyze radially symmetric localized bump solutions of an integro-differential neural field equation posed in Euclidean and hyperbolic geometry. The connectivity function and the nonlinear firing rate function are chosen such that radial spatial dynamics can be applied. Using integral transforms, we derive a PDE of the neural field equation in both geometries and then prove the existence of small amplitude radially symmetric spots bifurcating from the trivial state. Numerical continuation is then used to path-follow the spots and their bifurcations away from onset in parameter space. It is found that the radial bumps in Euclidean geometry are linearly stable in a larger parameter region than bumps in the hyperbolic geometry. We also find and path follow localized structures that bifurcate from branches of radially symmetric solutions with  $\mathbf{D}_6$ -symmetry and  $\mathbf{D}_8$ -symmetry in the Euclidean and hyperbolic cases, respectively. Finally, we discuss the applications of our results in the context of neural field models of short term memory and edges and textures selectivity in a hypercolumn of the visual cortex.

Mathematics Subject Classification : 35B22; 37G05; 92B20.

Keywords: Neural field equation; Localized patterns; Poincaré disk; Numerical continuation.

# 1 Introduction

A popular approach for understanding coarse-grained activity of large ensembles of neurons in cortex is to model cortical space as continuous. This natural approximation, based on the pioneer work of Wilson and Cowan [97, 98] and Amari [1, 2], gives rise to the notion of neural field models. These models typically take the form of integro-differential equations that describe the evolution of macroscopic variables such as the firing rate activity or average membrane voltage of populations of neurons. Unlike spiking neural network models, neural field models have the advantage that analysis techniques for partial differential equations (PDEs) can be adapted to study their nonlinear dynamics with a small set of parameters. Various types of dynamical behaviour have been observed in neural field models ranging from spatially and temporally periodic patterns [17, 35, 39, 61, 90], localized regions of activity [64, 70] to traveling waves [47, 48, 63, 71, 83]. Neural field equations have also successfully been used to model a wide range of neurobiological phenomena such as visual hallucinations [20, 21, 39, 43, 50], mechanisms for short term memory [37, 44, 53, 54, 69, 70] and feature selectivity in the visual cortex [14, 16, 19, 46, 55].

In this paper, we consider the canonical Wilson-Cowan-Amari neural field equation [1, 98] that describes the evolution of the average membrane voltage of a neuronal population,  $u(\mathbf{r}, t)$ , given by

$$\tau \partial_t u(\mathbf{r}, t) = -u(\mathbf{r}, t) + \int_{\Omega} W(\mathbf{r}, \mathbf{r}') S(u(\mathbf{r}', t)) \mathrm{dm}(\mathbf{r}')$$
(1.1)

where  $\tau$  is a temporal constant, that we set equal to one for mathematical simplicity, and  $\Omega$  can either model a piece of cortex or a feature space with the corresponding measure dm(·). The firing rate function S is generally chosen to be either a Heaviside step function [1, 32, 84], a piecewise linear function [53, 54] or a smooth function of sigmoidal type [31, 37, 40, 70]. The coupling of the neurons, described by the connectivity function W, is typically assumed to have a so-called "Mexican hat" or so-called "wizard hat" shape, although strictly positive connectivity functions have also been considered in, for example, [85]. The precise form of the connectivity function is usually chosen such that a Fourier transform can be applied to (1.1) to yield a PDE; such an approach goes back to the work of Jirsa and Haken in [60] and has been used by a number of other authors [31, 69, 70].

When modelling a piece of cortex,  $\Omega$  is taken to be a subset of  $\mathbb{R}^d$  with d = 1, 2 or 3 and naturally inherits a Euclidean geometry. For  $\Omega = \mathbb{R}$ , the most widely studied case, various analyses have found a range of solutions of (1.1) from localized states [32, 44, 53, 54, 70] to spatially periodic [6, 7, 30, 58]. However, the case  $\Omega = \mathbb{R}$  is of limited biological interest since the visual cortex is inherently a three dimensional object. Fortunately, a piece of cortex can be well approximated by a sheet as its thickness is much smaller than the other directions and this has lead various studies to look at the case where  $\Omega = \mathbb{R}^2$ ; see for instance [32, 47, 48, 61, 69, 82, 91, 96]. Recently, there has also been much interest in studying equation (1.1) in other Riemannian geometries such as hyperbolic geometry where  $\Omega$  can be taken to be the Poincaré disk  $\mathbb{D}^1$  and the neural field equation now models edge and texture selectivity in the visual cortex [26, 46]. Only periodic solutions in the Poincaré disk so far been investigated by Chossat and Faye [27, 45].

In the case of Euclidean geometry with a Heaviside firing function S, it is possible to explicitly construct general radially symmetric bump solutions to equation (1.1); see [48, 91, 96]. While this approach allows one to construct closed form formulae for the solutions, it has the drawback that it does not explain the nucleation, from a dynamical point of view, of radially symmetric bumps with a more realistic firing rate such as a smooth sigmoidal function.

The first study with more realistic firing function was carried out by Laing *et al.* [37, 69, 70] on both the real line and the Euclidean plane where numerical continuation methods have been used to path follow localized solutions. In the previous studies of Laing *et al.*, spatial dynamics methods can not be used to prove the existence of localized states due to the non-analyticity of the chosen firing rate function. A bifurcation analysis for more general analytic firing rate function was recently carried out by Faye *et al.* [44] where they were able to perform normal form analysis in one dimension near a Turing instability and describe the nucleation process of localized patterns. Furthermore, Faye *et al.* [44] found that the bifurcating pulses underwent homoclinic snaking away from onset similar to that seen in the Swift-Hohenberg equation; see [23, 34, 99] for a discussion of homoclinic snaking.

Most choices of connectivity functions lead to a PDE (via a Fourier transform) whose stationary states satisfy a "Swift-Hohenberg"-type PDE [44, 60, 69, 70]. Recently, there has been much progress made in understanding localized patterns in the Swift-Hohenberg equation on the plane where various planar (cellular hexagon, square and roll) fronts and fully localized two-dimensional patches have been investigated; we note that similar hexagonal patches have been computed in (1.1) by Laing & Troy [69]. For one-dimensional and planar localized patterns, a detailed geometric dynamical systems theory has been developed to explain the intricate bifurcation structure of such patterns [13, 24, 65, 66]. Unfortunately, the geometric theory breaks down for fully localized two-dimensional patches and only numerical studies exist. However, for the Swift-Hohenberg equation near the Turing instability, three types of small amplitude radially symmetric localized solutions have been proven to exist: a localized ring decaying to almost zero at the core, a spot with a maximum at the origin (called spot A; see Figure 1(a)) and a spot with minimum at the origin (called spot B; see Figure 1(b)); see [74, 77–79]. The proofs rely on matching, at  $O(1/r^2)$ , the "Core" manifold that describes solutions that remain bounded near r = 0 with the "Far-Field" manifold that describes how solutions decay to the trivial state for large r. The Core manifold is found by carrying out an asymptotic expansion involving Bessel functions while the Far-Field manifold is calculated by carrying out a radial

 $<sup>^{1}\</sup>mathbb{D}=\{z\in\mathbb{C}\ |\ 0\leq|z|<1\}.$ 

normal form expansion near  $r = \infty$ . The main technical difficulty is that the Far-Field normal form is only valid down to O(1/r) and so the manifold has to be carefully followed until  $O(1/r^2)$ . Localized rings occur due to a localized pulse in the Far-Field normal form equations and require that the bifurcation of rolls at  $r = \infty$  is subcritical. Spot A solutions occur due to the unfolding of a quadratic tangency of the core manifold and the cubic tangency of the far-field manifold with the trivial state at onset; see [74, Figure 4]. The spot B state is formed by 'gluing' the spot A and localized ring solution. Crucially, all these localized radial states are L<sup>2</sup>-functions that can not be found via a straight forward Lyapunov-Schmidt reduction. In this article, we will only be interested in the existence of spot A type of solutions for the neural field equation (1.1) as these localized solutions are the most relevant from a neuroscience point of view since numerics show that stable spots closely resemble a  $J_0(r)$ -Bessel function at the core with no ring structure [29, 49, 80]. For the localized ring and spot B proofs one requires an additional hypothesis that the stable and unstable manifolds of ring solution in the far-field normal form transversely intersect; see [79]. Spot B and localized rings have not been observed experimentally and we shall not discuss these further. From now on we shall use the terms, spot and bump interchangeably to refer to the spot A states.



Figure 1: Examples of spot solutions as computed for the neural field equations, which are solutions to (3.7); see §3.1.3 and §4 for further details. (a) Spot A solution that exists for  $\widehat{W}_c = 5$ ,  $\mu = 6$ . (b) Coexisting spot B solution.

The main aim of this paper is to prove the existence of radially symmetric localized solutions for the neural field equation (1.1) in the case where  $\Omega$  is equal to either  $\mathbb{R}^2$  or  $\mathbb{D}$  with a general analytic firing rate function. We will show that the existence proof of spot (spot A) solutions in the Swift-Hohenberg equation (in Euclidean geometry) can be adapted to finding radial spots in (1.1) in both Euclidean and hyperbolic geometry and hence we present both proofs together. These existence results provide a starting point for a numerical investigation of other fully localized patterns that bifurcate off the radial spots and are of interest from a biological point of view.

The spot existence proofs in the Euclidean case are a straight-forward adaption of those for the Swift-Hohenberg equation and provide the basic ideas for the existence proof in the hyperbolic case. Two major problems need to be overcome in order to prove the existence of spot A solutions in the hyperbolic case. Firstly, one needs to define a suitable connectivity function and integral transform such that a PDE can be derived where radial spatial dynamics can be applied. Recently, Faye [41] showed how to define a suitable connectivity function and integral transform for (1.1) posed on the Poincaré disk. Secondly, it is not clear how to define the Core and Far-Field manifolds in order to carry out the matching. In this paper, we show how this problem can be overcome. It turns out that the Far-Field manifold is easier to define than in the Euclidean case since there is no bifurcation in the Far-Field. However, calculating the Core manifold is significantly more involved than in the Euclidean case and constitutes the main challenge in the existence proof of spots in hyperbolic geometry.

Along with the analytical study, we present a numerical investigation of different types of 2D fully localized solutions found on connected solution branches related to the spot states using the numerical continuation

package AUTO07P [36]. Extensive novel numerical methods have been developed for the Swift-Hohenberg equation to carry out numerical continuation for fully localized 2D patterns in order to gain insight into the bifurcation structure away from onset; see [75]. We show how these methods can be applied to (1.1) and how stability can also be computed. We identify the exact regions of parameter space for which linearly stable localized solutions persist in terms of one parameter governing the shape of the nonlinearity and a second parameter governing the shape of the connectivity function. Furthermore, we investigate non-radially symmetric localized solutions that have symmetry properties particular to the specific geometry under study; in the Euclidean case  $\mathbf{D}_6$ -symmetric solutions and in the hyperbolic case  $\mathbf{D}_8$ -symmetric solutions.

This paper is divided into three parts. The first part is dedicated to the presentation of the models and the statement of the main theorems in §2.1 and §2.2. The second part, §3, is devoted to the theoretical analysis of our models and the proofs of our theorems. We apply techniques from Scheel [88] and Lloyd & Sandstede [74] to prove the existence of a bifurcated branch of spot (Spot A) solutions near a Turing instability for the Euclidean case and extend our results to the hyperbolic case. In the last part, §4, we use numerical continuation in order to extend the study by path following the solutions found in the analytical studies; we identify in terms of two parameters the regions of parameter space for which stable, radially symmetric, localized solutions exist. Furthermore, we find and path follow connected branches of non-radially symmetric solutions. Finally, in §5 we conclude with a discussion of the relevance of the localized states found to the study of neural field models and associated applications.

# 2 The neural field models and main results

### 2.1 Neural field model of short term memory on the Euclidean plane

The time evolution of the average membrane potential  $v(\mathbf{r}, t)$  at time t and position  $\mathbf{r}$  in the visual cortex, abstracted by  $\mathbb{R}^2$ , is governed by the following Wilson-Cowan neural field equation

$$\frac{d}{dt}v(\mathbf{r},t) = -v(\mathbf{r},t) + \int_{\mathbb{R}^2} W(\|\mathbf{r} - \mathbf{r}'\|) S(\mu v(\mathbf{r}',t)) d\mathbf{r}',$$
(2.1)

where  $\|\cdot\|$  is the standard Euclidean norm. The nonlinear firing rate function S is defined by

$$S(x) = \frac{1}{1 + e^{-x + \theta}} - \frac{1}{1 + e^{\theta}}.$$
(2.2)

 $\theta$  is a positive threshold and the parameters  $\mu$  in (2.1) describes the stiffness of the sigmoid S. The connectivity function W expresses interactions between populations of neurons located at position **r** and **r'** in the visual cortex and is defined through its *Hankel transform* in the following way

$$W(x) = \int_0^{+\infty} s\widehat{W}(s)J_0(xs)ds, \quad \widehat{W}(s) = 2\left(\frac{b_1\sigma_1}{\sigma_1^2 + s^2} - \frac{b_2\sigma_2}{\sigma_2^2 + s^2}\right).$$
 (2.3)

 $J_0(s)$  is the Bessel function of first kind and  $(b_1, b_2, \sigma_1, \sigma_2)$  are real parameters. One of the major modelling issues is to determine how localized solutions of equation (2.1) depend on the connectivity function W. Here we assume the connectivity function W has a "Mexican hat" shape. This type of connectivity function can be regarded as a two-layer model in which the inhibition is linear and very fast; see [83, 84]. The Fourier transform of  $W(\|\mathbf{r}\|)$  is  $\widehat{W}(\|\mathbf{k}\|)$  for all  $(\mathbf{r}, \mathbf{k}) \in \mathbb{R}^2$ . The parameters  $(b_1, b_2, \sigma_1, \sigma_2)$  are chosen such that Hypothesis 2.1 is satisfied.

**Hypothesis 2.1.** We assume the connectivity function W(x) and its corresponding Fourier transform  $\widehat{W}(\|\mathbf{k}\|)$  satisfy the following conditions:

- $(i) \ \widehat{W}(0) = \widehat{W}_0 < 0,$
- (*ii*) W(0) > 0,

(iii) there exists  $k_c > 0$  such that  $\widehat{W}_c \stackrel{def}{=} \widehat{W}(k_c) = \max_{k = \|\mathbf{k}\| \in \mathbb{R}^+} \widehat{W}(k) > 0$ ,

(*iv*)  $k_c = 1$ .

We define

$$\Gamma_1 = 2\sigma_1 \sigma_2 (b_1 \sigma_2 - b_2 \sigma_1),$$
  

$$\Gamma_2 = 2(b_1 \sigma_1 - b_2 \sigma_2),$$
(2.4)

such that  $\widehat{W}_c$  can be written

$$\widehat{W}_{c} = \frac{\Gamma_{1} + \Gamma_{2}}{1 + \sigma_{1}^{2} + \sigma_{2}^{2} + \sigma_{1}^{2}\sigma_{2}^{2}}.$$
(2.5)

The condition  $k_c = 1$ , which is equivalent to  $\frac{d}{ds}\widehat{W}(s)|_{s=1} = 0$ , reduces to

$$\Gamma_1(\sigma_1^2 + \sigma_2^2 + 2) + \Gamma_2(1 - \sigma_1^2 \sigma_2^2) = 0.$$
(2.6)

It is a straightforward computation to see that equations (2.5) and (2.6) imply that

$$\sigma_1^2 \sigma_2^2 - \Gamma_1 \widehat{W}_c^{-1} = 1,$$
  

$$\sigma_1^2 + \sigma_2^2 - \Gamma_2 \widehat{W}_c^{-1} = -2.$$
(2.7)

We define for all  $k \ge 1$ ,  $S^{(k)}(0) = s_k$ , the kth derivative of S evaluated at zero.

**Remark 2.1.** The dependence, with respect to the connectivity function, of the existence of stationary localized solution of equation (2.1) in the case where the visual cortex is further approximated by  $\mathbb{R}$  has been discussed in [44].

We can now state the first of our two theorems.

**Theorem 2.1** (Existence of spot solutions in Euclidean geometry). Fix  $\theta \geq 0$  for the threshold of the nonlinearity S and  $(b_1, b_2, \sigma_1, \sigma_2) \in \mathbb{R}^4$  such that the connectivity function W defined in equation (2.3) satisfies Hypotheses 2.1. Then there exists  $\mu_* < \mu_c$  such that the planar neural field equation (2.1) has a stationary localized radial solution  $v(r) \in L^2(\mathbb{R}^+, rdr)$  for each  $\mu \in ]\mu_*, \mu_c[$ : these solutions stay close to v = 0 and, for each fixed  $r_* > 0$ , we have the asymptotics

$$v(r) = \alpha \sqrt{\mu_c - \mu} J_0(r) + O(\mu - \mu_c) \text{ as } \mu \to \mu_c,$$
 (2.8)

uniformly in  $0 \leq r \leq r_*$  for an appropriate constant  $\alpha$  with  $sign(\alpha) = sign(s_2)$ .

# 2.2 Neural field model of edges and textures selectivity on the Poincaré disk

The selectivity of the responses of individual neurons to external features are often the basis of neuronal representations of the external world. For example, neurons in the primary visual cortex respond preferentially to visual stimuli that have a specific orientation [14, 18, 55], spatial frequency [16], velocity, direction of motion [81], and color [56]. A local network in the primary visual cortex, roughly 1 mm<sup>2</sup> of cortical surface, is assumed to consist of neurons coding for a given position in the retina for a full functional set of orientations and ocular dominance. These subgroups are the so-called Hubel and Wiesel hypercolumns of V1 [57]. A new approach to model the processing of image edges and textures in the hypercolumns of area V1 based on a nonlinear representation of the image first order derivatives called the structure tensor [15, 67] was introduced in [26, 46]. It was suggested that this structure tensor was represented by neuronal populations in the hypercolumns of V1 and that the time evolution of this representation was governed by equations similar to those proposed by Wilson and Cowan [97, 98]. Structure tensors are essentially  $2 \times 2$  symmetric, definite positive matrices. They therefore live in a solid open cone in  $\mathbb{R}^3$ , which is a Riemannian manifold foliated by hyperbolic planes. By a suitable change of coordinates, the hyperbolic plane can be further identified with the Poincaré disc  $\mathbb{D} = \{z \in \mathbb{C} \mid |z| < 1\}$ . Therefore, there is an isomorphism between the space of structure tensors and the product space  $\mathbb{R}^+_* \times \mathbb{D}^2$ , on which the distance is given by

$$d(\delta, z; \delta', z') = \sqrt{\log^2\left(\frac{\delta}{\delta'}\right) + d_{\mathbb{D}}(z, z')^2},$$
(2.9)

where the second term under the radical is the usual "hyperbolic" distance in  $\mathbb{D}$ :

$$d_{\mathbb{D}}(z,z') = 2 \operatorname{atanh}\left(\frac{|z-z'|}{|1-\bar{z}z'|}\right).$$
(2.10)

As shown in [27], we are able to neglect the dependence on  $\delta \in \mathbb{R}^+_*$  as it does not play a significant role in the analysis that follows. Hence, we can consider the following neural field equation set on the Poincaré disk

$$\frac{d}{dt}v(z,t) = -v(z,t) + \int_{\mathbb{D}} W(d_{\mathbb{D}}(z,z'))S(\mu v(z',t))\mathrm{dm}(z').$$
(2.11)

The measure element  $dm(\cdot)$  is given by the formula:

$$dm(z) = \frac{4dzd\bar{z}}{(1-|z|^2)^2}.$$
(2.12)

The nonlinearity S has already been defined in equation (2.2). We work in geodesic polar coordinates  $z = (\tau, \varphi) \in \mathbb{D}$ , with  $z = \tanh(\tau/2)e^{i\varphi}$ . In this coordinates system, the measure element defined in equation (2.12) is transformed into  $\operatorname{dm}(z) = \sinh(\tau)d\tau d\varphi$ . As we want to prove the existence of radial solutions ( $\tau$  being the radial coordinate), we define the equivalent of the Hankel transform for the hyperbolic disk, the *Mehler-Fock transform*, and we adopt the notations defined in [52].

**Definition 2.1.** We denote  $\mathcal{P}^{\mu}_{\nu}(z)$  and  $\mathcal{Q}^{\mu}_{\nu}(z)$  the two associated Legendre function solutions of the equation

$$(1-z^2)\partial_z^2 u(z) - 2z\partial_z u(z) + \left(\nu(\nu+1) - \frac{\mu^2}{1-z^2}\right)u(z) = 0.$$
(2.13)

Let  $f: [0, +\infty) = \mathbb{R}^+_* \to \mathbb{R}$  such that  $f(\tau)e^{\tau/2} \in L^1(\mathbb{R}^+_*)$ , then we define for all  $\rho > 0$  and all  $\tau > 0$  the Mehler-Fock transform of f as

$$\mathcal{M}f(\rho) = \int_0^{+\infty} f(\tau)\mathcal{P}_{-\frac{1}{2}+i\rho}(\cosh\tau)\sinh\tau d\tau.$$
(2.14)

The inversion formula states that

$$f(\tau) = \int_0^{+\infty} \mathcal{M}f(\rho)\mathcal{P}_{-\frac{1}{2}+i\rho}(\cosh\tau)\rho\tanh(\pi\rho)d\rho.$$
(2.15)

Moreover, we denote  $L^\tau_{\mathbb D}$  to be the radial part of the Laplace-Beltrami operator on  $\mathbb D$  defined by

$$L_{\mathbb{D}}^{\tau} = \frac{1}{\sinh \tau} \frac{d}{d\tau} \left( \sinh \tau \frac{d}{d\tau} \right).$$
(2.16)

It was shown in [52] that for all  $k \ge 1$ 

$$\begin{cases} (-1)^k \left(\frac{1}{4} + \rho^2\right)^k \mathcal{M}f(\rho) = \int_0^{+\infty} \left(L_{\mathbb{D}}^{\tau}\right)^k f(\tau) \mathcal{P}_{-\frac{1}{2} + i\rho}(\cosh \tau) \sinh \tau d\tau, \\ (-1)^k \left(L_{\mathbb{D}}^{\tau}\right)^k f(\tau) = \int_0^{+\infty} \left(\frac{1}{4} + \rho^2\right)^k \mathcal{M}f(\rho) \mathcal{P}_{-\frac{1}{2} + i\rho}(\cosh \tau)\rho \tanh(\pi\rho)d\rho. \end{cases}$$
(2.17)

Furthermore, for  $\Re a > 0$  the following formula holds (pp. 788 [86]):

$$\int_0^{+\infty} \frac{1}{a^2 + \rho^2} \mathcal{P}_{-\frac{1}{2} + i\rho}(\cosh \tau) \rho \tanh(\pi \rho) d\rho = \mathcal{Q}_{a-\frac{1}{2}}(\cosh \tau),$$

 $<sup>{}^{2}\</sup>mathbb{R}^{+}_{*}=(0,+\infty[$ 

where  $Q_{\nu}$  is the associated Legendre function of the second kind. From equations (2.14) and (2.15), we find that

$$\frac{1}{a^2 + \rho^2} = \int_0^{+\infty} \mathcal{Q}_{a-\frac{1}{2}}(\cosh \tau) \mathcal{P}_{-\frac{1}{2} + i\rho}(\cosh \tau) \sinh \tau d\tau.$$
(2.18)

A natural choice for the connectivity function W is then:

$$\mathbf{W}(z) = W(d_{\mathbb{D}}(z,0)) = \alpha_1 \mathcal{Q}_{a_1 - \frac{1}{2}}(\cosh d_{\mathbb{D}}(z,0)) - \alpha_2 \mathcal{Q}_{a_2 - \frac{1}{2}}(\cosh d_{\mathbb{D}}(z,0)),$$
(2.19)

with  $(\alpha_1, \alpha_2, a_1, a_2)$  satisfying the relations

$$\begin{cases}
\alpha_1 = 2\sigma_1 b_1, \\
\alpha_2 = 2\sigma_2 b_2, \\
a_1 = \sqrt{\sigma_1^2 + \frac{1}{4}}, \\
a_2 = \sqrt{\sigma_2^2 + \frac{1}{4}}.
\end{cases}$$
(2.20)

From equation (2.18) and the definition of the Mehler-Fock transform, we have that for all  $\rho > 0$ 

$$\widetilde{W}(\rho) \stackrel{def}{=} \mathcal{M}\mathbf{W}(\rho) = 2\left(\frac{b_1\sigma_1}{\sigma_1^2 + \frac{1}{4} + \rho^2} - \frac{b_2\sigma_2}{\sigma_2^2 + \frac{1}{4} + \rho^2}\right).$$
(2.21)

Comparing (2.21) with (2.3), we see that the transformed connectivity functions are equivalent.

As in the first section, we impose some conditions on the coefficients  $(b_1, b_2, \sigma_1, \sigma_2)$  and assume W satisfies Hypothesis 2.2

**Hypothesis 2.2.** We assume the connectivity function W and its corresponding Mehler-Fock transform  $\widetilde{W}(\rho)$  satisfy the following conditions:

- (i)  $\widetilde{W}(0) = \widetilde{W}_0 < 0$ ,
- (*ii*) W(0) > 0,

(iii) there exists  $\rho_c > 0$  such that  $\widetilde{W}_c \stackrel{def}{=} \widetilde{W}(\rho_c) = \max_{\rho \in \mathbb{R}^+} \widetilde{W}(\rho) > 0$ ,

(*iv*) 
$$\rho_c = \frac{\sqrt{3}}{2}$$
.

Using the constants  $\Gamma_1$  and  $\Gamma_2$  defined in equation (2.4),  $\widetilde{W}_c$  can be written as

$$\widetilde{W}_{c} = \frac{\Gamma_{1} + \Gamma_{2}}{1 + \sigma_{1}^{2} + \sigma_{2}^{2} + \sigma_{1}^{2}\sigma_{2}^{2}}.$$
(2.22)

The condition  $\rho_c = \frac{\sqrt{3}}{2}$ , reduces to equatin (2.6). It is also straightforward to see that equations (2.22) and (2.6) imply the relation given in equation (2.7).

We can now state the main result of the paper.

**Theorem 2.2** (Existence of spot solutions in hyperbolic geometry). Fix  $\theta \geq 0$  for the threshold of the nonlinearity S and  $(b_1, b_2, \sigma_1, \sigma_2) \in \mathbb{R}^4$  such that the connectivity function W defined in equation (2.19) satisfies Hypotheses 2.2. Then there exists  $\mu_* < \mu_c$  such that the planar neural field equation (2.11) has a stationary localized radial solution  $v(\tau) \in L^2(\mathbb{R}^+, \sinh(\tau)d\tau)$  for each  $\mu \in ]\mu_*, \mu_c[$ : these solutions stay close to v = 0 and, for each fixed  $\tau_* > 0$ , we have the asymptotics

$$v(\tau) = \beta \sqrt{\mu_c - \mu} \ \mathcal{P}_{-\frac{1}{2} + i\frac{\sqrt{3}}{2}}(\cosh \tau) + O(\mu - \mu_c) \ as \ \mu \to \mu_c, \tag{2.23}$$

uniformly in  $0 \leq \tau \leq \tau_*$  for an appropriate constant  $\beta$  with  $sign(\beta) = -sign(\Xi_{\tilde{d}_1^2})$ , where

$$\Xi_{\tilde{d}_{1}^{2}} = \frac{\mu_{c}^{2} s_{2}}{2} \left( -(\Gamma_{1} + 2\Gamma_{2}) \int_{0}^{\infty} (\mathcal{P}_{\nu}(\cosh(s))^{3} \sinh(s) ds + 2\Gamma_{2} \int_{0}^{\infty} \mathcal{P}_{\nu}(\cosh(s)) (\mathcal{P}_{\nu}^{1}(\cosh(s))^{2} \sinh(s) ds + o(1) \right) + 2\Gamma_{2} \int_{0}^{\infty} (\mathcal{P}_{\nu}(\cosh(s))^{3} \sinh(s) ds + o(1) \int_{0}^{\infty} (\mathcal{P$$

**Remark 2.2.** The matching arguments in §3 yield a similar theorem for the case  $\mu > \mu_c$  where one finds a bifurcating branch of solutions also given by (2.23). However, in these solutions are not  $L^2(\mathbb{R}^+, \sinh(\tau)d\tau)$ -functions.

# 3 Radial analysis

# 3.1 Localized radial solutions in Euclidean geometry

#### 3.1.1 Linear stability analysis of the trivial state

Linearising equation (2.1) about the trivial equilibrium v = 0 yields the equation

$$\frac{d}{dt}u(\mathbf{r},t) = -u(\mathbf{r},t) + \mu s_1 \int_{\mathbb{R}^2} W(\|\mathbf{r} - \mathbf{r}'\|) u(\mathbf{r}',t) d\mathbf{r}'.$$
(3.1)

Looking for a Turing instability we set  $u(\mathbf{r}, t) = e^{\sigma t} e^{i\mathbf{k}\cdot\mathbf{r}}$  to obtain the dispersion relation

$$\sigma(\|\mathbf{k}\|) = -1 + \mu s_1 \widehat{W}(\|\mathbf{k}\|). \tag{3.2}$$

The critical value  $\mu_c$  is then found to be given by

$$\mu_c = \frac{1}{s_1 \widehat{W}_c},\tag{3.3}$$

and for all  $\mu < \mu_c$ , the trivial solution v = 0 is stable. Hence, we set  $\lambda = \mu - \mu_c$  and rewrite equation (2.1)

$$\frac{d}{dt}v(\mathbf{r},t) = \mathbf{L}_{\mu_c}v(\mathbf{r},t) + \mathbf{R}(v(\mathbf{r},t),\lambda), \qquad (3.4)$$

where  $\mathbf{L}_{\mu_c}$  and  $\mathbf{R}$  are defined by

$$\begin{aligned} \mathbf{L}_{\mu_c} v(\mathbf{r},t) &= -v(\mathbf{r},t) + \mu_c s_1 \int_{\mathbb{R}^2} W(\|\mathbf{r}-\mathbf{r}'\|) v(\mathbf{r}',t) d\mathbf{r}', \\ \mathbf{R}(v(\mathbf{r},t),\lambda) &= \int_{\mathbb{R}^2} W(\|\mathbf{r}-\mathbf{r}'\|) \left[ S\left( (\lambda+\mu_c) v(\mathbf{r}',t) \right) - \mu_c s_1 v(\mathbf{r}',t) \right] d\mathbf{r}', \end{aligned}$$

with  $\mathbf{R}(0,0) = D_v \mathbf{R}(0,0) = 0$ . We define  $f(X,\lambda)$  by:

$$f(X,\lambda) = S\left((\lambda + \mu_c)X\right) - \mu_c s_1 X_2$$

where  $f(0,0) = \partial_X f(0,0) = 0$ .

# 3.1.2 PDE method

We assume that  $v \to v(\cdot, t) \in \mathcal{C}^1(\mathbb{R}^+, \mathcal{H}^4(\mathbb{R}^2))$  is a solution of (3.4), where  $\mathcal{H}^4(\mathbb{R}^2)$  is the Sobolev space defined as

$$\mathcal{H}^4(\mathbb{R}^2) = \{ u \in \mathcal{L}^2(\mathbb{R}^2) \mid \forall \ 0 \le |\alpha| \le 4 \ D^{\alpha} u \in \mathcal{L}^2(\mathbb{R}^2) \}.$$

Under this assumption, an application of the Fourier transform to equation (3.4) yields

$$(\partial_t + 1)\,\widehat{v}(\mathbf{k}, t) = \widehat{W}(\|\mathbf{k}\|) \left[ \mu_c s_1 \widehat{v}(\mathbf{k}, t) + \widehat{f(v, \lambda)}(\mathbf{k}, t) \right]$$

Using the inverse Fourier transform we then obtain

$$(\partial_t + 1) \mathcal{L}_{\mu_c}(v) = \mathcal{M}(v, \lambda), \tag{3.5}$$

with  $\mathcal{L}_{\mu_c}$  and  $\mathcal{M}$  defined by

$$\mathcal{L}_{\mu_c}(v) = (\sigma_1^2 \sigma_2^2 - \Gamma_1 \mu_c s_1)v - (\sigma_1^2 + \sigma_2^2 - \Gamma_2 \mu_c s_1)\Delta v + \Delta^2 v,$$

and

$$\mathcal{M}(v,\lambda) = \Gamma_1 f(v,\lambda) - \Gamma_2 \Delta f(v,\lambda),$$

where  $\Delta$  denotes the Laplcian on  $\mathbb{R}^2$ . From equations (2.7) and the fact that  $\mu_c = \left(s_1 \widehat{W}_c\right)^{-1}$ , the coefficients of  $\mathcal{L}_{\mu_c}$  reduce to

$$\sigma_1^2 \sigma_2^2 - \Gamma_1 \mu_c s_1 = 1, \sigma_1^2 + \sigma_2^2 - \Gamma_2 \mu_c s_1 = -2$$

We look for stationary radial solutions of equation (3.4), that is

$$v(r) + 2\Delta_r v(r) + \Delta_r^2 v(r) = \Gamma_1 f(v(r), \lambda) - \Gamma_2 \Delta_r f(v(r), \lambda)$$
(3.6)

with  $r = \|\mathbf{r}\| \in \mathbb{R}^+$  and  $\Delta_r = \partial_r^2 + \frac{1}{r}\partial_r$ . Note that  $\Delta_r f(v(r), \lambda)$  can be expressed as:

$$\Delta_r f(v,\lambda) = (\Delta_r v) f'(v,\lambda) + (\partial_r v)^2 f''(v,\lambda).$$

#### 3.1.3 The equation near the core

We rewrite equation (3.6) as a four dimensional spatial dynamical system where the radial coordinate plays the role of "time". To do this, we set

$$\partial_r u_1 = u_3,$$
  

$$\partial_r u_2 = u_4,$$
  

$$\left(\partial_r^2 + \frac{1}{r}\partial_r + 1\right)u_1 = u_2,$$
  

$$\left(\partial_r^2 + \frac{1}{r}\partial_r + 1\right)u_2 = \Gamma_1 f(u_1, \lambda) - \Gamma_2 \Delta_r f(u_1, \lambda),$$

and we can write

$$\partial_r U = \mathcal{A}(r)U + \mathcal{F}(U,\lambda), \qquad (3.7)$$

where  $U = (u_1, u_2, u_3, u_4)^T$ ,

$$\mathcal{A}(r) = \begin{pmatrix} 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \\ -1 & 1 & -\frac{1}{r} & 0 \\ 0 & -1 & 0 & -\frac{1}{r} \end{pmatrix}, \qquad \mathcal{F}(U,\lambda) = \begin{pmatrix} 0 \\ 0 \\ 0 \\ \mathcal{F}_4(u_1, u_2, u_3, u_4, \lambda) \end{pmatrix}$$

and

$$\mathcal{F}_4(u_1, u_2, u_3, u_4, \lambda) = \Gamma_1 f(u_1, \lambda) - \Gamma_2 \left[ (u_2 - u_1) f'(u_1, \lambda) + u_3^2 f''(u_1, \lambda) \right].$$

We can now apply the theory developed in [74, 88]. First, we set  $\lambda = 0$  and linearize (3.7) about U = 0 to get the linear system  $\partial_r U = \mathcal{A}(r)U$  which has four linearly independent solutions

$$V_{1}(r) = \sqrt{2\pi} (J_{0}(r), 0, -J_{1}(r), 0)^{\mathrm{T}}$$

$$V_{2}(r) = \sqrt{2\pi} (rJ_{1}(r), 2J_{0}(r), rJ_{0}(r), -2J_{1}(r))^{\mathrm{T}}$$

$$V_{3}(r) = \sqrt{2\pi} (Y_{0}(r), 0, -Y_{1}(r), 0)^{\mathrm{T}}$$

$$V_{4}(r) = \sqrt{2\pi} (rY_{1}(r), 2Y_{0}(r), rY_{0}(r), -2Y_{1}(r))^{\mathrm{T}}$$

Lemma 3.1, defines the centre-unstable "Core" manifold,  $\mathcal{W}^{cu}_{-}$  that describes all radial solutions that remain bounded as  $r \to 0$ .

**Lemma 3.1.** Fix  $r_0 > 0$ , then there exist constants  $\delta_0, \delta_1$  such that the set  $\mathcal{W}^{cu}_{-}(\lambda)$  of solutions U(r) of (3.7) for which  $\sup_{0 \le r \le r_0} |U(r)| < \delta_0$  is, for  $|\lambda| < \delta_0$ , is a smooth two-dimensional manifold. Furthermore,  $U \in \mathcal{W}^{cu}_{-}(\lambda)$  with  $|P^{cu}_{-}(r_0)U(r_0)| < \delta_1$  if and only if

$$U(r_0) = \tilde{d}_1 V_1(r_0) + \tilde{d}_2 V_2(r_0) + V_3(r_0) O_{r_0}(|\lambda| |\tilde{d}| + |\tilde{d}|^2) + V_4(r_0) \left( \Theta_{\tilde{d}_1^2} \tilde{d}_1^2 + \Theta_{\tilde{d}_1 \tilde{d}_2} \tilde{d}_1 \tilde{d}_2 + O(r_0)(|\lambda| |\tilde{d}| + |\tilde{d}_2|^2 + |\tilde{d}|^3) \right),$$
(3.8)

where

$$\Theta_{\tilde{d}_1^2} = \sqrt{2\pi}\mu_c^2 s_2 \frac{\Gamma_1 + \Gamma_2}{4} \left[ \frac{1}{\sqrt{3}} + O(r_0^{-1/2}) \right],$$
  
$$\Theta_{\tilde{d}_1 \tilde{d}_2} = -\sqrt{2\pi}\mu_c^2 s_2 \Gamma_2 \left[ \frac{1}{\sqrt{3}} + O(r_0^{-1/2}) \right],$$

for some  $\tilde{d} = (\tilde{d}_1, \tilde{d}_2) \in \mathbb{R}^2$  with  $|\tilde{d}| < \delta_1$ , where the right hand side in (3.8) depends smoothly on  $(\tilde{d}, \lambda)$ .

*Proof.* We observe that four independent solutions to the adjoint problem  $\partial_r U = -\mathcal{A}^T(r)U$  are given by

$$W_{1}(r) = \frac{\sqrt{2\pi}}{8} \left( -2rY_{1}(r), r^{2}Y_{0}(r), -2rY_{0}(r), -r^{2}Y_{1}(r) \right)^{\mathbf{T}}$$
  

$$W_{2}(r) = \frac{\sqrt{2\pi}}{8} \left( 0, -rY_{1}(r), 0, -rY_{0}(r) \right)^{\mathbf{T}},$$
  

$$W_{3}(r) = \frac{\sqrt{2\pi}}{8} \left( 2rJ_{1}(r), -r^{2}J_{0}(r), 2rJ_{0}(r), r^{2}J_{1}(r) \right)^{\mathbf{T}},$$
  

$$W_{4}(r) = \frac{\sqrt{2\pi}}{8} \left( 0, rJ_{1}(r), 0, rJ_{0}(r) \right)^{\mathbf{T}}.$$

It follows from

$$\frac{\pi}{2}r\left(J_1(r)Y_0(r) - J_0(r)Y_1(r)\right) = 1,$$

that

$$\langle V_i(r), W_j(r) \rangle_{\mathbb{R}^4} = \delta_{i,j} \quad i, j = 1, \dots, 4,$$

is independent of r. For given  $\tilde{d} = (\tilde{d}_1, \tilde{d}_2) \in \mathbb{R}^2$ , we consider the fixed-point equation

$$U(r) = \sum_{j=1}^{2} \tilde{d}_{j} V_{j}(r) + \sum_{j=1}^{2} V_{j}(r) \int_{r_{0}}^{r} \langle W_{j}(s), \mathcal{F}(U(s), \lambda) \rangle ds + \sum_{j=3}^{4} V_{j}(r) \int_{0}^{r} \langle W_{j}(s), \mathcal{F}(U(s), \lambda) \rangle ds, = \sum_{j=1}^{2} \tilde{d}_{j} V_{j}(r) + \sum_{j=1}^{2} V_{j}(r) \int_{r_{0}}^{r} W_{j,4}(s) \mathcal{F}_{4}(U(s), \lambda) ds + \sum_{j=3}^{4} V_{j}(r) \int_{0}^{r} W_{j,4}(s) \mathcal{F}_{4}(U(s), \lambda) ds$$
(3.9)

,

on  $\mathcal{C}^0([0, r_0], \mathbb{R}^4)$ , where  $W_{j,4}(r)$  (resp.  $\mathcal{F}_4(U(r), \lambda)$ ) denotes the fourth component of  $W_j(r)$  (resp.  $\mathcal{F}(U(r), \lambda)$ ). A direct adaptation of Lemma 1 in [74] gives:

- Each solution  $U \in \mathcal{C}^0([0, r_0], \mathbb{R}^4)$  of (3.9) gives a solution of (3.7) that is bounded on  $[0, r_0]$ .
- Every bounded solution  $U \in \mathcal{C}^0([0, r_0], \mathbb{R}^4)$  of (3.7) satisfies (3.9) provided that we add  $\tilde{d}_3 V_3(r) + \tilde{d}_4 V_4(r)$  to the right hand side for an appropriate  $\tilde{d} \in \mathbb{R}^4$ .
- Existence of solutions of (3.9) is given by the uniform contraction mapping principle for sufficiently small  $(\tilde{d}_1, \tilde{d}_2)$  and  $\lambda$ .
- The resulting solution U satisfies on  $[0, r_0]$

$$U(r) = \sum_{j=1}^{2} \tilde{d}_{j} V_{j}(r) + O_{r_{0}}(|\lambda||\tilde{d}| + |\tilde{d}|^{2}).$$

As in [74], we need to calculate the quadratic coefficient in  $\tilde{d}$  in front of  $V_4(r_0)$ , denoted  $\Theta$ . The quadratic term in U of  $\mathcal{F}(U,\lambda)$  at (0,0) is given by

$$\mathcal{F}_{20}(U,U) = \frac{\mu_c^2 s_2}{2} \left( 0, 0, 0, (\Gamma_1 + 2\Gamma_2) u_1^2 - 2\Gamma_2 u_1 u_2 - 2\Gamma_2 u_3^2 \right)^{\mathrm{T}}.$$

If we evaluate (3.9) at  $r = r_0$ , we arrive at (3.8) except that we need to calculate the quadratic coefficients in front of  $V_4(r_0)$ . Using a Taylor expansion, we find that these coefficients are given by

$$\Theta_{\tilde{d}_1^2} = \int_0^{\tau_0} W_{4,4}(s) \pi \mu_c^2 s_2 \left[ (\Gamma_1 + 2\Gamma_2) J_0(s)^2 - 2\Gamma_2 J_1(s)^2 \right] ds,$$

and

$$\Theta_{\tilde{d}_1\tilde{d}_2} = -4 \int_0^{r_0} W_{4,4}(s) \pi \mu_c^2 s_2 \Gamma_2 J_0(s)^2 ds.$$

Using the two formulas for integrals of Bessel functions

$$\int_{0}^{+\infty} J_{\nu}(as) J_{\nu}(bs) J_{\nu}(cs) s^{1-\nu} ds = \frac{2^{\nu-1} \Delta^{2\nu-1}}{(abc)^{\nu} \Gamma(\nu + \frac{1}{2}) \Gamma(\frac{1}{2})},$$
  
$$\int_{0}^{+\infty} J_{\mu}(as) J_{\nu}(bs) J_{\nu}(cs) s^{1-\mu} ds = \frac{(bc)^{\mu-1} \sin(A)^{\mu-1/2}}{\sqrt{2\pi} a^{\mu}} P_{\nu-\frac{1}{2}}^{\frac{1}{2}-\mu}(\cos(A)),$$

where  $\triangle$  is the area of triangle with lengths a, b and  $c, A = \arccos\left(\frac{b^2 + c^2 - a^2}{2bc}\right)$  and P is associated Legendre function (see [95]), we find with a = b = c = 1 and  $\nu = 0$  (i.e.,  $\triangle = \frac{\sqrt{3}}{4}$  and  $\Gamma(\frac{1}{2}) = \sqrt{\pi}$ )

$$\int_{0}^{+\infty} s J_{0}(s)^{3} ds = \frac{2}{\pi\sqrt{3}},$$
  
and with  $a = b = c = 1, \ \mu = 0$  and  $\nu = 1$  (*i.e.*,  $A = \frac{\pi}{3}$  and  $P_{\frac{1}{2}}^{\frac{1}{2}}(z) = \sqrt{\frac{2}{\pi}} \frac{z}{(1-z^{2})^{\frac{1}{4}}}$ )  
 $\int_{0}^{+\infty} s J_{0}(s) J_{1}(s)^{2} ds = \frac{1}{\pi\sqrt{3}}.$ 

We also have the estimates (see [77])

$$\int_{r_0}^{+\infty} s J_0(s)^3 ds = O(r_0^{-1/2}),$$

and

$$\int_{r_0}^{+\infty} s J_0(s) J_1(s)^2 ds = O(r_0^{-1/2}).$$

Thus we have

$$\begin{split} \Theta_{\tilde{d}_{1}^{2}} &= \frac{\sqrt{2\pi}\pi\mu_{c}^{2}s_{2}(\Gamma_{1}+2\Gamma_{2})}{8}\int_{0}^{r_{0}}sJ_{0}(s)^{3}ds - \frac{\sqrt{2\pi}\pi\mu_{c}^{2}s_{2}\Gamma_{2}}{4}\int_{0}^{r_{0}}sJ_{0}(s)J_{1}(s)^{2}ds, \\ &= \frac{\sqrt{2\pi}\pi\mu_{c}^{2}s_{2}(\Gamma_{1}+2\Gamma_{2})}{8}\left(\frac{2}{\pi\sqrt{3}}+O(r_{0}^{-1/2})\right) - \frac{\sqrt{2\pi}\pi\mu_{c}^{2}s_{2}\Gamma_{2}}{4}\left(\frac{1}{\pi\sqrt{3}}+O(r_{0}^{-1/2})\right), \\ &= \frac{\sqrt{2\pi}\mu_{c}^{2}s_{2}(\Gamma_{1}+\Gamma_{2})}{4}\left(\frac{1}{\sqrt{3}}+O(r_{0}^{-1/2})\right), \end{split}$$

and

$$\Theta_{\tilde{d}_1 \tilde{d}_2} = -\frac{\sqrt{2\pi} \pi \mu_c^2 s_2 \Gamma_2}{2} \int_0^{r_0} s J_0(s)^3 ds,$$
  
=  $-\sqrt{2\pi} \mu_c^2 s_2 \Gamma_2 \left(\frac{1}{\sqrt{3}} + O(r_0^{-1/2})\right).$ 

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#### 3.1.4 The far-field equations

We make the spatial dynamical system (3.7) autonomous by adding the equation  $\partial_r \alpha = -\alpha^2$  where  $\alpha = \frac{1}{r}$ . The augmented system becomes

$$\frac{d}{dr} \begin{pmatrix} u_1 \\ u_2 \\ u_3 \\ u_4 \\ \alpha \end{pmatrix} = \begin{pmatrix} u_3 \\ u_4 \\ -u_1 + u_2 - \alpha u_3 \\ -u_2 - \alpha u_4 + \mathcal{F}_4 (u_1, u_2, u_3, u_4, \lambda) \\ -\alpha^2 \end{pmatrix}.$$
 (3.10)

We now carry out a series of near-identity coordinate transformations to put (3.10) into normal form. To do this, we define the normal-form coordinates

$$U = \widetilde{A}\zeta_0 + \widetilde{B}\zeta_1 + \text{c.c.}, \quad U = (u_1, u_2, u_3, u_4)^T,$$

or equivalently

$$\begin{pmatrix} \widetilde{A} \\ \widetilde{B} \end{pmatrix} = \frac{1}{4} \begin{pmatrix} 2u_1 - i(2u_3 + u_4) \\ -u_4 - iu_2 \end{pmatrix},$$

where  $\mathcal{A}(\infty)\zeta_0 = i\zeta_0$  and  $\mathcal{A}(\infty)\zeta_1 = i\zeta_1 + \zeta_0$  with

$$\mathcal{A}(\infty) = \begin{pmatrix} 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \\ -1 & 1 & 0 & 0 \\ 0 & -1 & 0 & 0 \end{pmatrix}, \quad \zeta_0 = \begin{pmatrix} 1 \\ 0 \\ i \\ 0 \end{pmatrix}, \quad \zeta_1 = \begin{pmatrix} 0 \\ 2i \\ 1 \\ -2 \end{pmatrix}.$$

In these coordinates, equation (3.10) becomes

$$\partial_{r}\widetilde{A} = \left(i - \frac{\alpha}{2}\right)\widetilde{A} + \widetilde{B} + \frac{\alpha}{2}\overline{\widetilde{A}} + O((|\lambda| + |\widetilde{A}| + |\widetilde{B}|)(|\widetilde{A}| + |\widetilde{B}|)), 
\partial_{r}\widetilde{B} = \left(i - \frac{\alpha}{2}\right)\widetilde{B} - \frac{\alpha}{2}\overline{\widetilde{B}} + O((|\lambda| + |\widetilde{A}| + |\widetilde{B}|)(|\widetilde{A}| + |\widetilde{B}|)), 
\partial_{r}\alpha = -\alpha^{2}.$$
(3.11)

**Lemma 3.2.** Fix  $0 < m < \infty$ , then there exists a change of coordinates

$$\begin{pmatrix} A\\B \end{pmatrix} = e^{i\phi(r)} [1 + \mathcal{T}(\alpha)] \begin{pmatrix} \widetilde{A}\\\widetilde{B} \end{pmatrix} + O((|\lambda| + |\widetilde{A}| + |\widetilde{B}|)(|\widetilde{A}| + |\widetilde{B}|)),$$
(3.12)

so that (3.11) becomes

$$\partial_r A = -\frac{\alpha}{2}A + B + \mathcal{R}_A(A, B, \alpha, \lambda),$$
  

$$\partial_r B = -\frac{\alpha}{2}B + c_1^1 \lambda A + c_3^0 |A|^2 A + \mathcal{R}_B(A, B, \alpha, \lambda),$$
  

$$\partial_r \alpha = -\alpha^2.$$
(3.13)

The constants  $c_1^1$  and  $c_3^0$  are given by

$$c_1^1 = -\frac{s_1(\Gamma_1 + \Gamma_2)}{4},\tag{3.14}$$

and

$$c_3^0 = -\frac{\mu_c^3(\Gamma_1 + \Gamma_2)}{4} \left[ \frac{s_3}{2} + \frac{\mu_c s_2^2 (19\Gamma_1 + 4\Gamma_2)}{18} \right].$$
 (3.15)

The coordinate change is polynomial in  $(A, B, \alpha)$  and smooth in  $\lambda$  and  $\mathcal{T}(\alpha) = O(\alpha)$  is linear and upper triangular for each  $\alpha$ , while  $\phi(r)$  satisfies

$$\partial_r \phi(r) = 1 + O(|\lambda| + |\alpha|^3 + |A|^2), \quad \phi(0) = 0.$$

The remainder terms are given by

$$\mathcal{R}_{A}(A, B, \alpha, \lambda) = O\left(\sum_{j=0}^{2} |A^{j}B^{3-j}| + |\alpha|^{3}|A| + |\alpha|^{2}|B| + (|A| + |B|)^{5} + |\lambda||\alpha|^{m}(|A| + |B|)\right),$$
  
$$\mathcal{R}_{B}(A, B, \alpha, \lambda) = O\left(\sum_{j=0}^{1} |A^{j}B^{3-j}| + |\alpha|^{3}|B| + |\lambda|(|\lambda| + |\alpha|^{3} + |A|^{2})|A| + (|A| + |B|)^{5} + |\lambda||\alpha|^{m}|B|\right).$$

*Proof.* See Scheel [88] for the change of variables. The coefficients  $c_1^1$  and  $c_3^0$  have been computed in [44].  $\Box$ We set  $\lambda = -\epsilon$ ,  $\epsilon > 0$  and  $c_1^0 = -c_1^1$ . We define

$$A = \sqrt{\epsilon}a, \quad B = \epsilon b, r = \frac{\rho}{\sqrt{\epsilon}}, \tag{3.16}$$

for which (3.13) becomes

$$\partial_{\rho}a = b - \frac{a}{2\rho} + \mathcal{R}_{1}(a, b, \rho, \epsilon),$$
  

$$\partial_{\rho}b = -\frac{b}{2\rho} + c_{1}^{0}a + c_{3}^{0}|a|^{2}a + \mathcal{R}_{2}(a, b, \rho, \epsilon),$$
(3.17)

where

$$\mathcal{R}_{1}(a, b, \rho, \epsilon) = \epsilon^{-1} \mathcal{R}_{A}(\sqrt{\epsilon}a, \epsilon b, \frac{\sqrt{\epsilon}}{\rho}, \epsilon) = O(\epsilon(|a| + |b|)),$$
  
$$\mathcal{R}_{2}(a, b, \rho, \epsilon) = \epsilon^{-3/2} \mathcal{R}_{B}(\sqrt{\epsilon}a, \epsilon b, \frac{\sqrt{\epsilon}}{\rho}, \epsilon) = O(\epsilon(|a| + |b|)),$$
  
(3.18)

uniformly in  $\rho \ge \rho_1$  for each fixed  $\rho_1 > 0$ . We also use the variables

$$\begin{pmatrix} a \\ b \end{pmatrix} = \frac{1}{\sqrt{\rho}} \begin{pmatrix} \hat{a} \\ \hat{b} \end{pmatrix}, \tag{3.19}$$

in which (3.17) becomes

$$\partial_{\rho}\hat{a} = \hat{b} + O_{\rho}(\epsilon(|\hat{a}| + |\hat{b}|)),$$
  

$$\partial_{\rho}\hat{b} = c_{1}^{0}\hat{a} + \frac{c_{3}^{0}}{\rho}|\hat{a}|^{2}\hat{a} + O_{\rho}(\epsilon(|\hat{a}| + |\hat{b}|)).$$
(3.20)

The estimates for the remainder terms given above are valid for  $\rho \ge \rho_1$  for each fixed  $\rho_1 > 0$ . To capture the region  $r_0\sqrt{\epsilon} \le \rho \le \rho_1$ , we use the variables

$$\begin{pmatrix} \tilde{a} \\ \tilde{b} \end{pmatrix} = \begin{pmatrix} a \\ \rho(b - \frac{a}{2\rho}) \end{pmatrix}, \quad \tau = \log \rho,$$
(3.21)

so that  $\rho = e^{\tau}$  and  $\rho \to 0$  corresponds to  $\tau \to -\infty$ . In these variables, (3.20) becomes

$$\partial_{\tau}\tilde{a} = \tilde{b} + \tilde{\mathcal{R}}_{1}(\tilde{a}, \tilde{b}, \rho, \epsilon),$$
  

$$\partial_{\tau}\tilde{b} = \frac{\tilde{a}}{4} + \rho^{2} \left( c_{1}^{0}\tilde{a} + c_{3}^{0} |\tilde{a}|^{2}\tilde{a} \right) + \tilde{\mathcal{R}}_{2}(\tilde{a}, \tilde{b}, \rho, \epsilon),$$
  

$$\partial_{\tau}\rho = \rho,$$
  
(3.22)

with  $(\tilde{a}, \tilde{b}, \rho) \in \mathbb{C}^2 \times \mathbb{R}^+$ .

#### 3.1.5 Matching the core and the far-field

We start by linearizing the far-field equation (3.20) about  $(\hat{a}, \hat{b}) = 0$  to get the equation

$$\partial_{\rho} \begin{pmatrix} \hat{a} \\ \hat{b} \end{pmatrix} = \begin{pmatrix} 0 & 1 \\ c_1^0 & 0 \end{pmatrix} \begin{pmatrix} \hat{a} \\ \hat{b} \end{pmatrix} + O(\epsilon) \begin{pmatrix} \hat{a} \\ \hat{b} \end{pmatrix}.$$
(3.23)

For  $\epsilon = 0$ , the general solution of (3.23) is given by

$$\begin{pmatrix} \hat{a} \\ \hat{b} \end{pmatrix}(\rho) = q_1 e^{-\rho\sqrt{c_1^0}} \begin{pmatrix} 1 \\ -\sqrt{c_1^0} \end{pmatrix} + q_2 e^{\rho\sqrt{c_1^0}} \begin{pmatrix} 1 \\ \sqrt{c_1^0} \end{pmatrix}.$$

Thus, for each  $\rho = \rho_1 > 0$  and for sufficiently small  $\epsilon > 0$ , we can write the  $\rho = \rho_1$ -fiber of the stable manifold  $\mathcal{W}^s_+(\epsilon)$  of (3.20) near 0 as

$$\mathcal{W}^{s}_{+}(\epsilon) \mid_{\rho=\rho_{1}} : \left( \begin{array}{c} \hat{a} \\ \hat{b} \end{array} \right) = \eta \left( \begin{array}{c} 1 \\ -\sqrt{c_{1}^{0}} \end{array} \right) + O_{\rho_{1}}(\epsilon|\eta| + |\eta|^{3}) \left( \begin{array}{c} 1 \\ \sqrt{c_{1}^{0}} \end{array} \right),$$

where  $\eta \in \mathbb{C}$ . Using (3.19) and (3.21) and redefining  $\eta$ , we obtain the expression

$$\mathcal{W}^{s}_{+}(\epsilon) \mid_{\rho=\rho_{1}} : \left( \begin{array}{c} \tilde{a} \\ \tilde{b} \end{array} \right) = \eta \left( \begin{array}{c} 1 \\ -\frac{1}{2} - \rho_{1}\sqrt{c_{1}^{0}} \end{array} \right) + O_{\rho_{1}}(\epsilon|\eta| + |\eta|^{3}) \left( \begin{array}{c} 1 \\ -\frac{1}{2} + \rho_{1}\sqrt{c_{1}^{0}} \end{array} \right),$$

for  $\mathcal{W}^s_+(\epsilon)$  in the  $(\tilde{a}, \tilde{b})$ -coordinates. We introduce the coordinates

$$u = \begin{pmatrix} u_1 \\ u_2 \end{pmatrix} = \frac{1}{2} \begin{pmatrix} 1 & 2 \\ 1 & -2 \end{pmatrix} \begin{pmatrix} \tilde{a} \\ \tilde{b} \end{pmatrix}, \qquad (3.24)$$

and yield

$$\mathcal{W}^{s}_{+}(\epsilon) \mid_{\rho=\rho_{1}} : \begin{pmatrix} u_{1} \\ u_{2} \end{pmatrix} = \eta \begin{pmatrix} -\rho_{1}\sqrt{c_{1}^{0}} \\ 1+\rho_{1}\sqrt{c_{1}^{0}} \end{pmatrix} + O_{\rho_{1}}(\epsilon|\eta|+|\eta|^{3}) \begin{pmatrix} \rho_{1}\sqrt{c_{1}^{0}} \\ 1-\rho_{1}\sqrt{c_{1}^{0}} \end{pmatrix}.$$

The  $(u_1, u_2)$ -coordinates transform equation (3.22) into

$$\partial_{\tau} u = \left[ D + O(\epsilon e^{-2\tau}) \right] u + O((\sqrt{\epsilon} + e^{\tau})e^{\tau}|u|), \quad D = \begin{pmatrix} \frac{1}{2} & 0\\ 0 & -\frac{1}{2} \end{pmatrix}.$$
(3.25)

Lemma 3.3. The linear equation

$$\partial_{\tau} u = \left[ D + O(\epsilon e^{-2\tau}) \right] u \tag{3.26}$$

has an exponential dichotomy with exponents  $\pm \frac{1}{2}$  on  $[\tau_0, \tau_1]$ . Furthermore, the coordinate transformation  $u \to \tilde{u}$  that brings (3.26) into the form

$$\partial_{\tau}\tilde{u} = \left[D + O(\epsilon e^{-2\tau})Id\right]\tilde{u},\tag{3.27}$$

can be chosen such that

$$u(\tau_0) = \tilde{u}(\tau_0), \quad u(\tau_1) = \begin{pmatrix} 1 & O(\epsilon) \\ O(r_0^{-2}) & 1 \end{pmatrix} \tilde{u}(\tau_1).$$
(3.28)

*Proof.* See [74].

We recall that we have

$$\mathcal{W}^{s}_{+}(\epsilon) \mid_{\rho=\rho_{1}} : \left(\begin{array}{c} u_{1} \\ u_{2} \end{array}\right) = \eta \left(\begin{array}{c} -\rho_{1}\sqrt{c_{1}^{0}} \\ 1+\rho_{1}\sqrt{c_{1}^{0}} \end{array}\right) + O_{\rho_{1}}(\epsilon|\eta|+|\eta|^{3}) \left(\begin{array}{c} \rho_{1}\sqrt{c_{1}^{0}} \\ 1-\rho_{1}\sqrt{c_{1}^{0}} \end{array}\right).$$

for the stable manifold in the u-variables, which thanks to (3.28), becomes

$$\begin{split} \mathcal{W}^{s}_{+}(\epsilon) \mid_{\rho=\rho_{1}} &: \left( \begin{array}{c} \tilde{u}_{1} \\ \tilde{u}_{2} \end{array} \right) = \eta \left( \begin{array}{c} -\rho_{1}\sqrt{c_{1}^{0}} + O_{\rho_{1}}(\epsilon) \\ 1 + \rho_{1}\sqrt{c_{1}^{0}}(1 + O(r_{0}^{-2})) \end{array} \right) \\ &+ O_{\rho_{1}}(\epsilon|\eta| + |\eta|^{3}) \left( \begin{array}{c} \rho_{1}\sqrt{c_{1}^{0}} + O_{\rho_{1}}(\epsilon) \\ 1 - \rho_{1}\sqrt{c_{1}^{0}}(1 + O(r_{0}^{-2})) \end{array} \right), \end{split}$$

in the  $\tilde{u}\text{-variables}.$  Choosing  $0<\rho_1\ll 1$  sufficiently small, we can solve

$$\eta \left( 1 + \rho_1 \sqrt{c_1^0} (1 + O(r_0^{-2})) \right) + O_{\rho_1}(\epsilon |\eta| + |\eta|^3) \left( 1 - \rho_1 \sqrt{c_1^0} (1 + O(r_0^{-2})) \right) = \tilde{\eta},$$

for  $\eta$  so that

$$\eta = \frac{\eta}{1 + \rho_1 \sqrt{c_1^0} \left(1 + O(r_0^{-2})\right)} + O_{\rho_1}(\epsilon |\tilde{\eta}| + |\tilde{\eta}|^3),$$

and consequently

$$\mathcal{W}^{s}_{+}(\epsilon)|_{\rho=\rho_{1}} : \left(\begin{array}{c} \tilde{u}_{1} \\ \tilde{u}_{2} \end{array}\right) = \tilde{\eta} \left(\begin{array}{c} -\rho_{1}\sqrt{c_{1}^{0}}(1+O(\rho_{1}+r_{0}^{-2})) \\ 1 \end{array}\right) + O_{\rho_{1}}(\epsilon|\tilde{\eta}|+|\tilde{\eta}|^{3}) \left(\begin{array}{c} 1 \\ 0 \end{array}\right).$$
(3.29)

Using (3.29) and (3.28), we find that the stable manifold  $W^s_+(\epsilon)$  at  $r = r_0$  is given by

$$\mathcal{W}^s_+(\epsilon)\mid_{r=r_0} : u = \sqrt{\epsilon}r_0\hat{\eta} \left( \begin{array}{c} -\sqrt{c_1^0} + O(\rho_1 + r_0^{-2}) + O_{\rho_1}(\epsilon + \sqrt{\epsilon}|\hat{\eta}|^2) \\ 0 \end{array} \right) + \hat{\eta} \left( \begin{array}{c} 0 \\ 1 \end{array} \right).$$

We apply successive changes of variables to transform this expression into the (A, B)-coordinates and obtain

$$\mathcal{W}^{s}_{+}(\epsilon)|_{r=r_{0}} : \left(\begin{array}{c}A\\B\end{array}\right) = \epsilon \hat{\eta} \left[-\sqrt{c_{1}^{0}} + O(\rho_{1} + r_{0}^{-2}) + O_{\rho_{1}}(\epsilon + \sqrt{\epsilon}|\hat{\eta}|^{2})\right] \left(\begin{array}{c}r_{0}\\1\end{array}\right) + \sqrt{\epsilon} \hat{\eta} \left(\begin{array}{c}1\\0\end{array}\right). \quad (3.30)$$

We have now gathered all the results necessary for the proof of Theorem 2.1. In order to do this, we need to find nontrivial intersections of the stable manifold  $\mathcal{W}^s_+(\epsilon)$  with the center-unstable manifold  $\mathcal{W}^{cu}_-(\epsilon)$ . To this end, we write the expansion (3.8) for each fixed  $r_0 \gg 1$  in the  $(\tilde{A}, \tilde{B})$  coordinates and afterwards in the coordinates (A, B). Using the expansions of Bessel functions and the variables  $(d_1, d_2) = (\tilde{d}_1/\sqrt{r_0}, \sqrt{r_0}\tilde{d}_2)$ , we arrive at the expression

$$\begin{pmatrix} \tilde{A} \\ \tilde{B} \end{pmatrix} = e^{i(r_0 - \pi/4)} \begin{pmatrix} d_1(1 + O(r_0^{-1})) + d_2(-i + O(r_0^{-1})) \\ -d_2 r_0^{-1}(i + O(r_0^{-1})) - \left(\frac{1}{\sqrt{3}} + O(\frac{1}{\sqrt{r_0}})\right) \left(C_1 \sqrt{r_0} d_1^2 + \frac{C_2}{\sqrt{r_0}} d_1 d_2\right) \end{pmatrix} + e^{i(r_0 - \pi/4)} \begin{pmatrix} O_{r_0}(\epsilon |d| + |d|^2) \\ O_{r_0}(\epsilon |d| + |d_2|^2 + |d_1|^3) \end{pmatrix},$$
(3.31)

with  $C_1 = \sqrt{2\pi}\mu_c^2 s_2 \frac{\Gamma_1 + \Gamma_2}{4}$  and  $C_2 = -\sqrt{2\pi}\mu_c^2 s_2 \Gamma_2$ . We can apply the transformation (3.12) and obtain the expression

$$\mathcal{W}_{-}^{cu}(\epsilon) \mid_{r=r_{0}} : \begin{pmatrix} A \\ B \end{pmatrix} = e^{i(-\pi/4 + O(r_{o}^{-2}) + O_{r_{0}}(\epsilon|d| + |d|^{2})} \begin{pmatrix} O_{r_{0}}(\epsilon|d| + |d|^{2}) \\ O_{r_{0}}(\epsilon|d| + |d_{2}|^{2} + |d_{1}|^{3}) \end{pmatrix}$$
$$+ e^{i(-\frac{\pi}{4} + O(r_{o}^{-2}) + O_{r_{0}}(\epsilon|d| + |d|^{2})} \begin{pmatrix} d_{1}(1 + O(r_{0}^{-1})) + d_{2}(-i + O(r_{0}^{-1})) \\ -d_{2}r_{0}^{-1}(i + O(r_{0}^{-1})) \end{pmatrix}$$
$$+ e^{i(-\frac{\pi}{4} + O(r_{o}^{-2}) + O_{r_{0}}(\epsilon|d| + |d|^{2})} \begin{pmatrix} 0 \\ -\left(\frac{1}{\sqrt{3}} + O(\frac{1}{\sqrt{r_{0}}})\right) \begin{pmatrix} C_{1}\sqrt{r_{0}}d_{1}^{2} + \frac{C_{2}}{\sqrt{r_{0}}}d_{1}d_{2} \end{pmatrix} \end{pmatrix}.$$
(3.32)

After redefining  $\hat{\eta}$  to  $\eta$  to remove the phase in  $\mathcal{W}^{cu}_{-}(\epsilon)|_{r=r_0}$ , it remains to solve

$$\begin{cases} \sqrt{\epsilon}\eta + \epsilon r_0 \eta \left[ -\sqrt{c_1^0} + O(\rho_1 + r_0^{-2}) + O_{\rho_1}(\sqrt{\epsilon}) \right] = d_1(1 + O(r_0^{-1})) + d_2(-i + 0(r_0^{-1})) \\ + O_{r_0}(\epsilon |d| + |d|^2), \\ \epsilon \eta \left[ -\sqrt{c_1^0} + O(\rho_1 + r_0^{-2}) + O_{\rho_1}(\sqrt{\epsilon}) \right] = -d_2 r_0^{-1}(i + O(r_0^{-1})) + O_{r_0}(\epsilon |d| + |d_2|^2 + |d_1|^3) \\ - \left( \frac{1}{\sqrt{3}} + O(r_0^{-1/2}) \right) \left( C_1 \sqrt{r_0} d_1^2 + \frac{C_2}{\sqrt{r_0}} d_1 d_2 \right). \end{cases}$$

Setting  $d_j = \sqrt{\epsilon} \hat{d}_j$  and writing  $\eta = \eta_1 + i\eta_2$ , we obtain

$$\hat{d}_1 - i\hat{d}_2 - i\eta_2 + O(r_0^{-1})\hat{d} + O_{r_0}(\epsilon|\hat{d}| + \sqrt{\epsilon}|\hat{d}|^2) = \eta_1 + \sqrt{\epsilon}r_0O(\eta),$$

and

$$\begin{split} \sqrt{\epsilon}r_0\eta \left[ -\sqrt{c_1^0} + O(\rho_1 + r_0^{-2}) + O_{\rho_1}(\sqrt{\epsilon}) \right] &= -i\hat{d}_2 + O(r_0^{-1})\hat{d}_2 \\ &+ O_{r_0}(\epsilon|\hat{d}| + \sqrt{\epsilon}|\hat{d}_2|^2 + \epsilon|\hat{d}_1|^3) - \left(\frac{1}{\sqrt{3}} + O(r_0^{-1/2})\right) \left(C_1r_0^{3/2}\sqrt{\epsilon}\hat{d}_1^2 + C_2\sqrt{\epsilon}\sqrt{r_0}\hat{d}_1\hat{d}_2\right). \end{split}$$

Hence, we find the expansion

$$\hat{d}_1 = \eta_1 + O_{r_0}(\sqrt{\epsilon}\eta_1), \quad \hat{d}_2 = O_{r_0}(\sqrt{\epsilon}\eta_1), \quad \eta_2 = O(r_0^{-1})\eta_1 + O_{r_0}(\sqrt{\epsilon}\eta_1),$$

and we can solve the equation

$$\sqrt{\epsilon}r_0\eta_1\left[-\sqrt{c_1^0} + O(r_0^{-1}) + O(\rho_1) + O_{\rho_1}(\sqrt{\epsilon})\right] = -\left(\frac{1}{\sqrt{3}} + O(r_0^{-1/2})\right)C_1r_0^{3/2}\sqrt{\epsilon}\eta_1^2 + O_{r_0}(\epsilon\eta_1),$$

that has two solutions  $\eta_1 = 0$  (the trivial solution) and  $\eta_1$  small given to leading order by

$$\eta_1 = \frac{1}{C_1} \sqrt{\frac{3c_1^0}{r_0}} + O(\rho_1) + O(r_0^{-1}) + O_{\rho_1}(\sqrt{\epsilon})$$

This gives  $\tilde{d}_1$  and  $\tilde{d}_2$ 

$$\tilde{d}_1 = \frac{s_1}{\mu_c^2 s_2} \sqrt{\frac{3\epsilon}{2\pi c_1^0}}, \quad \tilde{d}_2 = O(\epsilon),$$

or equivalently  $(\epsilon=-\lambda,c_1^0=-c_1^1)$ 

$$\tilde{d}_1 = \frac{s_1}{\mu_c^2 s_2} \sqrt{\frac{3\lambda}{2\pi c_1^1}}, \quad \tilde{d}_2 = O(\lambda).$$

This completes the proof of Theorem 2.1.

# 3.2 Localized radial solutions in the Poincaré disk

### 3.2.1 Linear stability of the trivial state

We look at the linearization of (2.11) about the trivial state v = 0 governed by the equation

$$\frac{d}{dt}u(z,t) = -u(z,t) + \mu s_1 \int_{\mathbb{D}} W(d_{\mathbb{D}}(z,z'))u(z',t)\mathrm{dm}(z').$$
(3.33)

Looking at linear stability with to periodic solutions on the Poincaré disk  $u(z,t) = e^{\sigma t} e^{\left(\frac{1}{2}+i\rho\right)\langle z,b\rangle}$  (see [27]) we obtain the following dispersion relation

$$\sigma(\rho) = -1 + \mu s_1 \mathcal{M} \mathbf{W}(\rho). \tag{3.34}$$

The the critical value  $\mu_c$  is then found to occur at

$$\mu_c = \frac{1}{s_1 \widetilde{W}_c},\tag{3.35}$$

and for all  $\mu < \mu_c$ , the trivial solution v = 0 is stable.

We set  $\lambda = \mu - \mu_c$  and rewrite equation (2.11)

$$\frac{d}{dt}v(z,t) = \mathbf{L}_{\mu_c}v(z,t) + \mathbf{R}(v(z,t),\lambda), \qquad (3.36)$$

where  $\mathbf{L}_{\mu_c}$  and  $\mathbf{R}$  are defined by

$$\mathbf{L}_{\mu_c} v(z,t) = -v(z,t) + \mu_c s_1 \int_{\mathbb{D}} W(d_{\mathbb{D}}(z,z')) v(z',t) \mathrm{dm}(z'),$$
  
$$\mathbf{R}(v(z,t),\lambda) = \int_{\mathbb{D}} W(d_{\mathbb{D}}(z,z')) \left[ S\left( (\lambda + \mu_c) v(z',t) \right) - \mu_c s_1 v(z',t) \right] \mathrm{dm}(z'),$$

with  $\mathbf{R}(0,0) = D_v \mathbf{R}(0,0) = 0$ . We define  $f(X,\lambda)$  by

$$f(X,\lambda) = S\left((\lambda + \mu_c)X\right) - \mu_c s_1 X,$$

where  $f(0,0) = \partial_X f(0,0) = 0$ .

#### 3.2.2 PDE methods in the Poincaré disk

We assume that v is a sufficiently smooth radially symmetric solution of equation (3.36) such that if we apply the Mehler-Fock transform we obtain

$$\left(\frac{d}{dt}+1\right)\mathcal{M}v(\rho,t) = \widetilde{W}(\rho)\left[\mu_c s_1 \mathcal{M}v(\rho,t) + \mathcal{M}f(v,\lambda)(\rho,t)\right]$$

We also assume that we can apply an inverse Mehler-Fock transform, to yield [41]

$$\left(\frac{d}{dt}+1\right)\left(1+L_{\mathbb{D}}^{\tau}\right)^{2}v=\Gamma_{1}f(v,\lambda)-\Gamma_{2}L_{\mathbb{D}}^{\tau}f(v,\lambda).$$
(3.37)

 $L^{\tau}_{\mathbb{D}}$  has been defined in equation (2.16) and we have used the fact that

$$\sigma_1^2 \sigma_2^2 - \Gamma_1 \mu_c s_1 = 1, \sigma_1^2 + \sigma_2^2 - \Gamma_2 \mu_c s_1 = -2$$

We look for stationary radial solutions of equation (3.36), that is

$$(1+L_{\mathbb{D}}^{\tau})^{2} v(\tau) = \Gamma_{1} f(v(\tau), \lambda) - \Gamma_{2} L_{\mathbb{D}}^{\tau} f(v(\tau), \lambda).$$
(3.38)

Note that  $L^{\tau}_{\mathbb{D}}f(v(\tau),\lambda)$  can be expressed as

$$L^{\tau}_{\mathbb{D}}f(v,\lambda) = (L^{\tau}_{\mathbb{D}}v)f'(v,\lambda) + (\partial_{\tau}v)^2 f''(v,\lambda).$$

### 3.2.3 The equation near the core

We rewrite equation (3.38) as a four dimensional system of non-autonomous differential equations to yield

$$\partial_{\tau} u_1 = u_3,$$
  

$$\partial_{\tau} u_2 = u_4,$$
  

$$\left(\partial_{\tau}^2 + \coth(\tau)\partial_{\tau} + 1\right) u_1 = u_2,$$
  

$$\left(\partial_{\tau}^2 + \coth(\tau)\partial_{\tau} + 1\right) u_2 = \Gamma_1 f(u_1, \lambda) - \Gamma_2 L_{\mathbb{D}}^{\tau} f(u_1, \lambda),$$

and we may rewrite (3.38) as a spatial dynamical system

$$\partial_{\tau} U = \mathcal{A}(\tau) U + \mathcal{F}(U, \lambda), \tag{3.39}$$

where  $U = (u_1, u_2, u_3, u_4)^T$ 

$$\mathcal{A}(\tau) = \begin{pmatrix} 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \\ -1 & 1 & -\coth(\tau) & 0 \\ 0 & -1 & 0 & -\coth(\tau) \end{pmatrix} \text{ and, } \mathcal{F}(U,\lambda) = \begin{pmatrix} 0 \\ 0 \\ 0 \\ \mathcal{F}_4(u_1, u_2, u_3, u_4, \lambda), \end{pmatrix}$$

and

$$\mathcal{F}_4(u_1, u_2, u_3, u_4, \lambda) = \Gamma_1 f(u_1, \lambda) - \Gamma_2 \left[ (u_2 - u_1) f'(u_1, \lambda) + u_3^2 f''(u_1, \lambda) \right]$$

First, we set  $\lambda = 0$  and linearize (3.39) about U = 0 to get the linear system  $\partial_{\tau} U = \mathcal{A}(\tau)U$ . Proposition 3.1 defines the four linearly independent solutions of the linear system  $\partial_{\tau} U = \mathcal{A}(\tau)U$ .

**Proposition 3.1.** The linear system  $\partial_{\tau} U = \mathcal{A}(\tau)U$  has four linearly independent solutions given by

$$V_{1}(\tau) = \left(\mathcal{P}_{\nu}(\cosh\tau), 0, \mathcal{P}_{\nu}^{1}(\cosh\tau), 0\right)^{\mathbf{T}}$$

$$V_{2}(\tau) = \left(V_{2}^{1}(\tau), \mathcal{P}_{\nu}(\cosh\tau), V_{2}^{3}(\tau), \mathcal{P}_{\nu}^{1}(\cosh\tau)\right)^{\mathbf{T}}$$

$$V_{3}(\tau) = \left(\mathcal{Q}_{\nu}(\cosh\tau), 0, \mathcal{Q}_{\nu}^{1}(\cosh\tau), 0\right)^{\mathbf{T}}$$

$$V_{4}(\tau) = \left(V_{4}^{1}(\tau), \mathcal{Q}_{\nu}(\cosh\tau), V_{4}^{3}(\tau), \mathcal{Q}_{\nu}^{1}(\cosh\tau)\right)^{\mathbf{T}}$$

where

$$\nu = -\frac{1}{2} + i\frac{\sqrt{3}}{2} \tag{3.40}$$

and

$$V_{2}^{1}(\tau) = \mathcal{P}_{\nu}(\cosh \tau) \int_{0}^{\tau} \mathcal{P}_{\nu}(\cosh s) \mathcal{Q}_{\nu}(\cosh s) \sinh(s) ds - \mathcal{Q}_{\nu}(\cosh \tau) \int_{0}^{\tau} \left(\mathcal{P}_{\nu}(\cosh s)\right)^{2} \sinh(s) ds,$$

$$V_{2}^{3}(\tau) = \mathcal{P}_{\nu}^{1}(\cosh \tau) \int_{0}^{\tau} \mathcal{P}_{\nu}(\cosh s) \mathcal{Q}_{\nu}(\cosh s) \sinh(s) ds - \mathcal{Q}_{\nu}^{1}(\cosh \tau) \int_{0}^{\tau} \left(\mathcal{P}_{\nu}(\cosh s)\right)^{2} \sinh(s) ds,$$

$$V_{4}^{1}(\tau) = \mathcal{P}_{\nu}(\cosh \tau) \int_{0}^{\tau} \left(\mathcal{Q}_{\nu}(\cosh s)\right)^{2} \sinh(s) ds - \mathcal{Q}_{\nu}(\cosh \tau) \int_{0}^{\tau} \mathcal{P}_{\nu}(\cosh s) \mathcal{Q}_{\nu}(\cosh s) \sinh(s) ds,$$

$$V_{4}^{3}(\tau) = \mathcal{P}_{\nu}^{1}(\cosh \tau) \int_{0}^{\tau} \left(\mathcal{Q}_{\nu}(\cosh s)\right)^{2} \sinh(s) ds - \mathcal{Q}_{\nu}^{1}(\cosh \tau) \int_{0}^{\tau} \mathcal{P}_{\nu}(\cosh s) \mathcal{Q}_{\nu}(\cosh s) \sinh(s) ds.$$

*Proof.* From equation (2.13), the associated Legendre functions  $\mathcal{P}_{\nu}(\cosh \cdot)$  and  $\mathcal{Q}_{\nu}(\cosh \cdot)$  form a basis of solutions for the equation

$$\partial_{\tau}^{2}\Psi(\tau) + \coth(\tau)\partial_{\tau}\Psi(\tau) - \nu(\nu+1)\Psi(\tau) = 0$$

If  $\nu = -\frac{1}{2} + i\frac{\sqrt{3}}{2}$  then we have  $\nu(\nu+1) = -1$ . This implies that  $\mathcal{P}_{-\frac{1}{2}+i\frac{\sqrt{3}}{2}}(\cosh \tau)$  and  $\mathcal{Q}_{-\frac{1}{2}+i\frac{\sqrt{3}}{2}}(\cosh \tau)$  are solutions of

$$\partial_{\tau}^{2}\Psi(\tau) + \coth(\tau)\partial_{\tau}\Psi(\tau) + \Psi(\tau) = 0.$$

From now on  $\nu = -\frac{1}{2} + i\frac{\sqrt{3}}{2}$ . The solution of linear system  $\partial_{\tau}U = \mathcal{A}(\tau)U$  can be found by inspecting the equivalent system

$$\left(\partial_{\tau}^{2} + \coth(\tau)\partial_{\tau} + 1\right)u_{1} = u_{2}, \quad \left(\partial_{\tau}^{2} + \coth(\tau)\partial_{\tau} + 1\right)u_{2} = 0$$

Consequently, solutions  $V_1(\tau) = \left(\mathcal{P}_{\nu}(\cosh \tau), 0, \mathcal{P}_{\nu}^1(\cosh \tau), 0\right)^{\mathbf{T}}$  and  $V_3(\tau) = \left(\mathcal{Q}_{\nu}(\cosh \tau), 0, \mathcal{Q}_{\nu}^1(\cosh \tau), 0\right)^{\mathbf{T}}$ are found to be two linearly independent solutions, where we have used the relations

$$\partial_{\tau} \left( \mathcal{P}_{\nu}(\cosh \tau) \right) = \mathcal{P}_{\nu}^{1}(\cosh \tau), \quad \partial_{\tau} \left( \mathcal{Q}_{\nu}(\cosh \tau) \right) = \mathcal{Q}_{\nu}^{1}(\cosh \tau)$$

In order to find the other two linearly independent solutions, we have to solve the equation

$$\left(\partial_{\tau}^2 + \coth \tau \partial_{\tau} + 1\right) u(\tau) = \mathcal{P}_{\nu}(\cosh \tau).$$

Searching for solutions of the form  $u(\tau) = c_1(\tau)\mathcal{P}_{\nu}(\cosh \tau) + c_2(\tau)\mathcal{Q}_{\nu}(\cosh \tau)$ , we find that  $c_1(\tau)$  and  $c_2(\tau)$  satisfy the system

$$\begin{aligned} \dot{c}_1(\tau) \mathcal{P}_{\nu}(\cosh \tau) + \dot{c}_2(\tau) \mathcal{Q}_{\nu}(\cosh \tau) &= 0, \\ \dot{c}_1(\tau) \mathcal{P}_{\nu}^1(\cosh \tau) + \dot{c}_2(\tau) \mathcal{Q}_{\nu}^1(\cosh \tau) &= \mathcal{P}_{\nu}(\cosh \tau), \end{aligned}$$
(3.41)

where dots denote derivatives with respect to  $\tau$ . Rewriting (3.41) in matrix form yields

$$\begin{pmatrix} \mathcal{P}_{\nu}(\cosh\tau) & \mathcal{Q}_{\nu}(\cosh\tau) \\ \mathcal{P}_{\nu}^{1}(\cosh\tau) & \mathcal{Q}_{\nu}^{1}(\cosh\tau) \end{pmatrix} \begin{pmatrix} \dot{c}_{1}(\tau) \\ \dot{c}_{2}(\tau) \end{pmatrix} = \begin{pmatrix} 0 \\ \mathcal{P}_{\nu}(\cosh\tau) \end{pmatrix},$$
(3.42)

and from formula in [38, page 123], we obtain that

$$W(\tau) = \mathcal{P}_{\nu}(\cosh \tau)\mathcal{Q}_{\nu}^{1}(\cosh \tau) - \mathcal{P}_{\nu}^{1}(\cosh \tau)\mathcal{Q}_{\nu}(\cosh \tau) = -\frac{1}{\sinh \tau}$$

where  $W(\tau)$  is non vanishing for all  $\tau \in ]0; +\infty[$  such that we can invert the matrix system on left hand side of (3.42) to obtain:

$$\dot{c}_{1}(\tau) = \sinh \tau \mathcal{P}_{\nu}(\cosh \tau) \mathcal{Q}_{\nu}(\cosh \tau),$$
  
$$\dot{c}_{2}(\tau) = -\sinh \tau \left( \mathcal{P}_{\nu}(\cosh \tau) \right)^{2}.$$

Hence we have

$$u(\tau) = c_1 \mathcal{P}_{\nu}(\cosh \tau) + c_2 \mathcal{Q}_{\nu}(\cosh \tau) + \mathcal{P}_{\nu}(\cosh \tau) \int_0^{\tau} \mathcal{P}_{\nu}(\cosh s) \mathcal{Q}_{\nu}(\cosh s) \sinh(s) ds$$
$$-\mathcal{Q}_{\nu}(\cosh \tau) \int_0^{\tau} \left(\mathcal{P}_{\nu}(\cosh s)\right)^2 \sinh(s) ds,$$

where  $c_1, c_2 \in \mathbb{R}$  are two constants. Equivalently, the solutions of

$$\left(\partial_{\tau}^{2} + \coth \tau \partial_{\tau} + 1\right) u(\tau) = \mathcal{Q}_{\nu}(\cosh \tau)$$

are given by

$$u(\tau) = c_3 \mathcal{P}_{\nu}(\cosh \tau) + c_4 \mathcal{Q}_{\nu}(\cosh \tau) + \mathcal{P}_{\nu}(\cosh \tau) \int_0^{\tau} (\mathcal{Q}_{\nu}(\cosh s))^2 \sinh(s) ds$$
$$-\mathcal{Q}_{\nu}(\cosh \tau) \int_0^{\tau} \mathcal{P}_{\nu}(\cosh s) \mathcal{Q}_{\nu}(\cosh s) \sinh(s) ds,$$

with constants  $c_3, c_4 \in \mathbb{R}$ .

Chosing only linearly independent solutions, we finally obtain the result stated in the proposition.  $\Box$ 

In Table 1, we summarize the expansions of the associated Legendre functions in the limits  $\tau \to 0$  and  $\tau \to \infty$ ; Proposition 3.2 defines the constants in the  $\tau \to \infty$  limit. Thus,  $V_1(\tau)$  and  $V_2(\tau)$  stay bounded as  $\tau \to 0$ , while the norms of  $V_3(\tau)$  and  $V_4(\tau)$  behave like  $\ln \tau$  as  $\tau \to 0$ . We expect that the set of solutions of (3.39) that are bounded as  $\tau \to 0$  forms a two-dimensional manifold in  $\mathbb{R}^4$  for each fixed  $\tau > 0$ . We denote the projection  $P_{-}^{cu}(\tau_0)$  onto the space spanned by  $V_1(\tau_0), V_2(\tau_0)$  with null space given by the span of  $V_3(\tau_0), V_4(\tau_0)$ .

**Proposition 3.2.** The constants  $C_0, C_1, \Phi_0$  and  $\Phi_1$ , given in Table 1, are

$$C_0 = 2\sqrt{\frac{2\sqrt{3}}{3\pi \tanh\left(\frac{\sqrt{3}\pi}{2}\right)}} \text{ and } C_1 = \sqrt{\frac{2\pi\sqrt{3} \tanh\left(\frac{\sqrt{3}\pi}{2}\right)}{3}}, \tag{3.43}$$

$$\Phi_0 = \arg\left(\frac{\Gamma\left(i\frac{\sqrt{3}}{2}\right)}{\Gamma\left(\frac{1}{2} + i\frac{\sqrt{3}}{2}\right)}\right) \quad and \quad \Phi_0 + \Phi_1 = -\frac{\pi}{2}.$$
(3.44)

	$\tau \to 0$	$ au  ightarrow \infty$
$\mathcal{P}_{\nu}(\cosh \cdot)$	$1 + O(\tau^2)$	$C_0 \cos(\frac{\sqrt{3}\tau}{2} + \Phi_0)e^{-\frac{\tau}{2}} + O(e^{-\frac{3\tau}{2}})$
$\mathcal{P}^1_{\nu}(\cosh \cdot)$	$\tau\left(-\frac{1}{2}+O(\tau^2)\right)$	$C_0 \cos(\frac{\sqrt{3}\tau}{2} + \Phi_0 + \frac{2\pi}{3})e^{-\frac{\tau}{2}} + O(e^{-\frac{3\tau}{2}})$
$\mathcal{Q}_{\nu}(\cosh \cdot)$	$(-1 + O(\tau^2)) \ln \tau + O(1)$	$C_1 \cos(\frac{\sqrt{3}\tau}{2} - \Phi_1) e^{-\frac{\tau}{2}} + O(e^{-\frac{3\tau}{2}})$
$\mathcal{Q}^1_{\nu}(\cosh \cdot)$	$(1+O(\tau^2))\tau \ln \tau - \frac{1}{\tau} + O(1)$	$C_1 \cos(\frac{\sqrt{3}\tau}{2} - \Phi_1 + \frac{2\pi}{3})e^{-\frac{\tau}{2}} + O(e^{-\frac{3\tau}{2}})$

Table 1: Expansions of associated Legendre functions  $\mathcal{P}^k_{\nu}(\cosh \cdot)$  and  $\mathcal{Q}^k_{\nu}(\cosh \cdot)$  for  $\tau \to 0$  and  $\tau \to \infty$ ; see [38, 94].  $C_0, C_1, \Phi_0$  and  $\Phi_1$  are all real constants given in Proposition 3.2.

*Proof.* The proof of this proposition is based on the asymptotics as  $z \to \infty$  [92]

$$\mathcal{P}_{\nu}^{m}(z) \approx \frac{z^{m/2}}{\sqrt{\pi}(z-1)^{m/2}} \left( \frac{(2z)^{\nu} \Gamma(\nu+1/2)}{\Gamma(\nu-m+1)} \left(1+O(z^{-1})\right) + \frac{(2z)^{-\nu-1} \Gamma(-\nu-1/2)}{\Gamma(-m-\nu)} \left(1+O(z^{-1})\right) \right)$$

and

$$\begin{aligned} \mathcal{Q}_{\nu}^{m}(z) &\approx \frac{2^{-\nu-2}z^{-\nu-1}}{\sqrt{\pi}\Gamma(-m-\nu)\Gamma(\nu-m+1)} \left(\frac{1+z}{z-1}\right)^{m/2} \left((2z)^{2\nu+1}\Gamma(-m-\nu)\Gamma(\nu+1/2)\right) \\ &\times \left(\ln(1+z) - \ln(z-1)\right) \left(1 + O(z^{-1})\right) + \Gamma(-m+\nu+1)\Gamma(-\nu-1/2) \\ &\times \left(2\pi + \ln(1+z) - \ln(z-1)\right) \left(1 + O(z^{-1})\right) \right). \end{aligned}$$

We directly obtain the identities

$$C_{0} = \frac{2}{\sqrt{\pi}} \left| \frac{\Gamma\left(i\frac{\sqrt{3}}{2}\right)}{\Gamma\left(\frac{1}{2} + i\frac{\sqrt{3}}{2}\right)} \right|, \quad \Phi_{0} = \arg\left(\frac{\Gamma\left(i\frac{\sqrt{3}}{2}\right)}{\Gamma\left(\frac{1}{2} + i\frac{\sqrt{3}}{2}\right)}\right),$$

$$C_{1} = \sqrt{\pi} \left| \cot\left(\pi\left(-\frac{1}{2} + i\frac{\sqrt{3}}{2}\right)\right) \frac{\Gamma\left(i\frac{\sqrt{3}}{2}\right)}{\Gamma\left(\frac{1}{2} + i\frac{\sqrt{3}}{2}\right)} \right|,$$

$$\Phi_{1} = \arg\left[\sqrt{\pi}\cot\left(\pi\left(-\frac{1}{2} + i\frac{\sqrt{3}}{2}\right)\right) \frac{\Gamma\left(i\frac{\sqrt{3}}{2}\right)}{\Gamma\left(\frac{1}{2} + i\frac{\sqrt{3}}{2}\right)} \right].$$

The formulas for  $C_0$  and  $C_1$  are obtained by the standard identity involving the Gamma function

$$\left| \frac{\Gamma\left(i\frac{\sqrt{3}}{2}\right)}{\Gamma\left(\frac{1}{2}+i\frac{\sqrt{3}}{2}\right)} \right| = \sqrt{\frac{2\sqrt{3}}{3\tanh\left(\frac{\sqrt{3}\pi}{2}\right)}}$$

The formula (3.43) is then obtained from the above equation. The second relation in (3.44) can be verified from the above expressions of  $\Phi_0$  and  $\Phi_1$ .

**Remark 3.1.** One can deduce from (3.43) the following identity

$$\frac{\sqrt{3}C_0}{6} = \frac{C_1 C_0^2}{8}.$$
(3.45)

We are now able to present the hyperbolic equivalent of Lemma 3.1 for the centre-unstable manifold  $\mathcal{W}^{cu}_{-}$ .

**Lemma 3.4.** Fix  $\tau_0 > 0$ , then there exist constants  $\delta_0, \delta_1$  such that the set  $\mathcal{W}^{cu}_{-}(\lambda)$  of solutions  $U(\tau)$  of (3.39) for which  $\sup_{0 \le \tau \le \tau_0} |U(\tau)| < \delta_0$  is, for  $|\lambda| < \delta_0$ , is a smooth two-dimensional manifold. Furthermore,  $U \in \mathcal{W}^{cu}_{-}(\lambda)$  with  $|P^{cu}_{-}(\tau_0)U(\tau_0)| < \delta_1$  if and only if

$$U(\tau_0) = \tilde{d}_1 V_1(\tau_0) + \tilde{d}_2 V_2(\tau_0) + V_3(\tau_0) O_{\tau_0}(|\lambda| |\tilde{d}| + |\tilde{d}|^2) + V_4(\tau_0) \left( \Xi_{\tilde{d}_1^2} \tilde{d}_1^2 + \Xi_{\tilde{d}_1 \tilde{d}_2} \tilde{d}_1 \tilde{d}_2 + O_{\tau_0}(|\lambda| |\tilde{d}| + |\tilde{d}_2|^2 + |\tilde{d}|^3) \right),$$
(3.46)

where

$$\begin{aligned} \Xi_{\tilde{d}_{1}^{2}} &= \frac{\mu_{c}^{2} s_{2}}{2} \left[ -(\Gamma_{1} + 2\Gamma_{2}) \mathcal{I}_{1} + 2\Gamma_{2} \mathcal{I}_{2} + o(1) \right], \\ \Xi_{\tilde{d}_{1} \tilde{d}_{2}} &= \mu_{c}^{2} s_{2} \left[ \Gamma_{2} \mathcal{I}_{1} + o(1) \right], \end{aligned}$$

with

$$\mathcal{I}_{1} = \int_{0}^{\infty} \left(\mathcal{P}_{\nu}(\cosh s)\right)^{3} \sinh s ds < \infty,$$
  
$$\mathcal{I}_{2} = \int_{0}^{\infty} \mathcal{P}_{\nu}(\cosh s) \left(\mathcal{P}_{\nu}^{1}(\cosh s)\right)^{2} \sinh s ds < \infty.$$

for some  $\tilde{d} = (\tilde{d}_1, \tilde{d}_2) \in \mathbb{R}^2$  with  $|\tilde{d}| < \delta_1$ , where the right hand side in (3.46) depends smoothly on  $(\tilde{d}, \lambda)$ .

*Proof.* The proof is similar to that of Lemma 3.1. We observe that four independent solutions to the adjoint problem  $\partial_{\tau} U = -\mathcal{A}^{T}(\tau)U$  are given by

$$W_{1}(\tau) = \sinh \tau \left( \mathcal{Q}_{\nu}^{1}(\cosh \tau), W_{1}^{2}(\tau), -\mathcal{Q}_{\nu}(\cosh \tau), W_{1}^{4}(\tau) \right)^{\mathbf{T}},$$
  

$$W_{2}(\tau) = \sinh \tau \left( 0, \mathcal{Q}_{\nu}^{1}(\cosh \tau), 0, -\mathcal{Q}_{\nu}(\cosh \tau) \right)^{\mathbf{T}},$$
  

$$W_{3}(\tau) = \sinh \tau \left( \mathcal{P}_{\nu}^{1}(\cosh \tau), W_{3}^{2}(\tau), -\mathcal{P}_{\nu}(\cosh \tau), W_{3}^{4}(\tau) \right)^{\mathbf{T}},$$
  

$$W_{4}(\tau) = \sinh \tau \left( 0, \mathcal{P}_{\nu}^{1}(\cosh \tau), 0, -\mathcal{P}_{\nu}(\cosh \tau) \right)^{\mathbf{T}},$$

where

$$W_1^2(\tau) = \mathcal{P}_{\nu}^1(\cosh\tau) \int_0^{\tau} \mathcal{P}_{\nu}^1(\cosh s) \mathcal{Q}_{\nu}^1(\cosh s) \sinh(s) ds - \mathcal{Q}_{\nu}^1(\cosh\tau) \int_0^{\tau} \left(\mathcal{P}_{\nu}^1(\cosh s)\right)^2 \sinh(s) ds,$$

$$W_1^4(\tau) = -\mathcal{Q}_{\nu}(\cosh\tau) - \mathcal{P}_{\nu}(\cosh\tau) \int_0^{\tau} \mathcal{P}_{\nu}^1(\cosh s) \mathcal{Q}_{\nu}^1(\cosh s) \sinh(s) ds + \mathcal{Q}_{\nu}(\cosh\tau) \int_0^{\tau} \left(\mathcal{P}_{\nu}^1(\cosh s)\right)^2 \sinh(s) ds,$$

$$W_3^2(\tau) = \mathcal{P}_{\nu}^1(\cosh\tau) \int_0^{\tau} \left(\mathcal{Q}_{\nu}^1(\cosh s)\right)^2 \sinh(s) ds - \mathcal{Q}_{\nu}^1(\cosh\tau) \int_0^{\tau} \mathcal{P}_{\nu}^1(\cosh s) \mathcal{Q}_{\nu}^1(\cosh s) \sinh(s) ds,$$

$$W_3^4(\tau) = -\mathcal{P}_{\nu}(\cosh\tau) - \mathcal{P}_{\nu}(\cosh\tau) \int_0^{\tau} \left(\mathcal{Q}_{\nu}^1(\cosh s)\right)^2 \sinh(s) ds - \mathcal{Q}_{\nu}(\cosh\tau) \int_0^{\tau} \mathcal{P}_{\nu}^1(\cosh s) \mathcal{Q}_{\nu}^1(\cosh s) \sinh(s) ds.$$
The electric of  $W(\tau)$  and the full state of the s

The calculation of  $W_i(\tau)$  uses the following facts:

• the following relations are satisfied [38]:

$$\partial_{\tau} \left( \sinh \tau \mathcal{P}^{1}_{\nu} (\cosh \tau) \right) = - \sinh \tau \mathcal{P}_{\nu} (\cosh \tau), \partial_{\tau} \left( \sinh \tau \mathcal{Q}^{1}_{\nu} (\cosh \tau) \right) = - \sinh \tau \mathcal{Q}_{\nu} (\cosh \tau),$$

 $\bullet\,$  the solutions of

$$\left( \partial_{\tau}^{2} - \coth \tau \partial_{\tau} + 1 \right) q_{1}(\tau) = \sinh \tau \mathcal{P}_{\nu}^{1}(\cosh \tau), \left( \partial_{\tau}^{2} - \coth \tau \partial_{\tau} + 1 \right) q_{2}(\tau) = \sinh \tau \mathcal{Q}_{\nu}^{1}(\cosh \tau),$$

are given by

$$q_{1}(\tau) = c_{1} \sinh \tau \mathcal{P}_{\nu}^{1}(\cosh \tau) + c_{2} \sinh \tau \mathcal{Q}_{\nu}^{1}(\cosh \tau) + \sinh \tau \mathcal{P}_{\nu}^{1}(\cosh \tau) \int_{0}^{\tau} \mathcal{P}_{\nu}^{1}(\cosh s) \mathcal{Q}_{\nu}^{1}(\cosh s) \sinh(s) ds - \sinh \tau \mathcal{Q}_{\nu}^{1}(\cosh \tau) \int_{0}^{\tau} \left( \mathcal{P}_{\nu}^{1}(\cosh s) \right)^{2} \sinh(s) ds, q_{2}(\tau) = c_{3} \sinh \tau \mathcal{P}_{\nu}^{1}(\cosh \tau) + c_{4} \sinh \tau \mathcal{Q}_{\nu}^{1}(\cosh \tau) + \sinh \tau \mathcal{P}_{\nu}^{1}(\cosh \tau) \int_{0}^{\tau} \left( \mathcal{Q}_{\nu}^{1}(\cosh s) \right)^{2} \sinh(s) ds - \sinh \tau \mathcal{Q}_{\nu}^{1}(\cosh \tau) \int_{0}^{\tau} \mathcal{P}_{\nu}^{1}(\cosh s) \mathcal{Q}_{\nu}^{1}(\cosh s) \sinh(s) ds,$$

with four real constants  $c_1, c_2, c_3, c_4 \in \mathbb{R}$ .

It follows from

$$\sinh \tau \left( \mathcal{P}^{1}_{\nu}(\cosh \tau) \mathcal{Q}_{\nu}(\cosh \tau) - \mathcal{P}_{\nu}(\cosh \tau) \mathcal{Q}^{1}_{\nu}(\cosh \tau) \right) = 1,$$

that

$$\langle V_i(\tau), W_j(\tau) \rangle_{\mathbb{R}^4} = \delta_{i,j} \quad i, j = 1, \dots, 4,$$

is independent of  $\tau$ . For a given  $\tilde{d} = (\tilde{d}_1, \tilde{d}_2) \in \mathbb{R}^2$ , we consider the fixed-point equation:

$$U(\tau) = \sum_{j=1}^{2} \tilde{d}_{j} V_{j}(\tau) + \sum_{j=1}^{2} V_{j}(\tau) \int_{\tau_{0}}^{\tau} \langle W_{j}(s), \mathcal{F}(U(s), \lambda) \rangle ds + \sum_{j=3}^{4} V_{j}(\tau) \int_{0}^{\tau} \langle W_{j}(s), \mathcal{F}(U(s), \lambda) \rangle ds$$
  
$$= \sum_{j=1}^{2} \tilde{d}_{j} V_{j}(\tau) + \sum_{j=1}^{2} V_{j}(\tau) \int_{\tau_{0}}^{\tau} W_{j,4}(s) \mathcal{F}_{4}(U(s), \lambda) ds + \sum_{j=3}^{4} V_{j}(\tau) \int_{0}^{\tau} W_{j,4}(s) \mathcal{F}_{4}(U(s), \lambda) ds$$
  
(3.47)

on  $\mathcal{C}^{0}([0, \tau_{0}], \mathbb{R}^{4})$ , where  $W_{j,4}(\tau)$  (resp.  $\mathcal{F}_{4}(U(\tau), \lambda)$ ) denotes the fourth component of  $W_{j}(\tau)$  (resp.  $\mathcal{F}(U(\tau), \lambda)$ ). Hence, we have that:

- Each solution  $U \in \mathcal{C}^0([0, \tau_0], \mathbb{R}^4)$  of (3.47) gives a solution of (3.39) that is bounded on  $[0, \tau_0]$ .
- Every bounded solution  $U \in \mathcal{C}^0([0, \tau_0], \mathbb{R}^4)$  of (3.39) satisfies (3.47) provided that we add  $\tilde{d}_3 V_3(\tau) + \tilde{d}_4 V_4(\tau)$  to the right hand side for an appropriate  $\tilde{d} \in \mathbb{R}^4$ .
- Existence of solutions of (3.47) is given by the uniform contraction mapping principle for sufficiently small  $(\tilde{d}_1, \tilde{d}_2)$  and  $\lambda$ .
- The resulting solution U satisfies  $U(\tau) = \sum_{j=1}^{2} \tilde{d}_j V_j(\tau) + O_{\tau_0}(|\lambda||\tilde{d}| + |\tilde{d}|^2)$  on  $[0, \tau_0]$ .

We also need to compute the quadratic coefficient in  $\tilde{d}$  in front of  $V_4(\tau_0)$ , denoted by  $\Xi$ . The quadratic term in U of  $\mathcal{F}(U, \lambda)$  at (0, 0) is given by

$$\mathcal{F}_{20}(U,U) = \frac{\mu_c^2 s_2}{2} \left( 0, 0, 0, (\Gamma_1 + 2\Gamma_2) u_1^2 - 2\Gamma_2 u_1 u_2 - 2\Gamma_2 u_3^2 \right)^{\mathbf{T}}.$$

If we evaluate (3.47) at  $\tau = \tau_0$ , we arrive at (3.46) except that we need to calculate the quadratic coefficients in front of  $V_4(\tau_0)$ . Using a Taylor expansion, we find that these coefficients are given by

$$\Xi_{\tilde{d}_{1}^{2}} = \frac{\mu_{c}^{2} s_{2}}{2} \int_{0}^{\tau_{0}} W_{4,4}(s) \left[ \left( \Gamma_{1} + 2\Gamma_{2} \right) \left( \mathcal{P}_{\nu}(\cosh s) \right)^{2} - 2\Gamma_{2} \left( \mathcal{P}_{\nu}^{1}(\cosh s) \right)^{2} \right] ds,$$

and

$$\Xi_{\tilde{d}_1\tilde{d}_2} = -\mu_c^2 s_2 \Gamma_2 \int_0^{\tau_0} W_{4,4}(s) \left(\mathcal{P}_{\nu}(\cosh s)\right)^2 ds.$$

Hence, we have

$$\begin{split} \Xi_{\tilde{d}_{1}^{2}} &=& \frac{\mu_{c}^{2}s_{2}}{2}\left[-(\Gamma_{1}+2\Gamma_{2})\mathcal{I}_{1}+2\Gamma_{2}\mathcal{I}_{2}+o(1)\right]\\ \Xi_{\tilde{d}_{1}\tilde{d}_{2}} &=& \mu_{c}^{2}s_{2}\left[\Gamma_{2}\mathcal{I}_{1}+o(1)\right], \end{split}$$

where

$$\mathcal{I}_{1} = \int_{0}^{\infty} \left(\mathcal{P}_{\nu}(\cosh s)\right)^{3} \sinh s ds < \infty,$$
  
$$\mathcal{I}_{2} = \int_{0}^{\infty} \mathcal{P}_{\nu}(\cosh s) \left(\mathcal{P}_{\nu}^{1}(\cosh s)\right)^{2} \sinh s ds < \infty.$$

#### 3.2.4 The far-field equations and matching

We make spatial dynamical system (3.39) autonomous by augmenting the system with the equation  $\partial_{\tau} \alpha = -\alpha(2+\alpha)$  where  $\alpha = \coth \tau - 1$  to yield the new system

$$\frac{d}{d\tau} \begin{pmatrix} u_1 \\ u_2 \\ u_3 \\ u_4 \\ \alpha \end{pmatrix} = \begin{pmatrix} u_3 \\ -u_1 + u_2 - (1+\alpha)u_3 \\ -u_2 - (1+\alpha)u_4 + \mathcal{F}_4(u_1, u_2, u_3, u_4, \lambda, \alpha) \\ -\alpha(2+\alpha) \end{pmatrix}.$$
 (3.48)

In the remainder of this section, we focus on the regime  $\alpha \approx 0$  which corresponds to the far field  $\tau \gg 1$ . We denote  $\mathcal{A}(\infty, \lambda)$  by the matrix

$$\mathcal{A}(\infty,\lambda) = \begin{pmatrix} 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & 1 \\ -1 & 1 & -1 & 0 \\ s_1(\Gamma_1 + \Gamma_2)\lambda & -1 - s_1\Gamma_2\lambda & 0 & -1 \end{pmatrix},$$

where  $\partial_{\tau} U = \mathcal{A}(\tau, \lambda)U$  is the linearisation of (3.48) about the trivial state. We find that the matrix  $\mathcal{A}(\infty, 0)$ has four eigenvalues  $\nu, \bar{\nu}$  with multiplicity two ( $\nu$  is defined in equation (3.40)); see Figure 2. As  $\Re(\nu) = -1/2$ , the trivial state U = 0 is asymptotically stable at  $\lambda = 0$  and then there is no bifurcation at the far field. Recall that in the Euclidean case a Turing instability occurs at infinity. In Figure 2, we summarize how the eigenvalues  $\ell$  of  $\mathcal{A}(\infty, \lambda)$  split close to  $\lambda = 0$ . For  $\lambda > 0$ , there exist four complex conjugate eigenvalues with  $\Re(\ell) = -1/2$ . For  $\lambda < 0$ , there exist four complex conjugate eigenvalues with  $\Re(\ell) \neq -1/2$  and stable manifold  $\mathcal{W}^s_+$  is the union of the stable fast manifold (which we denote by  $\mathcal{W}^{sf}_+$ ) and the stable slow manifold (which we denote by  $\mathcal{W}^{ss}_+$ ) corresponding to the fast and slow decay to the trivial state.



Figure 2: Schematic splitting of the eigenvalues  $\ell$  of  $\mathcal{A}(\infty, \lambda)$  for different values of  $\lambda$ . At  $\lambda = 0$ , the multiplicity is two. Eigenvalues in the red box correspond to the stable fast manifold  $\mathcal{W}^{sf}_{+}(\lambda)$ .

First we argue that the centre-unstable manifold  $\mathcal{W}^{cu}_{-}$  and stable manifold  $\mathcal{W}^s_{+}$  should intersect. We have that  $V_1(\tau)$  and  $V_3(\tau)$  decay like  $e^{-\tau/2}$ , while  $V_2(\tau)$  and  $V_4(\tau)$  decay like  $\tau e^{-\tau/2}$  as  $\tau \to \infty$ . Hence the tangent space of the stable manifold at  $(u, \lambda) = 0$  is spanned by  $(V_1(\tau), V_2(\tau), V_3(\tau), V_4(\tau))$ . On the other hand, we showed in Lemma 3.4 that the tangent space of the core manifold is spanned by  $V_1(\tau)$  and  $V_2(\tau)$ . Then these tangent spaces would intersect along the two-dimensional subspace spanned by  $V_1(\tau)$  and  $V_2(\tau)$ .

In order to show that the centre-unstable manifold  $\mathcal{W}^{cu}_{-}$  intersects with the stable fast manifold  $\mathcal{W}^{sf}_{+}$ , we need to find an explicit description of  $\mathcal{W}^{sf}_{+}$ . To do this, we use successive, well chosen change of variables to put (3.48) into normal form. We first define the linear change of coordinates

$$U = \widetilde{A} \begin{pmatrix} 1 \\ 0 \\ \nu \\ 0 \end{pmatrix} + \widetilde{B} \begin{pmatrix} 0 \\ 2\nu + 1 \\ 1 \\ \nu(2\nu + 1) \end{pmatrix} + \text{c.c.},$$

or equivalently,

$$\begin{pmatrix} \widetilde{A} \\ \widetilde{B} \end{pmatrix} = \begin{pmatrix} \frac{1}{2}u_1 + i\frac{\sqrt{3}}{3}\left(-\frac{1}{2}u_1 - \frac{1}{3}u_2 - u_3 - \frac{2}{3}u_4\right) \\ -\frac{1}{3}\left(\frac{1}{2}u_2 + u_4\right) - i\frac{\sqrt{3}}{6}u_2 \end{pmatrix}, \text{ and } U = (u_1, u_2, u_3, u_4)^T$$

In these coordinates, the linear part of (3.48) becomes at  $\lambda = 0$ ,

$$\partial_{\tau}\widetilde{A} = \left(-\frac{1}{2} - \frac{\alpha}{2} + i\left[\frac{\sqrt{3}}{2} - \frac{\alpha\sqrt{3}}{6}\right]\right)\widetilde{A} + \left(1 + \frac{\alpha}{3}\right)\widetilde{B} + \alpha\left(\frac{1}{2} - \frac{\alpha\sqrt{3}}{6}\right)\overline{\widetilde{A}} - \frac{\alpha}{3}\overline{\widetilde{B}},$$
  

$$\partial_{\tau}\widetilde{B} = \left(-\frac{1}{2} - \frac{\alpha}{2} + i\left[\frac{\sqrt{3}}{2} - \frac{\alpha\sqrt{3}}{6}\right]\right)\widetilde{B} + \alpha\left(-\frac{1}{2} + \frac{\alpha\sqrt{3}}{6}\right)\overline{\widetilde{B}},$$
  

$$\partial_{\tau}\alpha = -\alpha(2 + \alpha).$$
(3.49)

We can find a similar transformation to (3.12) for the Euclidean case, given in Lemma 3.5.

**Lemma 3.5.** Fix  $0 < m < \infty$ , then there exists a change of coordinates

$$\begin{pmatrix} A\\B \end{pmatrix} = e^{-i\phi(r)} [1 + \mathcal{T}(\alpha)] \begin{pmatrix} \widetilde{A}\\\widetilde{B} \end{pmatrix} + O((|\lambda| + |\widetilde{A}| + |\widetilde{B}|)(|\widetilde{A}| + |\widetilde{B}|)),$$
(3.50)

so that (3.49) becomes

$$\partial_{\tau}A = \left(-\frac{1}{2} - \frac{\alpha}{2}\right)A + B + h.o.t.,$$
  

$$\partial_{\tau}B = \left(-\frac{1}{2} - \frac{\alpha}{2}\right)B + c_{1}\lambda A + h.o.t.,$$
  

$$\partial_{\tau}\alpha = -\alpha(2+\alpha).$$
(3.51)

The constant  $c_1$  is given by

$$c_1 = -s_1 \frac{\Gamma_1 + \Gamma_2}{3}.$$

The coordinate change is polynomial in  $(A, B, \alpha)$  and smooth in  $\lambda$  and  $\mathcal{T}(\alpha) = O(\alpha)$  is linear and upper triangular for each  $\alpha$ , while  $\phi(r)$  satisfies

$$\partial_r \phi(r) = \frac{\sqrt{3}}{2} + O(|\lambda| + |\alpha| + |A|^2), \quad \phi(0) = 0.$$

Note that at  $(\alpha, \lambda) = (0, 0)$ , the trivial state (A, B) = (0, 0) is hyperbolic such that the higher order terms in equation (3.51) are exponentially small for  $\tau \gg 1$  and  $\lambda$  small enough and can be neglected. We can also directly solve the linear part of equation (3.51) to obtain

$$\begin{pmatrix} A(\tau) \\ B(\tau) \end{pmatrix} = \frac{1}{\sqrt{\sinh(\tau)}} \left[ q_1 e^{-\tau\sqrt{c_1\lambda}} \begin{pmatrix} 1 \\ -\sqrt{c_1\lambda} \end{pmatrix} + q_2 e^{\tau\sqrt{c_1\lambda}} \begin{pmatrix} 1 \\ \sqrt{c_1\lambda} \end{pmatrix} \right].$$
(3.52)

We want to find solutions that have a finite energy density with respect to the hyperbolic measure, i.e. functions that are in  $L^2(\mathbb{R}^+, \sinh(\tau)d\tau)$ . This restriction implies that we need to track the stable fast manifold  $\mathcal{W}^{sf}_+(\lambda)$  of equation (3.52) which corresponds to eigenvalues  $\ell$  of  $\mathcal{A}(\infty, \lambda)$  with real part less than  $-\frac{1}{2}$  as shown in Figure 2. Thus, for each fixed  $\tau_0 \gg 1$  and for all sufficiently small  $\lambda < 0$ , we can write the  $\tau = \tau_0$ -fiber of the stable fast manifold  $\mathcal{W}^{sf}_+(\lambda)$  of equation (3.52) near U = 0 as

$$\mathcal{W}_{+}^{sf}(\lambda)|_{\tau=\tau_{0}} : \begin{pmatrix} A\\ B \end{pmatrix} = e^{-\tau_{0}/2} \left[ -\eta \sqrt{c_{1}\lambda} \left(1 + O_{\tau_{0}}(|\lambda|)\right) \begin{pmatrix} \tau_{0}\\ 1 \end{pmatrix} + \sqrt{c_{1}\lambda}\eta \begin{pmatrix} 1\\ 0 \end{pmatrix} \right], \quad (3.53)$$

for  $\eta \in \mathbb{C}$ .

We can now finish the proof of Theorem 2.2. To do this, we need to find nontrivial intersections of the stable fast manifold  $\mathcal{W}^{sf}_+(\lambda)$  with the centre-unstable manifold  $\mathcal{W}^{cu}_-(\lambda)$ . To this end, we write the expansion (3.46) for each fixed  $\tau_0 \gg 1$  in the  $(\tilde{A}, \tilde{B})$  coordinates and afterwards in the coordinates (A, B). Using the expansions of the associated Legendre functions given in Table 1 we arrive at the expression

$$\begin{pmatrix} \tilde{A} \\ \tilde{B} \end{pmatrix} = e^{-\tau_0/2} \left[ e^{i\left(\frac{\sqrt{3}}{2}\tau_0 + \Phi_0\right)} \begin{pmatrix} \frac{C_0}{2}\tilde{d}_1(1+O(1)) + \tau_0\tilde{d}_2\left(-i\frac{\sqrt{3}C_0}{6} + O(1)\right) \\ -\tilde{d}_2(i\frac{\sqrt{3}C_0}{6} + O(1)) + \frac{C_1\sqrt{3}}{6}\left(\Xi_{\tilde{d}_1^2}\tilde{d}_1^2 + \Xi_{\tilde{d}_1\tilde{d}_2}\tilde{d}_1\tilde{d}_2\right) \end{pmatrix} + e^{i\left(\frac{\sqrt{3}}{2}\tau_0 + \Phi_0\right)} \begin{pmatrix} O_{\tau_0}(\lambda|\tilde{d}| + |\tilde{d}|^2) \\ O_{\tau_0}(\lambda|\tilde{d}| + |\tilde{d}_2|^2 + |\tilde{d}_1|^3) \end{pmatrix} \right].$$
(3.54)

We can apply the transformation (3.50) to equation (3.54) and obtain the expansion

$$\mathcal{W}_{-}^{cu}(\lambda) \mid_{\tau=\tau_{0}} : \begin{pmatrix} A \\ B \end{pmatrix} = e^{i(\Phi_{0}+O(\tau_{0}^{-2})+O_{\tau_{0}}(\lambda|\tilde{d}|+|\tilde{d}|^{2})} \begin{pmatrix} O_{\tau_{0}}(\lambda|\tilde{d}|+|\tilde{d}|^{2}) \\ O_{\tau_{0}}(\lambda|\tilde{d}|+|\tilde{d}_{2}|^{2}+|\tilde{d}_{1}|^{3}) \end{pmatrix} + e^{i(\Phi_{0}+O(\tau_{0}^{-2})+O_{\tau_{0}}(\lambda|\tilde{d}|+|\tilde{d}|^{2})} \begin{pmatrix} \frac{C_{0}}{2}\tilde{d}_{1}(1+O(1))+\tau_{0}\tilde{d}_{2}(-i\frac{\sqrt{3}C_{0}}{6}+O(1)) \\ -\tilde{d}_{2}(i\frac{\sqrt{3}C_{0}}{6}+O(1)) \end{pmatrix} + e^{i(\Phi_{0}+O(\tau_{0}^{-2})+O_{\tau_{0}}(\lambda|\tilde{d}|+|\tilde{d}|^{2})} \begin{pmatrix} 0 \\ \frac{C_{1}\sqrt{3}}{6}\left(\Xi_{\tilde{d}_{1}^{2}}\tilde{d}_{1}^{2}+\Xi_{\tilde{d}_{1}\tilde{d}_{2}}\tilde{d}_{1}\tilde{d}_{2}\right) \end{pmatrix}.$$
(3.55)

As in the Euclidean case, the final step of the analysis consists in finding nontrivial intersections of the stable fast manifold  $\mathcal{W}^{sf}_{+}(\lambda)$  given above in equation (3.53) and the core manifold  $\mathcal{W}^{cu}_{-}(\lambda)$  given in (3.55). We can easily solve this problem in  $(\tilde{d}_1, \tilde{d}_2)$  to find that

$$\tilde{d}_1 = -\frac{C_0}{C_1 \Xi_{\tilde{d}_1^2}} \sqrt{3\lambda c_1}, \quad \text{and} \quad \tilde{d}_2 = O(\lambda).$$

This completes the proof of Theorem 2.2.

# 4 Numerical analysis

In this section, we describe the use of numerical continuation (and the continuation package AUTO07P [36]) to compute solutions of the systems of ODEs described by (3.7) in the Euclidean case and (3.39) in the hyperbolic case. Solutions of these spatial dynamical systems correspond to steady states of the neural field equation (2.1) where the radial coordinate r in the Euclidean case or the geodesic polar radial coordinate  $\tau$  in the hyperolic case has been recast as time in AUTO07P's boundary value problem (BVP) solver. The BVP is set up on the domain  $r \in [0, L]$  (or  $\tau \in [0, L]$ ) with homogeneous Neumann boundary conditions given by  $u_2(0) = u_4(0) = u_1(L) = u_3(L) = 0$ . Typical parameters for the AUTO07P radial computations are L = 1000 and AUTO07P's NTST=400 with standard relative tolerances that are specified in AUTO07P's manual. In order to compute initial solutions to start AUTO07P computations and to compute solutions that are not radially symmetric, we solve equation (3.6) in the Euclidean case and (3.38) in the hyperbolic case in polar coordinates where we discretise in the radial variable with finite-differences and a pseudo-spectral Fourier method in the angular variable; see Avitabile et al. [9]. The resulting discretisation is then solved in Matlab using the FSOLVE routine with a Levenberg-Marquardt algorithm, the Jacobian implemented and standard tolerances. For  $D_6$ -symmetric solution branches in the Euclidean case and  $D_8$ -symmetric solution branches in the Hyperbolic case FSOLVE was used in conjunction with standard secant continuation method [68]. For the full 2D computations, we compute on the positive quadrant and take the radial truncation to be L = 60 with 500 finite-difference points and 20 Fourier modes in the angular direction.

#### 4.1 Tuning the parameters

Before numerically studying stationary localized radial solutions on the two geometries, we first reduce the set of parameters governing the shape of the connectivity function and the firing rate function. In both geometries, the connectivity function W, defined in equations (2.3) and (2.19), depends upon the same set of parameters  $(b_1, b_2, \sigma_1, \sigma_2)$  and satisfies equivalent Hypotheses 2.1 and 2.2, respectively. Note that the sigmoidal function, defined in equation (2.2), depends upon two parameters  $(\mu, \theta)$ . Based on the study conducted by Faye *et al.* [44] on the real line, we fix the value of the threshold to  $\theta = 3.5$ .

Following the same lines as in [44], space can be rescaled such that, without loss of generality,  $\sigma_1 = 1$ . The important quantities for the connectivity function are  $\widehat{W}_0, \widehat{W}_c$  (resp.  $\widetilde{W}_0, \widetilde{W}_c$ ), which determine the overall shape of the Hankel transform (resp. Mehler-Fock transform) of W. The connectivity function can then be reparameterized in terms of  $(\widehat{W}_0, \widehat{W}_c)$  (resp.  $(\widetilde{W}_0, \widetilde{W}_c)$ ) solely by the the parameters:

$$\begin{split} \sigma_2 &= \sqrt{\frac{\widehat{W}_c}{\widehat{W}_c - \widehat{W}_0}}, \\ b_1 &= -\frac{2\widehat{W}_c(\widehat{W}_c - \widehat{W}_0)}{\widehat{W}_0}, \\ b_2 &= -\frac{(2\widehat{W}_c - \widehat{W}_0)^2}{2\widehat{W}_0}\sqrt{\frac{\widehat{W}_c}{\widehat{W}_c - \widehat{W}_0}} \end{split}$$

Finally, in order to express the connectivity function in terms of a single parameter, we fix  $\widehat{W}_0 = -1$ . Recall that  $\widehat{W}_0$  has to be negative in order to ensure the existence of a unique trivial solution v = 0 of equation (2.1). Hence, the connectivity function only depends upon  $\widehat{W}_c$  (resp.  $\widetilde{W}_c$ ).

### 4.2 Numerical computation of the stability of localized solutions

In the Euclidean case, in order to numerically investigate the radial stability of a stationary localized bump solution  $V_0(r)$  of equation (2.1), we have to solve the following eigenvalue problem. We linearize equation (2.1) around  $V_0(r)$  to find the linearized equation

$$\frac{d}{dt}v(\mathbf{r},t) = -v(\mathbf{r},t) + \mu \int_{\mathbb{R}^2} W(\|\mathbf{r} - \mathbf{r}'\|) S'(\mu V_0(\|\mathbf{r}'\|)) v(\mathbf{r}',t) d\mathbf{r}'.$$
(4.1)

Looking for radial perturbation of the form  $v(\mathbf{r}, t) = e^{\sigma t}q(r)$  and applying the PDE method developed in §3.1.2 to equation (4.1), we obtain:

$$(\sigma+1)\mathcal{L}_0(q(r)) = \mu\left(\Gamma_1 - \Gamma_2\Delta_r\right)\left[S'(\mu V_0(r)q(r))\right].$$

$$(4.2)$$

The linear operator  $\mathcal{L}_0$  is given by  $\mathcal{L}_0 = \Delta_r^2 - (\sigma_1^2 + \sigma_2^2)\Delta_r + \sigma_1^2\sigma_2^2$ . For all  $q \in \mathcal{H}^4(\mathbb{R}^+_*) = \{u \in L^2(\mathbb{R}^+_*) \mid \forall k \leq 4 \ \partial_{r^k}^k u(r) \in L^2(\mathbb{R}^+_*)\}$  and  $p \in L^2(\mathbb{R}^+_*)$ , we set  $\mathcal{L}_0 q = p$ . As the spectrum of  $\mathcal{L}_0$  is given by  $\operatorname{spec}_{\mathcal{L}_0} = \{\rho^4 + (\sigma_1^2 + \sigma_2^2)\rho^2 + \sigma_1^2\sigma_2^2 \mid \rho \in \mathbb{R}\} \subset [\sigma_1^2\sigma_2^2, +\infty[, \mathcal{L}_0 \text{ is an invertible operator and } q = \mathcal{L}_0^{-1}p$ . It follows that equation (4.2) can be rewritten:

$$\sigma p(r) = -p(r) + \mu \left(\Gamma_1 - \Gamma_2 \Delta_r\right) \left[ S'(\mu V_0(r)) \mathcal{L}_0^{-1} p(r) \right] = \mathcal{B}(p(r)).$$

$$(4.3)$$

Now, for every solution  $V_0$  discretized on a domain [0, L], we compute the eigenvalues  $\sigma$  of the corresponding discretized version of the linear operator  $\mathcal{B}$  where we use finite differences methods to approximate the Laplacian operator  $\Delta_r$ . The results in [87] imply that spectral stability also implies linear (asymptotic) stability. Hence, when all the eigenvalues of  $\mathcal{B}$  have negative real part, we say the solution is radially stable and unstable otherwise.

The radial stability in the hyperbolic case is computed in a similar fashion. If  $V_0(\tau)$  is a stationary localized bump solution of equation (2.11) then the linearized equation around  $V_0(\tau)$  is:

$$\frac{d}{dt}v(z,t) = -v(z,t) + \mu \int_{\mathbb{D}} W(d_{\mathbb{D}}(z,z'))S'(\mu V_0(d_{\mathbb{D}}(z',0)))v(z',t)\mathrm{dm}(z').$$
(4.4)



Figure 3: Radial solutions in the Euclidean case. (a) Bifurcation diagram in  $\mu$  where stable solution branch segments are black and unstable segments are gray. The trivial solution branch at  $||u||_{L^2} = 0$  is dashed and the two branches bifurcating at the Turing instability K are solid. The first branch undergoes three fold bifurcations  $L_1$ ,  $L_2$  and  $L_3$  and is stable between  $L_1$  and  $L_2$ ; the second branch undergoes a fold at  $L_4$ ; both branches asymptote at  $\mu = 3.94$ . The inset shows the solution profile on the asymptote. (b) Solution profile at  $\mu = 6$  (vertical dashed line) on the branch segment between  $L_2$  and  $L_3$ . (c) Solution profile at  $\mu = 6$  on the branch segment between  $L_1$  and  $L_2$ . (d) Solution profile at  $\mu = 6$  on the branch segment between Kand  $L_4$ . (d) Solution profile at  $\mu = 6$  on the branch segment between K and  $L_1$ .

Looking for radial perturbation of the form  $v(z,t) = e^{\sigma t}q(\tau)$  and applying the PDE method developed in §3.2.2 to equation (4.4), we obtain:

$$(\sigma+1)\mathcal{L}_{0}^{\mathbb{D}}\left(q\left(\tau\right)\right) = \mu\left(\Gamma_{1} - \Gamma_{2}L_{\mathbb{D}}^{\tau}\right)\left[S'(\mu V_{0}\left(\tau\right)q\left(\tau\right)\right].$$
(4.5)

 $L_{\mathbb{D}}^{\tau}$  is defined in equation (2.16) and the linear operator  $\mathcal{L}_{0}^{\mathbb{D}}$  is given by  $\mathcal{L}_{0} = (L_{\mathbb{D}}^{\tau})^{2} - (\sigma_{1}^{2} + \sigma_{2}^{2})L_{\mathbb{D}}^{\tau} + \sigma_{1}^{2}\sigma_{2}^{2}$  is aslo invertible, as its spectrum is positive. If we denote  $q = (\mathcal{L}_{0}^{\mathbb{D}})^{-1} p$ . It follows that equation (4.5) can be rewritten:

$$\sigma p(\tau) = -p(\tau) + \mu \left(\Gamma_1 - \Gamma_2 L_{\mathbb{D}}^{\tau}\right) \left[S'(\mu V_0(\tau)) \left(\mathcal{L}_0^{\mathbb{D}}\right)^{-1} p(\tau)\right] = \mathcal{B}^{\mathbb{D}}(p(\tau)).$$
(4.6)

Again, for every solution  $V_0$  discretized on a domain [0, L], we compute the eigenvalues  $\sigma$  of the corresponding discretized version of the linear operator  $\mathcal{B}^{\mathbb{D}}$  where we use finite differences methods to approximate the Laplace-Beltrami operator  $L^{\tau}_{\mathbb{D}}$ . When all eigenvalues have negative real part, then the solution is said to be radially stable, otherwise it is unstable.

We note that stability with respect to other perturbations (in particular  $\mathbf{D}_n$ -perturbations) can similarly be investigated by carrying out a polar discretisation of the differentiation operators in  $\mathcal{B}$  and  $\mathcal{B}^{\mathbb{D}}$  as described above.



Figure 4: Two-parameter solution structure in the Euclidean case. (a) Two parameters bifurcation diagram in the  $(\mu, \widehat{W}_c)$ -plane. The locus  $\widehat{K}$  of the bifurcation point K shown in Figure 3(a) is gray; the loci  $\widehat{L}_1$ and  $\widehat{L}_2$  of the respective fold points  $L_1$  and  $L_2$  are black. A stable spot solution, as shown in Figure 3(c), persists in the region between  $\widehat{L}_1$  and  $\widehat{L}_2$ , which is shaded in gray. Dashed lines are the one dimensional slices corresponding to the one-parameter diagrams shown in Figure 3(a) and panels (b) and (c). (b) One parameter diagram in  $\widehat{W}_c$  at  $\mu = 6$ . (c) One parameter diagram in  $\mu$  at  $\widehat{W}_c = 6$ .

### 4.3 Euclidean case

### 4.3.1 Radially symmetric solution branches

We present branches of radial solutions on the Euclidean plane for the reduced system of ODEs given in (3.7). In this section, we fix the critical Fourier mode  $\widehat{W}_c = 3.25$  that fixes the critical value of the sigmoidal slope to  $\mu_c = 10.81$  (given by (3.3)) for which there is a Turing instability for the trivial state  $u_1 = 0$ . Figure 3(a) shows a bifurcation diagram value  $\mu$  where the solution branches are represented in terms of the L<sup>2</sup>-norm  $||u||_{L^2} = (\int_0^\infty u(r)^2 r dr)^{1/2}$ . The inset 3(a) and panels 3(b)-3(e) show solution profiles at different points on the bifurcation diagram as indicated. Two branches of spatially localized solutions bifurcate from the trivial state at the bifurcation point K where  $\mu = \mu_c$ . The first branch represents a bump solution, as shown in panel (e) for  $\mu = 6$ , on the segment between K and  $L_1$  and this solution is initially unstable. The branch becomes stable and the bump solution becomes a spot solution, as shown in panel (c) for  $\mu = 6$ , at the fold bifurcation  $L_1$ . The branch loses stability and a ring is added to the spot, as shown in panel (b) for  $\mu = 6$ , at the fold point  $L_2$ . After a further fold point  $L_3$  the ring starts to travel away from the core, as shown in the inset of (a), and the solution branch asymptotes at  $\mu = 3.94$  in a gluing bifurcation. Such glueing bifurcations involving radial spots have been observed in other reaction-diffusion systems [10] (though it is an open problem to explain why the bifurcation is co-dimension one as the pulse in the  $r = \infty$ invariant subspace is co-dimension zero). The second branch represents an unstable bump solution with a ring and in a similar fashion the branch asymptotes in a gluing bifurcation where the ring travels away from the core.

So far, we have identified a particular range of  $\mu$  for which localized radial solutions exist, for specific values of the critical Fourier mode  $\widehat{W}_c = 3.25$ . We now show that localized bump solutions persist over a range of the parameters  $(\mu, \widehat{W}_c)$  and are not an isolated phenomena in parameter space. First we observe that, as shown in Figure 3(a), the branch of localized bump solutions is stable with respect to radial perturbations between the two fold points  $L_1$  and  $L_2$  for all the associated  $\mu$ -values as as indicated by the black curve. Therefore, assuming that this is also the case when  $\widehat{W}_c$  is varied, we can find bounds of regions for which localized bump solutions exist; in the  $(\mu, \widehat{W}_c)$ -plane it is sufficient to track the loci of the fold points  $L_1$  and  $L_2$  under the simultaneous variation of those two parameters. We denote  $\widehat{K}$  the locus of the bifurcation point K shown in Figure 3(a) and we denote  $\widehat{L}_1$  and  $\widehat{L}_2$  the loci of the respective fold points  $L_1$  and  $L_2$ .

Figure 4(a) shows curves that are the loci of bifurcations in the  $(\mu, \widehat{W}_c)$  parameter plane. The curve  $\widehat{K}$  is the locus of the Turing bifurcation at  $\mu_c$ , which is determined analytically by the expression  $\mu_c = 1/(s_1 \widehat{W}_c)$ .



Figure 5:  $\mathbf{D}_6$ -symmetric solutions in the Euclidean case. (a) Zoom of Figure 3(a) close to  $L_2$  with a branch segment of  $\mathbf{D}_6$ -symmetric solutions plotted. The  $\mathbf{D}_6$ -branch bifurcates from the radially stable branch at  $P^{\mathbf{D}_6}$ , close to  $L_2$ ; the  $\mathbf{D}_6$ -branch undergoes a fold bifurcation at  $L^{\mathbf{D}_6}$  and further fold bifurcations at a  $\mu$ value slightly smaller than that of  $L_2$ . Further branch segments that are not shown are indicated by arrows. (b)–(e) Solution profiles from different points on the  $\mathbf{D}_6$ -branch plotted in the euclidean plane.

The curves  $\widehat{L}_1$  and  $\widehat{L}_2$  bound the region for which localized bump solutions persist (shaded in gray). Under the variation of  $\mu$  and  $\widehat{W}_c$ , shown in Figure 4(a), there is a channel in parameter space for which localized solutions exist. The vertical line in Figure 4(a) corresponds to the default value  $\mu = 6$  and the corresponding bifurcation diagram as  $\widehat{W}_c$  is varied is shown in Figure 4(b). We note that this bifurcation diagram has the same characteristics as the one obtained when varying  $\mu$  (see Figure 3(a)) with  $\widehat{W}_c = 3.25$ . We also remark that in both cases the localized bump solutions always exist before the Turing bifurcation. Therefore, the localized bump solutions coexist with a stable trivial solution. One horizontal line in Figure 4(a) corresponds to the default value of  $\widehat{W}_c$  used in Figure 3 and the other one at  $\widehat{W}_c = 6$  is shown in Figure 4(c). The bifurcation diagram in Figure 4(c) shows that the folds  $L_3$  and  $L_4$  have undergone a saddle transition creating two separate branches, one connected back to K and one that asymptotes at  $\mu = 2.10$ . Similar saddle transition scenario has been encountered in [9] for branches of fully localized stripe patterns (worms) in the planar Swift-Hohenberg equation with cubic/quintic nonlinearity; see [9].

#### 4.3.2 D<sub>6</sub>-symmetric solution branch

In order to detect possible bifurcations from the radial solution branch to non-radially symmetric solutions, the radial ODE system is augmented with the linearized system with respect to  $\mathbf{D}_6$  perturbations; see [75]. We concentrate on  $\mathbf{D}_6$ -perturbations since hexagons are usually the most stable patterns in Euclidean geometry due to close-packing arguments. Figure 5 shows detail of Figure 3(a) near the fold point  $L_1$  and a possible bifurcation to a  $\mathbf{D}_6$ -symmetric branch was detected in AUTO07P at the point  $P^{\mathbf{D}_6}$  (note that when traced in the  $(\mu, \widehat{W}_c)$ -plane, the point  $P^{\mathbf{D}_6}$  was found to follow  $L_1$  very closely). In order to find a solution on the  $\mathbf{D}_6$ -branch it was necessary to reconverge the solution from the bifurcation point in Matlab and confirm that there is locally a zero-crossing of an eigenvalue  $\sigma$  associated with an eigenfunction q of (4.2) that is  $\mathbf{D}_6$ -symmetric. Note that  $\sigma$  and q are computed numerically from the Jacobian at the equilibrium point. Using the eigenfunction as a perturbation at a parameter value of  $\mu$  less than at  $P^{\mathbf{D}_6}$  we are able to converge



Figure 6: Decay rates of radial solutions in hyperbolic case. In panels (a) and (b) respectively, the solutions  $u(\tau)$  at  $\mu = 7$  and  $\mu = 2$  are black and the same solutions rescaled as  $u(\tau)e^{\tau/2}$  are gray. Panels (c) and (d) show the decay rate of the tails of the solutions in black at  $\mu = 7$  and  $\mu = 2$ , respectively. Linear best fits are shown in gray with the linear coefficients  $p_1$  as indicated.

a solution on the  $D_6$ -symmetric branch. A standard secant continuation method is then used to compute the new branch.

Figure 5 shows the branch of  $\mathbf{D}_6$ -symmetric solutions that bifurcates off the original radial branch at  $P^{\mathbf{D}_6}$ . The  $\mathbf{D}_6$ -branch undergoes a fold bifurcation at  $L^{\mathbf{D}_6}$  and further fold bifurcations at a  $\mu$ -value slightly smaller than that of  $L_2$ . We see that as  $\mu$  decreases, an inner ring of 12 cells and a outer ring of 6 cells with small amplitude is glued to the initial bump solution. The amplitude of these cells is largest as  $L^{\mathbf{D}_6}$  is approached. For the branch shown the numerical accuracy was tested by recomputing it with double the number of angular mesh points. The branches carry on as indicated by the arrows in Figure 5 but they undergo a complicated sequence of folds that we leave for further study.

### 4.4 Hyperbolic radial case

Before tracing out the bifurcation diagrams for the radial spots, we verify that the spots found in Theorem 2.2 are  $L^2(\mathbb{R}^+, \sinh(\tau)d\tau)$  functions. In Figure 6, we show the radial states and the decay rate of their tails. Figure 6(a) and (b) show radial spot solutions  $u(\tau)$  close to  $\mu_c$  at  $\mu = 7$  and on a connected radially stable branch segment at  $\mu = 2$ , respectively. In each case we also show the same solutions rescaled by  $e^{\tau/2}$ ; these rescaled solutions still decay, which shows that in each case the decay rate is less than  $e^{-\tau/2}$  as predicted in §3.2.4. Close to onset at  $\mu = 7$ , Figure 6 shows that the bifurcating pulse has highly oscillatory tails as also predicted by the analysis. Plotting the decay rate of the tails of the radial states, Figure 6(c) and (d) confirm that for  $\mu = 7$  and  $\mu = 2$ , both states decay faster than  $e^{-\tau/2}$  and hence are in  $L^2(\mathbb{R}^+, \sinh(\tau)d\tau)$ . In particular, we see in Figure 6(c), that close to onset the decay rate converges to a half as predicted.

We now present branches of radial solutions on the Poincaré disk for the reduced system of ODEs given in (3.39). In Figure 7(a), we fix the critical Fourier mode  $\widetilde{W}_c = 5$  that fixes the critical value of the sigmoidal slope to  $\mu_c = 7.03$  (given by (3.35)). Figure 7(a) shows a bifurcation diagram in  $\mu$  where the solution



Figure 7: Radial solutions in the hyperbolic case. (a) One-parameter bifurcation diagram in  $\mu$  at  $W_c = 5$ . A single solution branch originating from K (not shown) has a stable segment between  $L_1$  and  $L_2$  and asymptotes at  $\mu = 1.95$ . (c) Solution profile on the asymptote. (d) Solution profile at  $\mu = 2$  on the branch segment between  $L_2$  and the asymptote. (e) Solution profile at  $\mu = 2$  on the branch segment between  $L_1$ and  $L_2$ . (f) Solution profile at  $\mu = 2$  on the branch segment between K and  $L_1$ .

branches are represented in terms of the Euclidean  $L^2$ -norm  $||u||_{L^2} = (\int_0^\infty u(\tau)^2 \tau d\tau)^{1/2}$ . This is due to numerical difficulties of computing the hyperbolic  $L^2$  norm presented by the sinh-function in the integrand but we believe the Euclidean radial norm to be a good solution measure. The panels (c)–(f) show solution profiles at different points on the bifurcation diagram as indicated. The branch represents a bump solution, as shown in panel (f) for  $\mu = 2$ , on the segment between K and  $L_1$  and this solution is initially unstable (K is only shown in figure 7(b)). The branch becomes stable and the bump solution becomes a spot solution at the fold bifurcation  $L_1$ ; see panel (e) for  $\mu = 2$ . The branch loses stability and the spot decays at the core at the fold point  $L_2$ ; see panel (d) for  $\mu = 2$ . After a further fold point  $L_3$  the spot starts to travel away from the core, as shown in the inset of (c), and the solution branch asymptotes at  $\mu = 1.95$ .

We denote  $\widehat{K}$  the locus of the bifurcation point K and we denote  $\widehat{L}_1$  and  $\widehat{L}_2$  the loci of the respective fold points  $L_1$  and  $L_2$ . Figure 7(b) show curves that are the loci of bifurcations in the  $(\mu, \widetilde{W}_c)$  parameter plane. The curve  $\widehat{K}$  is the locus of the Turing bifurcation at  $\mu_c$ , which is determined analytically by the expression  $\mu_c = 1/(s_1\widetilde{W}_c)$ . The curves  $\widehat{L}_1$  and  $\widehat{L}_2$  bound the region for which localized bump solutions persist (shaded in gray). Under the variation of  $\mu$  and  $\widetilde{W}_c$ , shown in Figure 7(b), there is a channel in parameter space for which localized solutions exist. The horizontal line in Figure 7(b) corresponds to the default value of  $\widetilde{W}_c$  used in Figure 7(a). We can see that the region for which localized bump solutions persist is much thinner in the hyperbolic case than in the Euclidean case. This implies that radial stationary spots are more likely to be seen in the Euclidean plane than in the Poincaré disk. Furthermore, the bifurcation where the asymptote occurs, is fundamentally different. In the Euclidean case, the asymptote is due to a glueing bifurcation between a bump solution at the core and a symmetric pulse in the  $r = \infty$  invariant subspace. In the Hyperbolic case, the bifurcation is due to the fact that there is co-dimension one homoclinic bifurcation where an asymmetric pulse occurs in the  $\tau = \infty$ -invariant subspace. The reason why the bifurcation in the hyperbolic case is co-dimension one is due to the fact that in the  $\tau = \infty$  invariant subspace, the stationary equations no longer have spatial reflectional symmetry, hence an odd pulse; since the ODE system is not



Figure 8:  $\mathbf{D}_8$ -symmetric solutions in the Hyperbolic case. (a) Zoom of Figure 7(a) where radial solution branches are now dashed gray. Two new branches of  $\mathbf{D}_8$ -symmetric solutions bifurcate from the radial branch between  $L_1$  and  $L_2$  at  $P^{\mathbf{D}8}$ . The upper segment is gray and undergoes a fold bifurcation  $L_1^{\mathbf{D}8}$  followed by a series of unlabelled fold bifurcations. The lower segment is black and undergoes a single fold bifurcation  $L_2^{\mathbf{D}8}$ . Panels (b)–(e) show solution profiles plotted polar geodesic coordinates (left) and in the hyperbolic disk (right). Panels (b) and (e) are from the lower  $\mathbf{D}_8$ -branch at points above and below  $L_2^{\mathbf{D}8}$ , respectively. Panels (c) and (d) are from upper branch at points above and below  $L_1^{\mathbf{D}8}$ , respectively.

conservative such pulses are generically co-dimension one.

#### 4.4.1 D<sub>8</sub>-symmetric solution branch

In [27, 45], Chossat and Faye studied the bifurcations of periodic solutions in the Poincaré disk with octagonal symmetry for neural field equations similar to (2.11). They were able to, for each absolutely irreducible representation, classify all the corresponding isotropy subgroups satisfying the Equivariant Branching Lemma [28, 51];  $\mathbf{D}_8$  appears naturally as such an isotropy subgroup [27]. Hence, it is natural to look for  $\mathbf{D}_8$ -symmetric localized solutions. Note, we have found that the radial spot undergoes other dihedral instabilities (in particular  $\mathbf{D}_4$  and  $\mathbf{D}_6$  instabilities) but we concentrate on the  $\mathbf{D}_8$  instability.

With a similar approach as was taken in §4.3.2, we find a branch of  $\mathbf{D}_8$ -symmetric solutions bifurcating from the main, stable, radial branch that was shown in Figure 7(a). Figure 8(a) shows a zoom of Figure 7 with a  $\mathbf{D}_8$ -symmetric branch of solutions that bifurcates from the radial branch at  $P^{\mathbf{D}_8}$ . The bifurcated branch exists for  $\mu$ -values below  $P^{\mathbf{D}_8}$  and has an upper and a lower segment. The lower segment undergoes a single fold bifurcation at  $L_2^{\mathbf{D}_8}$  before reconnecting back to the trivial solution branch at the Turing instability K(not shown in Figure 8(a)). The upper branch undergoes a series of fold bifurcations starting with  $L_1^{\mathbf{D}_8}$ and also connects back to K. The solution profile on the lower branch close to  $P^{\mathbf{D}_8}$  in panel (b) shows a  $\mathbf{D}_8$ -symmetric spot solution. After the fold bifurcation  $L_2^{\mathbf{D}_8}$ , one can see that the solution has formed a spot solution surrounded by a ring of 8 smaller cells; see panel (e). On the upper branch, there is also a  $\mathbf{D}_8$ -symmetric spot solution as shown in panel (c). After the fold bifurcation  $L_1^{\mathbf{D}_8}$  we see that the solution has 4 out of 8 nodes at an elevated activity level. For the  $\mathbf{D}_6$  solution branches discussed in §4.3.2, we found that at the bifurcation from the radial branch, a ring of cells was glued to the main radially symmetric solution. Here, in contrast, we find that a modified spot solution that is  $\mathbf{D}_8$ -symmetric; although a ring of 8 cells can also form about the spot as was shown in Figure 8(e).

# 5 Conclusion

In this paper, we investigated radially localized solutions of a neural field equation set on the Euclidean plane and on the Poincaré disk equipped with their corresponding natural metrics. Applying an appropriate integral transform in both geometries we were able to reduce the neural field equations to partial differential equations and look for stationary radially symmetric solutions. We applied the techniques from Scheel [88] and Lloyd & Sandstede [74] to prove the existence of a branch of bump solutions bifurcating near onset for both Euclidean and the hyperbolic geometries. In the Poincaré disk, the analysis near the core manifold required significant changes compared with the approach used for the Euclidean case; in hyperbolic geometry it was necessary to developed a detailed knowledge of the asymptotics of the associated Legendre functions. It turns out that the main difference between the two geometries comes from the far field. At infinity, Bessel function  $J_0(r)$  scales in term of the radial coordinate r as  $1/\sqrt{r}$  whereas the associated Legendre function  $\mathcal{P}_{\nu}(\cosh(\tau))$  scales in term of the polar geodesic coordinate  $\tau$  as  $e^{-\tau/2}$  with  $\nu = -\frac{1}{2} + i\frac{\sqrt{3}}{2}$ . Moreover, in the Euclidean case, the trivial state at infinity, there is a Turing instability bifurcation. However, in the hyperbolic case, the trivial state is always asymptotically stable at infinity and this simplifies the resulting matching problem. Other localized radial states are also expected and the theory developed in [77, 79] can be applied to yield the existence of rings and another spot solution not investigated in this paper.

Numerical continuation was used to track branches of spot solutions away from onset and we implemented a method to compute their radial stability. We identified regions of parameter space where the spot solutions are radially stable for both geometries. We also found branches of localized  $\mathbf{D}_6$  symmetric and  $\mathbf{D}_8$ -symmetric solutions bifurcating off the stable radial branches in the Euclidean and hyperbolic geometries, respectively. In terms of the radial spot solutions, the region of radial stability in the Euclidean case is significantly larger than in the hyperbolic case. Although the region of stability is smaller in the hyperbolic case, we expect to find other types of solution by looking at different radial perturbations, for example, time-periodic localized states; see [26]. In [44], the 1D analogue of the present model was studied and stationary localized structures were shown to undergo homoclinic snaking, however, this does not occur for both of the 2D geometries studied here. This appears to be due to the connectivity function creating almost spike-like multi-pulse localized states similar to that seen in Lloyd & O'Farrell [73] for a different PDE system and we believe this is due to the presence of a singular limit nearby in parameter space. For the connectivity function chosen in the Laing & Troy model [69], snaking-like behavior was found on the Euclidean plane for radial localized solutions, we expect that the localized hexagon patches that were found in their study could also undergo homoclinic snaking similar to that observed in the planar Swift-Hohenberg equation [75].

For the neural field model set on the Euclidean plane, radially localized solutions can be interpreted as spatially localized regions of high activity on the cortex. In particular, this type of solutions is of interest in modeling working memory: the ability to remember information over a time-scale of a few seconds. Indeed, some experiments on primates [29, 49, 80] reveal that there are spatially localized regions of neurons with elevated firing rates during the period that the animal is remembering some aspect of an object or event. In this paper, we have been able to prove the existence of such solutions. More precisely, we have demonstrated that there exists radially localized solutions which bifurcate off the trivial state using dynamical system's methods developed by Scheel [88] and Lloyd & Sandstede [74]. This result provides a better understanding of the formation of this type of solutions in planar neural field equations (though near onset such states are unstable) and is complementary to the constructive approach used in [48, 91, 96]. We also identified a region in the parameter plane (one parameter for the connectivity function and one parameter for the slope of the firing rate function) where these solutions are stable with respect to radial perturbations and certain periodic perturbations, which generalizes the results presented in [69, 82].

Our results show that a cortical network with local excitatory and inhibitory connections can spontaneously

produce regions of localized high activity. Another interpretation of these regions of localized activity is related to the recent work of Chavane *et al.* [25] where it was found, using voltage sensitive dye imaging in the cat primary visual cortex, that localized oriented inputs in the visual field produce localized regions of activity in the cortical feedforward imprint of the input. Suppose now that we add to equation (2.1) an external input of the form  $\epsilon I_{\text{ext}}(||\mathbf{r}||)$  with  $0 < \epsilon \ll 1$  where  $I_{\text{ext}}$  is radially localized. This breaks the translational symmetry but preserves the reflectional and rotational symmetry. Then, for small enough  $\epsilon$ , we still expect to find radially localized solution with a high region of activity for  $(\mu, \widehat{W}_c)$  in the grey region of Figure 4(a). This type of result is well known for the ring model of orientation [14, 55, 93] and for the one dimensional version of equation (2.1) [44]. A more detailed study of the effects of localized external inputs in equation (2.1) will be the subject of future work. We have also demonstrated in the Euclidean case the existence of radially localized solutions with a ring of glued  $\mathbf{D}_6$ -symmetric patches that is another possible response to localized inputs. Indeed, such patchy localized responses have been experimentally recorded in the primary visual cortex [3, 76].

The neural field equation (2.11) set on the Poincaré disk has been used to model orientation and texture selectivity for a hypercolumn of the primary visual cortex [26, 42, 46] where the angle of a point in the Poincaré disk represents the preferred orientation and its modulus the corresponding selectivity. In this model, radially localized solutions can be interpreted as the analog of the tuning curves found for the ring model of orientation [14, 55, 93] or the analog of the tuning surfaces of the spherical model developed by Bressloff and Cowan [16, 19]. We have proved that the neural field model set on the Poincaré disk is able to spontaneously produce radially localized solutions i.e. tuning surfaces. Then, due to the equivariance of equation (2.11) with respect to hyperbolic translations [26, 46], there exists a family of radially localized solutions which are translated copies of the solution given by Theorem 2.2. Each translated copy is a tuning surface with a high activity region centered at some given point in the Poincaré disk. As for the Euclidean case, we found that localized solutions stable to radial perturbations exist in a region of the parameter plane  $(\mu, \widetilde{W}_c)$  (the grey region of Figure 7(b)); in the hyperbolic case, the region is significantly smaller. The next step of the analysis would be to introduce into equation (2.11) an external input, weakly tuned to a specific preferred texture, and study if our network is able to produce a tuned response.

A natural extension of this work would be to study the effects of spike frequency adaptation [47, 48, 71] and synaptic depression [22, 61, 62] on spatiotemporal dynamics of our neural field models, where we expect to find spirals, breathers and traveling waves; we leave this for further work.

Finally, we note that spectral analysis on the real hyperbolic space  $\mathbb{H}^n$ ,  $n \geq 2$ , has received much interest in the areas of quantum chaos [8, 11, 33, 89] and cosmology [59, 72]. More recently, some aspects of dispersive and concentration phenomena have been studied for evolution equations such as the nonlinear Schrödinger equation and the wave equation posed on  $\mathbb{H}^n$  [4, 5, 12]. Our study of radial states in (1.1) in this paper constitutes, from a mathematical point of view, a framework to highlight the inherent differences and similarities between the Euclidean and the hyperbolic geometry that may be of interest in other areas of physics.

#### Acknowledgments

GF and JR were partially funded by the ERC advanced grant NerVi number 227747 and by the European Union Seventh Framework Program (FP7/2007- 2013) under grant agreement no. 269921 (BrainScaleS). DL acknowledges support from an EPSRC grant EP/H05040X/1.

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