

A second order all Mach number IMEX finite volume solver for the three dimensional Euler equations

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Abstract

This article deals with the development of a numerical method for the compressible Euler system valid for all Mach numbers: from extremely low to high regimes. In classical fluid dynamic problems, one faces both situations in which the flow is subsonic, and consequently acoustic waves are very fast compared to the velocity of the fluid, and situations in which the fluid moves at high speed and compressibility may generate shock waves. Standard explicit fluid solvers such as Godunov method fail in the description of both flows due to time step restrictions caused by the stiffness of the equations which leads to prohibitive computational costs. In this work, we develop a new method for the full Euler system of gas dynamics based on partitioning the equations into a fast and a low scale. Such a method employs a time step which is independent of the speed of the pressure waves and works uniformly for all Mach numbers. Cell centered discretization on Cartesian meshes is proposed. Numerical results up to the three dimensional case show the accuracy, the robustness and the effectiveness of the proposed approach.

Keywords: All Mach number flow solver, Asymptotic preserving, Implicit-Explicit Runge-Kutta schemes, Incompressible flows, Multidimensional Euler equations.

1. Introduction

The equations which model high speed and low speed inviscid flows are well known and have been studied for more than a century [61]. These models are respectively referred to as the compressible Euler equations and the incompressible ones. However, there are situations in which the fluid may be only slightly compressible or situations in which the fluid exhibits very fast pressure waves in some regions of the domain and not in others, or even fast waves are present only for short times compared to the scale of observation. In all these cases, it is interesting to analyze from a theoretical point of view what happens if one performs the passage from incompressible to compressible scaling and vice versa [56, 57, 81, 2, 62, 1]. It is probably equally interesting from the point of view of the applications to dispose of numerical methods which are able to deal both with the incompressible and the compressible regimes [50, 44, 74, 91, 58, 20, 42, 69, 66, 45, 41, 73, 67, 24, 27, 25, 23, 43, 16, 18, 39, 55, 28, 30]. The compressible Euler equations describe the conservation of density, momentum and energy, while the corresponding incompressible model deals with a constraint on the divergence of the velocity, with an elliptic equation for the pressure and a linear transport equation for the density moving at the average speed of the flow. These two models can be connected by a rescaling procedure in which the Mach number, i.e. the ratio between the flow and the sound speeds, is put into evidence in the compressible model and then is set to zero through a limit analysis. This situation implies a fluid flow almost at rest, and, the pressure waves to move at large speed compared to the average speed of the gas.

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Unfortunately, the aforementioned models cannot deal with all occurrences encountered in practical cases. In fact, in many problems, the presence of low and high Mach number flows becomes evident during the study of the phenomena and cannot be predicted before starting the simulation. Thus, one needs to be able to describe these situations with the same numerical scheme since switching the scheme during the time marching is not a strategy that can be followed. However, this causes many difficulties from the numerical point of view because standard numerical methods suffer from the stiffness of the equations, while domain decomposition techniques which may couple the two models are difficult to be set into practice, as pointed out for instance in [7]. Since the problem involves a fast time scale, one possible solution is given by the use of implicit time integrators. On the other hand, it is well known that implicit schemes are not simple to be applied due to the non linearity of the system. Furthermore, the necessity to employ high order in space schemes complicates the realization of such methods. In the recent past, an alternative strategy to treat problems with multiple time scales has been proposed, which consists in treating implicitly only one part of the system to be solved while keeping the remaining explicit. This approach considerably facilitates the development of the schemes since it diminishes the complexity of the non linear systems to be inverted in order to advance the solution in time. We refer to this class of schemes as Implicit-Explicit methods [3, 8, 72], that are a particular case of the so-called partitioned schemes [48]. Thanks to this choice, it is possible to develop numerical schemes which are stable for all values of the fast scale, independently on the prescribed value of the time step. Moreover, if these schemes share also the property of being consistent with the limit model, they are often referred to as Asymptotic Preserving methods. In the context of compressible-incompressible fluids, these techniques have been developed by several authors in the last decade [26, 24, 25, 23, 43, 85, 16, 68, 18, 30, 29]. A semi-implicit finite volume scheme introduced in [35], which shares similarities with the one here proposed, makes use of a staggered grid spatial discretization. Staggered meshes in the context of semi-implicit schemes for incompressible fluids have also been considered in [86, 88, 11, 12]. In two recent works [30, 29], first and second order Asymptotic Preserving methods have been developed for the low Mach number limit for the simplified isentropic Euler case. In [94] a semi-implicit scheme for atmospheric flows is proposed where the separation of acoustic and advective components of the hyperbolic flux is carried out in the characteristic space, while in [95] a correction to deal with non-hyperbolic flux splitting is presented. High order discontinuous Galerkin methods on staggered grids for compressible flows have been presented in [87].

In the present work, we extend the previous study to the full Euler equations with the intent of simulating up to three dimensional flows. Compared to the isentropic Euler situation, here the difficulties are related to search for the minimum number of terms to treat implicitly in the full Euler system, and, to deal with the non linearity of the system which arises in the scheme. All these aspects increase in complexity by the fact that multidimensional problems are solved. In the following, we first present a first order in time and space method and then we extend our analysis to general second order in time Implicit-Explicit (IMEX) methods coupled with a second order space discretization approach. The time marching technique is based on IMEX Runge Kutta method for the time and TVD second order discretization for the space [60]. In [29], we focused on the analysis of the development of L^2 and L^∞ stable methods and on the differences that arise if one stability condition is assured instead of the other. In this work, we only consider the weaker L^2 stability requirement for sake of shortness and simplicity. We postpone the development of schemes which are also able to assure the L^∞ stability condition to future investigations. We only recall that it has been proved [40] that the L^∞ and the Total Variation Diminishing (TVD), also named Strong Stability Preserving (SSP), properties cannot be assured for an unconstrained implicit in time scheme of order greater than one, and, thus, remedy to this negative result should be studied. In [29], a solution constructing high resolution algorithms by combining schemes of different nature has been proposed. We intend to extend and study in details this possibility for the full Euler case in some next future work.

The article is organized as follows. In Section 2, we recall the governing equations, the low Mach scaling and we give the details of the formal passage from the compressible to the incompressible model. In Section 3, we introduce a first order scheme which works for all values of the Mach number. We start by discussing the semi discrete in time case in 3.1, then we give the details of the full time and space discretization for the one dimensional case in Section 3.2 and we end with the three dimensional first order case in Section 3.4. The second order extension in time and space of our method is proposed in Section 4, namely the time semi discretization is discussed in Section 4.1 while the space second order extension in Section 4.2. The numerical results showing the effectiveness of the proposed approach are given in Section 5. A concluding Section finalizes the article.

2. Governing equations

We consider a bounded domain $\Omega \in \mathbb{R}^d$ with the space dimension $d \in \{1, 2, 3\}$. The space and time variables are respectively denoted by $\mathbf{x} \in \Omega$ and $t \in \mathbb{R}_+$. The compressible Euler equations read

$$\partial_t \rho + \nabla_{\mathbf{x}} \cdot \mathbf{q} = 0, \quad (1)$$

$$\partial_t \mathbf{q} + \nabla_{\mathbf{x}} \cdot \left(\frac{\mathbf{q} \otimes \mathbf{q}}{\rho} \right) + \nabla_{\mathbf{x}} p = 0, \quad (2)$$

$$\partial_t E + \nabla_{\mathbf{x}} \cdot \left((E + p) \frac{\mathbf{q}}{\rho} \right) = 0, \quad (3)$$

where $\rho(t, \mathbf{x}) > 0$ is the density of the fluid, $\mathbf{q}(t, \mathbf{x}) = (\rho(t, \mathbf{x}) \mathbf{u}(t, \mathbf{x})) \in \mathbb{R}^d$ the momentum, $\mathbf{u}(t, \mathbf{x})$ represents the velocity vector, $E(t, \mathbf{x})$ is the total energy and $p \equiv p(\rho, \mathbf{q}, E)$ is the fluid pressure defined by the following equation of state

$$p \equiv p(\rho, \mathbf{q}, E) = (\gamma - 1) \left(E - \frac{1}{2} \frac{|\mathbf{q}|^2}{\rho} \right). \quad (4)$$

The parameter $\gamma = c_p/c_v \geq 1$ is a characterization of the fluid and represents the ratio of specific heats, namely the heat capacity at constant volume c_v and at constant pressure c_p . In the following, the state variables are often gathered into one $(d + 2)$ -vector $\mathbf{Q} = (\rho, \mathbf{q}, E)$. The low Mach number limit can be obtained from this model by rescaling the equations in the following way:

$$\tilde{\rho} = \rho/\rho_0, \quad \tilde{\mathbf{u}} = \mathbf{u}/u_0, \quad \tilde{p} = p/p_0, \quad \tilde{E} = E/p_0, \quad \tilde{\mathbf{x}} = \mathbf{x}/x_0, \quad \tilde{t} = t/t_0, \quad (5)$$

where ρ_0, p_0, x_0, t_0 and $u_0 = x_0/t_0$ are typical values encountered in the problems under consideration. Now, by defining $\varepsilon = \rho_0 u_0^2 / p_0$, inserting (5) into system (1)-(4) and omitting the tildes, one obtains the rescaled Euler equations

$$\partial_t \rho + \nabla_{\mathbf{x}} \cdot \mathbf{q} = 0, \quad (6)$$

$$\partial_t \mathbf{q} + \nabla_{\mathbf{x}} \cdot \left(\frac{\mathbf{q} \otimes \mathbf{q}}{\rho} \right) + \frac{1}{\varepsilon} \nabla_{\mathbf{x}} p = 0, \quad (7)$$

$$\partial_t E + \nabla_{\mathbf{x}} \cdot \left((E + p) \frac{\mathbf{q}}{\rho} \right) = 0, \quad (8)$$

and the corresponding rescaled equation of state

$$p = (\gamma - 1) \left(E - \frac{\varepsilon}{2} \frac{|\mathbf{q}|^2}{\rho} \right). \quad (9)$$

The quantity $\varepsilon > 0$ is related to the Mach number M through the relation $M = u_0/c = \sqrt{\varepsilon/\gamma}$, where c is the sound speed given by $c^2 = \left(\frac{\partial p}{\partial \rho} \right)_s = \gamma \frac{p}{\rho}$, s being the entropy [56, 64].

The Low-Mach regimes correspond to situations in which the sound speed in the fluid is very high compared to the speed of the fluid itself, and consequently to small values of ε . In such situations, if an explicit scheme is used, the rescaled sound speed dictates a severe CFL stability condition that drastically limits the maximum admissible size of the time step. Indeed, for instance for $d = 3$ and setting $\mathbf{x} = (x, y, z)$ and $\mathbf{u} = (u, v, w)$, the eigenvalues of the Jacobian matrix in the direction x are $\lambda_{1x} = u - c/\sqrt{\varepsilon}$, $\lambda_{2x} = u$ (multiplicity 3) and $\lambda_{3x} = u + c/\sqrt{\varepsilon}$ (the same results hold for the directions y and z changing u by v and w , respectively). The time-step $\Delta t = t^{n+1} - t^n$ must then satisfy the CFL condition, which is given by

$$\Delta t \leq \frac{1}{\frac{\max(|u \pm c/\sqrt{\varepsilon}|)}{\Delta x} + \frac{\max(|v \pm c/\sqrt{\varepsilon}|)}{\Delta y} + \frac{\max(|w \pm c/\sqrt{\varepsilon}|)}{\Delta z}},$$

where Δx , Δy and Δz are the respective space steps for the different directions. Then, Δt is of order $\sqrt{\varepsilon}$ and tends to 0 with ε . Furthermore, even if this constraint is satisfied, it is well known (see [42], [41] or [27]) that explicit schemes suffer from a consistency problem in the limit $\varepsilon \rightarrow 0$. Indeed, they are not capable to capture the right asymptotic regime.

A possible way to bypass these limitations is to use the incompressible Euler equations (IE) obtained as the low Mach number limit of the compressible Euler equations (CE), i.e. $\lim_{\varepsilon \rightarrow 0} CE = IE$. This limit is formally recalled in the next Section. However, this solution can be used only if the fluid is incompressible in the entire computational domain. For situations such that compressible zones coexist with incompressible ones, alternative remedies should be studied. This is the case of asymptotic stable schemes which are free of time step constraints related to fast temporal scales. To construct such schemes, it is necessary to determine what are the terms, in the compressible Euler equations, that carry the sound waves information. In [90], it has been shown that the pressure term in the momentum equations and the enthalpy term in the energy equation are those generating these fast sound waves that constrain the time step. Therefore, in the sequel we choose to implicitly discretize these terms to avoid a restrictive stability condition for low Mach number flows.

We end this section by providing notations useful in the sequel. The specific enthalpy h is given by

$$h = e + p, \quad (10)$$

where e is the internal energy e given by

$$e = \frac{p}{\gamma - 1}. \quad (11)$$

Then, the total energy is the sum of the internal energy e and the kinetic energy $k = \frac{\varepsilon}{2}\rho|\mathbf{u}|^2$ (respectively $k = \frac{1}{2}\rho|\mathbf{u}|^2$ in physical variables), and the total energy flux in (3) can be reformulated according to [90] as

$$\mathbf{u}(E + p) = k\mathbf{u} + h\mathbf{u}. \quad (12)$$

The compressible Euler equations can be written in expanded form in the three dimensional case by the equations

$$\mathbf{Q}_t + \mathbf{F}_x + \mathbf{G}_y + \mathbf{W}_z = \mathbf{0}, \quad (13)$$

where the vectors of conservative variables \mathbf{Q} and the fluxes ($\mathbf{F}(\mathbf{Q})$, $\mathbf{G}(\mathbf{Q})$, $\mathbf{W}(\mathbf{Q})$) explicitly write

$$\mathbf{Q} = \begin{bmatrix} \rho \\ \rho u \\ \rho v \\ \rho w \\ E \end{bmatrix}, \quad \mathbf{F}(\mathbf{Q}) = \begin{bmatrix} \rho u \\ \rho u^2 + \frac{1}{\varepsilon} p \\ \rho uv \\ \rho uw \\ u(k+h) \end{bmatrix}, \quad \mathbf{G}(\mathbf{Q}) = \begin{bmatrix} \rho v \\ \rho vu \\ \rho v^2 + \frac{1}{\varepsilon} p \\ \rho vw \\ v(k+h) \end{bmatrix}, \quad \mathbf{W}(\mathbf{Q}) = \begin{bmatrix} \rho w \\ \rho wu \\ \rho vw \\ \rho w^2 + \frac{1}{\varepsilon} p \\ w(k+h) \end{bmatrix}, \quad (14)$$

with $\mathbf{x} = (x, y, z) \in \Omega$ and $\mathbf{u} = (u, v, w)$. Additional useful relations (respectively the scaled and the non scaled ones) are

$$E = e + k = e + \frac{\varepsilon}{2}\rho|\mathbf{u}|^2, \quad E = e + k = e + \frac{1}{2}\rho|\mathbf{u}|^2. \quad (15)$$

Before introducing the numerical scheme, we give the details of the formal derivation of the low-Mach (or incompressible) limit of the compressible Euler equations which will be useful in the analysis of the numerical scheme.

2.1. Low Mach limit and Incompressible Euler equations

The rigorous low-Mach limit of the compressible Euler system has been widely studied in the last years [2, 56, 57, 62, 81]. In the isentropic case, the solutions are known to converge to the solution of the corresponding incompressible Euler equations independently from the choice of the initial data (well prepared/consistent or not to the limit regime). Results in the case of non-isentropic Euler equations with general initial data can be found in [64] in the free space $\Omega = \mathbb{R}^d$, in [1] for an exterior domain and in a bounded toroidal domain in [65].

Here, we briefly recall the formal limit for a general bounded domain in the simplified case of well-prepared initial conditions. We supplement system (6)-(9), or equivalently (9)-(15), with the permeability boundary condition

$$\mathbf{u}(t, \mathbf{x}) \cdot \boldsymbol{\nu}(\mathbf{x}) = 0, \quad \forall \mathbf{x} \in \partial\Omega, \quad t > 0, \quad (16)$$

where $\boldsymbol{\nu}(\mathbf{x})$ denotes the unit outward normal vector to the boundary $\partial\Omega$ of the domain Ω . We consider well prepared boundary conditions

$$(\rho, \mathbf{u}, p)(t = 0, \mathbf{x}) = (\rho^0(\mathbf{x}), \mathbf{u}^0(\mathbf{x}), p^0(\mathbf{x})), \quad \forall \mathbf{x} \in \Omega, \quad (17)$$

where

$$\rho^0(\mathbf{x}) \geq 0, \quad p^0(\mathbf{x}) = \bar{p}_0 + \varepsilon \tilde{p}_0(\mathbf{x}) \geq 0, \quad \mathbf{u}^0(\mathbf{x}) = \bar{\mathbf{u}}_0(\mathbf{x}) + \sqrt{\varepsilon} \tilde{\mathbf{u}}_0(\mathbf{x}), \quad \nabla_{\mathbf{x}} \cdot \bar{\mathbf{u}}_0(\mathbf{x}) = 0, \quad (18)$$

where \bar{p}_0 is a fixed constant. We assume that all quantities ρ , \mathbf{u} , p and E , solutions of (6)-(9), converge respectively towards some values ρ_0 , \mathbf{u}_0 , p_0 and E_0 when ε tends to 0. Then, multiplying the momentum equations and passing to the limit lead to $\nabla_{\mathbf{x}} p_0 = 0$ and, consequently, $p_0(t, \mathbf{x}) = p_0(t)$. The low Mach limit of the equation of state (9) yields moreover $E_0(t, \mathbf{x}) = E_0(t) = p_0(t)/(\gamma - 1)$. Integrating now the limit of the energy equation (8) on the domain, using the Green formula and the limit permeability condition (16), gives

$$|\Omega| E_0'(t) + (E_0(t) + p_0(t)) \int_{\partial\Omega} \mathbf{u}_0(\mathbf{x}, t) \cdot \boldsymbol{\nu}(\mathbf{x}) d\sigma(\mathbf{x}) = |\Omega| E_0'(t) = 0,$$

where $d\sigma$ is the $d - 1$ -dimensional Lebesgue measure on $\partial\Omega$. Then, $E_0(t) = E_0(0) = \bar{p}_0/(\gamma - 1)$ and $p_0(t) = \bar{p}_0$.

By considering again the energy equation (8), one notices that the constant limit pressure implies the incompressibility condition $\nabla_{\mathbf{x}} \cdot \mathbf{u}_0 = 0$. Finally, assuming that $p_1 = \lim_{\varepsilon \rightarrow 0} \frac{1}{\varepsilon} (p - \bar{p}_0)$ exists, we recover the incompressible limit system [1], that is

$$\partial_t \rho_0 + \nabla_{\mathbf{x}} \cdot \mathbf{q}_0 = 0, \quad (19)$$

$$\partial_t \mathbf{q}_0 + \nabla_{\mathbf{x}} \cdot \left(\frac{\mathbf{q}_0 \otimes \mathbf{q}_0}{\rho_0} \right) + \nabla_{\mathbf{x}} p_1 = 0, \quad (20)$$

$$\nabla_{\mathbf{x}} \cdot \mathbf{u}_0 = 0. \quad (21)$$

We can observe that $p = \bar{p}_0 + \varepsilon p_1 + o(\varepsilon^2)$ and that p_1 is implicitly defined by the constraint $\nabla_{\mathbf{x}} \cdot \mathbf{u}_0 = 0$ and explicitly given by the equation

$$-\nabla_{\mathbf{x}} \cdot \left(\frac{1}{\rho_0} \nabla_{\mathbf{x}} p_1 \right) = \nabla_{\mathbf{x}} \cdot (\mathbf{u}_0 \cdot \nabla_{\mathbf{x}} \mathbf{u}_0).$$

In addition, the density conservation equation reduces to $\partial_t \rho_0 + \mathbf{u}_0 \cdot \nabla_{\mathbf{x}} \rho_0 = 0$, which is a pure transport equation at the velocity \mathbf{u}_0 . Note that if ρ_0 is initially constant then ρ_0 remains constant for all times and the explicit equation for p_1 becomes

$$-\Delta_{\mathbf{x}} p_1 = \nabla_{\mathbf{x}} : (\rho_0 \mathbf{u}_0 \otimes \mathbf{u}_0).$$

The incompressible system can also be expressed using the energy variable, remarking that

$$\lim_{\varepsilon \rightarrow 0} \frac{1}{\varepsilon} \left(E - \frac{\bar{p}_0}{\gamma - 1} \right) = \lim_{\varepsilon \rightarrow 0} \frac{1}{\varepsilon} (E - \bar{E}_0) = \frac{1}{\gamma - 1} \lim_{\varepsilon \rightarrow 0} \left(\frac{1}{\varepsilon} (p - \bar{p}_0) + \frac{1}{2} \frac{|\mathbf{q}|^2}{\rho} \right) = \frac{p_1}{\gamma - 1} + \frac{1}{2} \frac{|\mathbf{q}_0|^2}{\rho_0}.$$

Then, defining E_1 by $E_1 = \frac{p_1}{\gamma - 1} + \frac{1}{2} \frac{|\mathbf{q}_0|^2}{\rho_0}$, we obtain $E = \bar{E}_0 + \varepsilon E_1 + o(\varepsilon^2)$ and system (19)-(21) becomes

$$\partial_t \rho_0 + \nabla_{\mathbf{x}} \cdot \mathbf{q}_0 = 0, \quad (22)$$

$$\partial_t \mathbf{q}_0 + \nabla_{\mathbf{x}} \cdot \left(\frac{\mathbf{q}_0 \otimes \mathbf{q}_0}{\rho_0} \right) + (\gamma - 1) \nabla_{\mathbf{x}} \tilde{E}_1 - \frac{\gamma - 1}{2} \nabla_{\mathbf{x}} \left(\frac{|\mathbf{q}_0|^2}{\rho_0} \right) = 0, \quad (23)$$

$$\nabla_{\mathbf{x}} \cdot \mathbf{u}_0 = 0. \quad (24)$$

3. A first order numerical scheme

In this Section, we introduce and discuss a first order in time and space discretization of the Euler equations (13). We start from a semi discretization in time in the general d -dimensional space, then we derive a full space-time discretization in one space dimension and we end by giving the details of the full discretized three dimensional algorithm.

3.1. A first order time discretization

In this part, we discuss the semi-implicit time discretization we decide to follow and we study its properties. The class of schemes considered hereafter belong to the so-called IMEX (IMplicit-EXplicit) time discretization methods (see [24, 25, 85, 30, 29, 8, 10, 9, 3, 72]). In particular, following [90], the fluxes in the governing equations (14) are split into a non stiff part $(\mathbf{F}^e, \mathbf{G}^e, \mathbf{W}^e)$, which will be discretized explicitly, and a stiff part $(\mathbf{F}^i, \mathbf{G}^i, \mathbf{W}^i)$, which will be discretized implicitly. This partition write

$$\begin{aligned} \mathbf{F} &= \mathbf{F}^e + \mathbf{F}^i, & \mathbf{G} &= \mathbf{G}^e + \mathbf{G}^i, & \mathbf{W} &= \mathbf{W}^e + \mathbf{W}^i, \\ &= \begin{bmatrix} \rho u \\ \rho u^2 \\ \rho uv \\ \rho uw \\ uk \end{bmatrix} + \begin{bmatrix} 0 \\ p \\ \varepsilon \\ 0 \\ 0 \\ hu \end{bmatrix}, & &= \begin{bmatrix} \rho v \\ \rho vu \\ \rho v^2 \\ \rho vw \\ vk \end{bmatrix} + \begin{bmatrix} 0 \\ 0 \\ p \\ \varepsilon \\ 0 \\ hv \end{bmatrix}, & &= \begin{bmatrix} \rho w \\ \rho wu \\ \rho wv \\ \rho w^2 \\ wk \end{bmatrix} + \begin{bmatrix} 0 \\ 0 \\ 0 \\ p \\ \varepsilon \\ hw \end{bmatrix}. \end{aligned} \quad (25)$$

The explicit fluxes contribute to the advection sub-system, while the implicit ones correspond to the pressure (or sound) waves in the fluid. In [90] a detailed study of the eigenstructure of both explicit-advection and implicit-pressure sub-systems prove that both are hyperbolic systems. Therefore, the problem can be recast as

$$\mathbf{Q}_t + (\mathbf{F}^e + \mathbf{F}^i)_x + (\mathbf{G}^e + \mathbf{G}^i)_y + (\mathbf{W}^e + \mathbf{W}^i)_z = \mathbf{0}, \quad (26)$$

that yields the first order in time semi-discretization

$$\underbrace{\mathbf{Q}^{n+1} + \Delta t \left(\mathbf{F}_x^i(\mathbf{Q}^{n+1}) + \mathbf{G}_y^i(\mathbf{Q}^{n+1}) + \mathbf{W}_z^i(\mathbf{Q}^{n+1}) \right)}_{\text{Implicit}} = \underbrace{\mathbf{Q}^n - \Delta t \left(\mathbf{F}_x^e(\mathbf{Q}^n) + \mathbf{G}_y^e(\mathbf{Q}^n) + \mathbf{W}_z^e(\mathbf{Q}^n) \right)}_{\text{Explicit}}. \quad (27)$$

For the sake of clarity and compact notation in the description of the methods, let us call $\mathcal{L}_i = (\mathbf{F}^i, \mathbf{G}^i, \mathbf{W}^i)$ the implicit operator and $\mathcal{L}_e = (\mathbf{F}^e, \mathbf{G}^e, \mathbf{W}^e)$ the explicit one. In this way, the scheme (27) is rewritten as

$$\mathbf{Q}^{n+1} = \mathbf{Q}^n - \Delta t \nabla_x \cdot \mathcal{L}_i(\mathbf{Q}^{n+1}) - \Delta t \nabla_x \cdot \mathcal{L}_e(\mathbf{Q}^n). \quad (28)$$

We show in Lemma 1 that this choice is enough to ensure the asymptotic consistency of the semi-discrete scheme, i.e. the consistency with the incompressible limit. In [30, 29] it has been shown on a simplified problem mimicking the isentropic model that stability can be also ensured with a CFL related to the eigenvalues of the explicit part of the matrix. We believe that such a study can be extended to the present system in its linearised form. Nevertheless, we did not perform a stability analysis in this work and we postpone such analysis to future investigations. Instead, we show in the numerical results presented in Section 5 that stability is indeed assured independently on the speed of the fast pressure waves.

The construction of numerical schemes which are capable to capture a limit model starting from another set of equations, such as the example described in the previous Section of the compressible equations which reduce to the incompressible ones, is closely connected with the notion of asymptotic-preserving schemes (see Figure 1 for a visual explanation). Here, in agreement with [25, 26, 72], we give the following definitions of asymptotic preserving methods for system (14).

Definition 1. *A consistent time discretization method for (14) supplemented by the equation of state (9) of stepsize Δt is said*

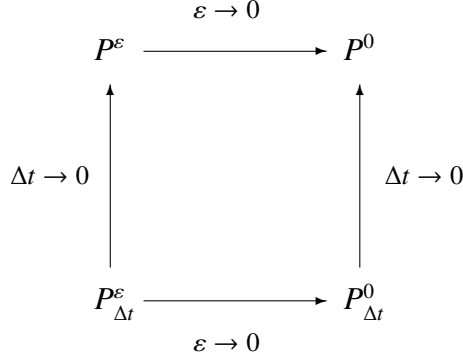


Figure 1: P^ϵ is the original singular perturbation problem, i.e. the compressible Euler model, and $P_{\Delta t}^\epsilon$ its numerical approximation characterized by a discretization parameter Δt . The asymptotic-preserving (AP) property corresponds to the request that for a given Δt independent of ϵ , when $\epsilon \rightarrow 0$, $P_{\Delta t}^\epsilon$ is a consistent and stable discretization of P^0 , which is the incompressible Euler model.

- *Strong asymptotic preserving if, independently of the initial data, it is asymptotically stable and consistent. That is if, it is stable for time steps Δt which can be chosen independently of ϵ , and it becomes a consistent time discretization method for the limit system (19)-(21) in the limit $\epsilon \rightarrow 0$.*
- *Weak asymptotic preserving if, it is asymptotically stable and it is asymptotically consistent when the initial data are well prepared to the limit equations. That is if, it is stable for time steps Δt which can be chosen independently of ϵ , and if it becomes a consistent time discretization method for the limit system (19)-(21) in the limit $\epsilon \rightarrow 0$.*

These definitions do not imply that the scheme preserves the order of accuracy in time in the stiff limit $\epsilon \rightarrow 0$. In this latter case we say that the scheme is asymptotically accurate (AA). The definition of consistent or well-prepared initial data is given in the following.

Definition 2. *The initial data for equation (14) and the relative equation of state (9) are said consistent or well-prepared if*

$$p_\epsilon(t=0, \mathbf{x}) = \bar{p}_0 + \epsilon p_1(\mathbf{x}), \quad \bar{p}_0, p_1(\mathbf{x}) \in \mathbb{R}^+, \quad \mathbf{u}_\epsilon(t=0, \mathbf{x}) = \mathbf{u}_0(\mathbf{x}) + \epsilon \mathbf{u}_1(\mathbf{x}), \quad \nabla \cdot \mathbf{u}_0(\mathbf{x}) = 0.$$

That is, if the pressure field reduces to a constant value in the limit, and the divergence of the velocity field is equal to zero in the same limit.

In what follows, we will consider the development of asymptotic preserving and asymptotically accurate schemes using the general setting of IMEX Runge-Kutta methods [3, 71]. We start by giving a formal proof that this first order in time explicit-implicit approach (28) provides a consistent discretization of the limit model when $\epsilon \rightarrow 0$.

Lemma 1. *We denote by $(\rho_\epsilon^n(\mathbf{x}), \mathbf{u}_\epsilon^n(\mathbf{x}), E_\epsilon^n(\mathbf{x}))_{n \geq 1}$ the solution of the semi discrete numerical scheme (28) with the permeability boundary condition $\mathbf{u}^n(\mathbf{x}) \cdot \mathbf{\nu}(\mathbf{x}) = 0, \forall \mathbf{x} \in \partial\Omega, \forall n \geq 0$, Then:*

1. *assuming that initial data are well prepared to the incompressible regime, i.e. for all $\mathbf{x} \in \Omega, p^0(\mathbf{x}) = \bar{p}_0 \in \mathbb{R}^+$ and $\nabla \cdot \mathbf{u}^0(\mathbf{x}) = \nabla \cdot \mathbf{u}_0(\mathbf{x}) = 0$, the semi discrete numerical scheme (28) is asymptotic consistent, i.e. in the limit $\epsilon \rightarrow 0$ it becomes a consistent discretization of system (19)-(21). This discretization is given for all $n \geq 0$ by*

$$\frac{\rho^{n+1}(\mathbf{x}) - \rho^n(\mathbf{x})}{\Delta t} + \mathbf{u}^n(\mathbf{x}) \cdot \nabla_{\mathbf{x}} \rho^n(\mathbf{x}) = 0, \tag{29}$$

$$\frac{(\rho^{n+1} \mathbf{u}^{n+1})(\mathbf{x}) - (\rho^n \mathbf{u}^n)(\mathbf{x})}{\Delta t} + \nabla_{\mathbf{x}} \cdot ((\rho^n \mathbf{u}^n)(\mathbf{x}) \otimes \mathbf{u}^n(\mathbf{x})) + \nabla_{\mathbf{x}} p_1^{n+1}(\mathbf{x}) = 0, \tag{30}$$

$$\nabla_{\mathbf{x}} \cdot \mathbf{u}^{n+1}(\mathbf{x}) = 0. \tag{31}$$

2. For general initial data, the semi discrete numerical scheme (28) is strongly asymptotically consistent from the second time iteration ($n \geq 2$). This means that in the limit $\varepsilon \rightarrow 0$ it becomes a consistent discretization of system (19)-(21). This is given for all $n \geq 1$ by equation (30) and

$$\frac{\rho^{n+1}(\mathbf{x}) - \rho^n(\mathbf{x})}{\Delta t} + \nabla_{\mathbf{x}} \cdot (\rho^n \mathbf{u}^n(\mathbf{x})) = 0, \quad \text{if } n = 1 \quad (32)$$

$$\frac{\rho^{n+1}(\mathbf{x}) - \rho^n(\mathbf{x})}{\Delta t} + \mathbf{u}^n \cdot \nabla_{\mathbf{x}}(\rho^n(\mathbf{x})) = 0, \quad \text{if } n \geq 2 \quad (33)$$

$$\nabla_{\mathbf{x}} \cdot \mathbf{u}^{n+1}(\mathbf{x}) = \begin{cases} 0, & \text{if } n \geq 2, \\ -\frac{\int_{\Omega} \bar{p}^0(\mathbf{x}) d\mathbf{x} - |\Omega| \bar{p}^0(\mathbf{x})}{\Delta t \gamma \int_{\Omega} \bar{p}^0(\mathbf{x}) d\mathbf{x}}, & \text{if } n = 1. \end{cases} \quad (34)$$

Proof. We assume that for all $n \geq 1$ all quantities $\rho_{\varepsilon}^n(\mathbf{x})$, $\mathbf{u}_{\varepsilon}^n(\mathbf{x})$, $p_{\varepsilon}^n(\mathbf{x})$ and $E_{\varepsilon}^n(\mathbf{x})$, which are solutions of (28), converge respectively towards some values $\rho^n(\mathbf{x})$, $\mathbf{u}^n(\mathbf{x})$, $p^n(\mathbf{x})$ and $E^n(\mathbf{x})$ when ε tends to 0. Under this hypothesis, multiplying by ε the momentum equations of (28) and letting ε tend to zero, gives, for all $n \geq 0$, $\nabla_{\mathbf{x}} p^{n+1}(\mathbf{x}) = 0$, i.e. the pressure at time index $(n + 1)$ is constant independently on the initial data. From the equation of state (28), we have

$$E_{\varepsilon}^{n+1}(\mathbf{x}) = e_{\varepsilon}^{n+1}(\mathbf{x}) + \frac{\varepsilon}{2} \rho_{\varepsilon}^{n+1}(\mathbf{x}) |\mathbf{u}_{\varepsilon}^{n+1}(\mathbf{x})|^2 = \frac{p_{\varepsilon}^{n+1}(\mathbf{x})}{(\gamma - 1)} + \frac{\varepsilon}{2} \rho_{\varepsilon}^{n+1}(\mathbf{x}) |\mathbf{u}_{\varepsilon}^{n+1}(\mathbf{x})|^2,$$

that gives $E^{n+1}(\mathbf{x}) = p^{n+1}(\mathbf{x})/(\gamma - 1)$ for all $n \geq 1$, in the limit $\varepsilon \rightarrow 0$. The limit of the energy equation yields, for all $n \geq 1$, all $\mathbf{x} \in \Omega$ and all initial data

$$E^{n+1}(\mathbf{x}) = E^n(\mathbf{x}) - \Delta t (e^{n+1}(\mathbf{x}) + p^{n+1}(\mathbf{x})) \nabla_{\mathbf{x}} \cdot \mathbf{u}^{n+1}(\mathbf{x}). \quad (35)$$

Thus, integrating this equation on the domain, using the Green formula and the limit of the permeability boundary condition leads to

$$|\Omega| E^{n+1}(\mathbf{x}) = \int_{\Omega} E^n(\mathbf{x}) d\mathbf{x} - \Delta t (e^{n+1}(\mathbf{x}) + p^{n+1}(\mathbf{x})) \int_{\partial\Omega} \mathbf{u}^{n+1}(\mathbf{x}) \cdot \mathbf{v}(\mathbf{x}) d\sigma(\mathbf{x}) = \int_{\Omega} E^n(\mathbf{x}) d\mathbf{x}. \quad (36)$$

Now, if initial data are well prepared, this implies that for the first time step we have

$$\frac{p^1(\mathbf{x})}{\gamma - 1} = E^1(\mathbf{x}) = \frac{1}{|\Omega|} \int_{\Omega} E^0(\mathbf{x}) d\mathbf{x} = \frac{1}{|\Omega|} \int_{\Omega} \frac{p^0(\mathbf{x})}{\gamma - 1} d\mathbf{x} = \frac{1}{|\Omega|} \int_{\Omega} \frac{\bar{p}_0}{\gamma - 1} d\mathbf{x} = \frac{\bar{p}_0}{\gamma - 1}.$$

Consequently, for all $n \geq 1$, both pressure and energy remain constant: $p^{n+1} = \bar{p}_0$, $E^{n+1} = \bar{p}_0/(\gamma - 1)$. Using again (35), gives now for all $n \geq 0$, $\nabla_{\mathbf{x}} \cdot \mathbf{u}^{n+1}(\mathbf{x}) = 0$, i.e. equations (29)-(31). Assuming finally that the following limit for $\varepsilon \rightarrow 0$ exists

$$\lim_{\varepsilon \rightarrow 0} \frac{1}{\varepsilon} (p_{\varepsilon}^{n+1} - p^{n+1}) = \lim_{\varepsilon \rightarrow 0} \frac{1}{\varepsilon} \left(p_{\varepsilon}^{n+1} - \frac{1}{|\Omega|} \int_{\Omega} \bar{p}^0(\mathbf{x}) d\mathbf{x} \right) = p_1^{n+1},$$

and passing to the limit into the momentum equation of (28), we obtain (30).

We discuss now the case in which initial data are not well prepared to the limit incompressible system. In this case, we have again that for $n = 1$ the energy is projected over a constant state: $E^1(\mathbf{x}) = \frac{1}{|\Omega|} \int_{\Omega} E^0(\mathbf{x}) d\mathbf{x}$. However, since the initial condition does not force constant values for pressure and consequently for the energy, we have that the incompressibility condition is not preserved neither at the beginning $\nabla_{\mathbf{x}} \cdot \mathbf{u}^0 \neq 0$ nor for the first time step. This is given precisely by

$$\nabla_{\mathbf{x}} \cdot \mathbf{u}^1 = -\frac{E^1 - E^0(\mathbf{x})}{\Delta t (e^1 + p^1)}. \quad (37)$$

Nevertheless, from $n = 2$, following equation (36), we have $E^{n+1} = E^n$ which means that energy is constant in space and thus $\nabla_{\mathbf{x}} \cdot \mathbf{u}^{n+1}(\mathbf{x}) = 0$, i.e. we are solving equations (29)-(31). The equation for p_1^{n+1} is obtained as done for the case of well prepared initial data passing to the limit in equation (28). This concludes the proof. \square

In the next Section, we give the details of the one dimensional time and space discretization and we discuss in practice how we solve the implicit terms in the non linear system (28).

3.2. A full time and space Asymptotic Preserving scheme for the Euler equations

We present now the full space and time discretization of system (26) for the one-dimensional case. We will then extend this scheme in Section 3.4 to the full three dimensional setting. In one space dimension, the governing equations read

$$\mathbf{Q}_t + \mathbf{F}_x = \mathbf{0}, \quad x \in \Omega \subset \mathbb{R}, \quad t \in \mathbb{R}_0^+, \quad \mathbf{Q} \in \Omega_Q \subset \mathbb{R}^3, \quad (38)$$

with the vector of conservative variables given by $\mathbf{Q} = [\rho, \rho u, E]^T$. The flux vector is split into an explicit and implicit part according to (25), hence obtaining

$$\mathbf{F} = \mathbf{F}^e + \mathbf{F}^i = \begin{bmatrix} \rho u \\ \rho u^2 \\ uk \end{bmatrix} + \begin{bmatrix} 0 \\ \frac{p}{\varepsilon} \\ \frac{h}{\rho} \\ -q \end{bmatrix}. \quad (39)$$

The computational domain reduces to $\Omega = [x_{\min}, x_{\max}]$ and it is discretized with a total number N_x of control volumes of uniform size $\Delta x = \frac{x_{\max} - x_{\min}}{N_x}$. A cell-centered discretization is used, where index $i \in [1, N_x]$ denotes the cell, while $i \pm 1/2$ refers to the interface between cell i and $i \pm 1$, respectively. The cell center is located at x_i and the interface is consistently located at $x_{i+1/2} = \frac{1}{2}(x_i + x_{i+1}) = x_i + \Delta x/2$. Because of the Implicit-Explicit approach, we choose a time step which is restricted by a CFL condition that is only driven by the eigenvalues of explicit part and, as such, it is only dependent on the fluid velocity u , and *not* by the rescaled sound speed $\frac{c}{\sqrt{\varepsilon}}$, that is

$$\Delta t \leq \text{CFL} \frac{\Delta x}{\max_i |2u_i^n|}, \quad (40)$$

thus making our numerical method particularly well suited for the discretization of low Mach number flows. We stress that a similar approach but on staggered grids has been recently proposed in [35] for general equations of state.

Let us introduce the explicit operator $F^e [m_i^n]$ which applies to a generic cell quantity m_i^n :

$$F^e [m_i^n] = m_i^n - \frac{\Delta t}{\Delta x} (f_{i+1/2}^m - f_{i-1/2}^m), \quad (41)$$

where the numerical fluxes $f_{i\pm 1/2}^m$ are chosen to be of Rusanov-type

$$\begin{aligned} f_{i+1/2}^m &= \frac{1}{2} (f(m_{i+1}^n) + f(m_i^n)) - \frac{1}{2} |a_{i+1/2}^n| (m_{i+1}^n - m_i^n), & a_{i+1/2}^n &= \max(|f'(m_{i+1}^n)|, |f'(m_i^n)|), \\ f_{i-1/2}^m &= \frac{1}{2} (f(m_i^n) + f(m_{i-1}^n)) - \frac{1}{2} |a_{i-1/2}^n| (m_i^n - m_{i-1}^n), & a_{i-1/2}^n &= \max(|f'(m_i^n)|, |f'(m_{i-1}^n)|), \end{aligned} \quad (42)$$

with $f(\cdot)$ being the physical flux and $|f'(m_i^n)|$ denoting the maximum eigenvalue of the Jacobian matrix associated to the flux $f(\cdot)$, i.e. the maximum speed of the flow u at the interface at time index n , since we have separated the pressure and the average speed contributions. The governing system (38) with the flux splitting (39) is then discretized using (41) for \mathbf{F}^e and a centred numerical flux for the implicit part \mathbf{F}^i . Following [30], we believe that this choice is sufficient for obtaining an L^2 stable scheme. We postpone such a study to future investigations.

In details, the resulting numerical scheme reads as follows.

1. Density equation with $m_i^n \equiv \rho_i^n$:

$$\rho_i^{n+1} = F^e [\rho_i^n]. \quad (43)$$

2. Momentum equation with $m_i^n \equiv q_i^n = (\rho u)_i^n$:

$$q_i^{n+1} = F^e [q_i^n] - \frac{\Delta t}{2\varepsilon \Delta x} (p_{i+1}^{n+1} - p_{i-1}^{n+1}), \quad (44)$$

where the implicit discretization of the pressure corresponds to the use of the Rusanov flux without numerical dissipation.

3. Energy equation with $m_i^n \equiv k_i^n$:

$$E_i^{n+1} = F^e [E_i^n] - \frac{\Delta t}{2\Delta x} \left(\tilde{h}_{i+1/2}^{n+1} (q_{i+1}^{n+1} + q_i^{n+1}) - \tilde{h}_{i-1/2}^{n+1} (q_i^{n+1} + q_{i-1}^{n+1}) \right), \quad (45)$$

where the enthalpy on the interfaces is defined as

$$\tilde{h}_{i+1/2}^{n+1} = \frac{\frac{h_i^{n+1}}{\rho_i^{n+1}} q_i^{n+1} + \frac{h_{i+1}^{n+1}}{\rho_{i+1}^{n+1}} q_{i+1}^{n+1}}{q_i^{n+1} + q_{i+1}^{n+1}}. \quad (46)$$

If the fluid is at rest, thus the denominator vanishes in (46), the enthalpy at the interface is simply computed by the arithmetic average: $\tilde{h}_{i+1/2}^{n+1} = \frac{1}{2} \left(\frac{h_i^{n+1}}{\rho_i^{n+1}} + \frac{h_{i+1}^{n+1}}{\rho_{i+1}^{n+1}} \right)$. Note that the implicit flux in the discrete energy equation (45) is not a standard centred approximation but instead it corresponds to a mean centred approximation. This choice is enough to avoid a checkerboard effect in the resulting elliptic equation for the pressure [80] we will derive later on in this article.

The proposed numerical method is built along the lines of [35] and it leads to a non linear system to be solved in order to advance the solution in time from n to $(n + 1)$. This solution can be obtained by splitting the total energy E_i^{n+1} into an internal and a kinetic contribution according to (15). Then, one can insert the momentum equation (44) into the energy equation (45) obtaining

$$\begin{aligned} e_i^{n+1} + k_i^{n+1} &= F^e [E_i^n] \\ &- \frac{1}{2} \frac{\Delta t}{\Delta x} \tilde{h}_{i+1/2}^{n+1} \left[(F^e [q_{i+1}^n] + F^e [q_i^n]) - \frac{1}{2} \frac{\Delta t}{\varepsilon \Delta x} (p_{i+2}^{n+1} - p_i^{n+1} + p_{i+1}^{n+1} - p_{i-1}^{n+1}) \right] \\ &+ \frac{1}{2} \frac{\Delta t}{\Delta x} \tilde{h}_{i-1/2}^{n+1} \left[(F^e [q_i^n] + F^e [q_{i-1}^n]) - \frac{1}{2} \frac{\Delta t}{\varepsilon \Delta x} (p_{i+1}^{n+1} - p_{i-1}^{n+1} + p_i^{n+1} - p_{i-2}^{n+1}) \right], \end{aligned} \quad (47)$$

where the kinetic energy at time t^n is computed as

$$k_i^{n+1} = \frac{1}{2} \rho_i^{n+1} \left(\frac{q_i^{n+1}}{\rho_i^{n+1}} \right)^2. \quad (48)$$

In the energy equation (47), the variables which are already known at time t^{n+1} are the density ρ_i^{n+1} , because the continuity equation is solved explicitly by (43), and the explicit operators $F^e [q_i^n]$ and $F^e [E_i^n]$. The resulting system of equations is strongly nonlinear. In order to reduce this non linearity, one can first observe that for an ideal gas, the internal energy is a linear function of p , namely $e = \frac{p}{\gamma - 1}$, according to (11). Then, one can rely on a Picard algorithm to converge to the solution at time t^{n+1} by selecting only the pressure terms as implicit ones in the iterative method. Thanks to this strategy, the following mildly nonlinear system is obtained for the pressure $p_i^{n+1,r+1}$ as the only unknown at each Picard iteration r :

$$\begin{aligned} &- \frac{1}{4} \frac{\Delta t^2}{\Delta x} \tilde{h}_{i+1/2}^{n+1,r} p_{i+2}^{n+1,r+1} - \frac{1}{4} \frac{\Delta t^2}{\Delta x} (\tilde{h}_{i+1/2}^{n+1,r} - \tilde{h}_{i-1/2}^{n+1,r}) p_{i+1}^{n+1,r+1} \\ &+ \left(\frac{\varepsilon \Delta x}{\gamma - 1} + \frac{1}{4} \frac{\Delta t^2}{\Delta x} (\tilde{h}_{i+1/2}^{n+1,r} + \tilde{h}_{i-1/2}^{n+1,r}) \right) p_i^{n+1,r+1} \\ &+ \frac{1}{4} \frac{\Delta t^2}{\Delta x} (\tilde{h}_{i+1/2}^{n+1,r} - \tilde{h}_{i-1/2}^{n+1,r}) p_{i-1}^{n+1,r+1} - \frac{1}{4} \frac{\Delta t^2}{\Delta x} \tilde{h}_{i-1/2}^{n+1,r} p_{i-2}^{n+1,r+1} = \varepsilon \Delta x b_i^r, \end{aligned} \quad (49)$$

where we have multiplied by ε each sides of the equations, and b_i^r is the known right hand side:

$$b_i^r = \left(F^e [E_i^n] - k_i^{n+1,r} \right) - \frac{\Delta t}{2\Delta x} \left(\tilde{h}_{i+1/2}^{n+1,r} F^e [q_{i+1}^n] + (\tilde{h}_{i+1/2}^{n+1,r} - \tilde{h}_{i-1/2}^{n+1,r}) F^e [q_i^n] - \tilde{h}_{i-1/2}^{n+1,r} F^e [q_{i-1}^n] \right). \quad (50)$$

The scope of the next Lemma is to prove that the system (49) admits an unique solution for all $\varepsilon > 0$.

Lemma 2. *Let us assume Dirichlet boundary conditions, then system (49) is well posed for all $\varepsilon > 0$.*

Proof. In the following, for simplicity, we omit all the exponents in the variables. System (49) can be rewritten for all $i = 1, \dots, N_x$ as

$$\frac{\varepsilon \Delta x}{\gamma - 1} p_i - \frac{\Delta t^2}{2 \Delta x} \left[\tilde{h}_{i+1/2} \left(\frac{p_{i+2} + p_{i+1}}{2} - \frac{p_i + p_{i-1}}{2} \right) - \tilde{h}_{i-1/2} \left(\frac{p_{i+1} + p_i}{2} - \frac{p_{i-1} + p_{i-2}}{2} \right) \right] = \varepsilon \Delta x b_i.$$

Multiplying this equation by $(p_i + p_{i-1})/2 + (p_{i+1} + p_i)/2$ and summing over the index i , gives

$$\begin{aligned} \frac{\varepsilon}{\gamma - 1} \sum_{i=1}^{N_x} \Delta x p_i \left(\frac{p_i + p_{i-1}}{2} + \frac{p_{i+1} + p_i}{2} \right) - \frac{\Delta t^2}{2 \Delta x} \left[\sum_{i=1}^{N_x} \tilde{h}_{i+1/2} \left(\frac{p_{i+2} + p_{i+1}}{2} - \frac{p_i + p_{i-1}}{2} \right) \left(\frac{p_i + p_{i-1}}{2} + \frac{p_{i+1} + p_i}{2} \right) \right. \\ \left. - \sum_{i=1}^{N_x} \tilde{h}_{i-1/2} \left(\frac{p_{i+1} + p_i}{2} - \frac{p_{i-1} + p_{i-2}}{2} \right) \left(\frac{p_i + p_{i-1}}{2} + \frac{p_{i+1} + p_i}{2} \right) \right] = \sum_{i=1}^{N_x} \varepsilon \Delta x b_i \left(\frac{p_i + p_{i-1}}{2} + \frac{p_{i+1} + p_i}{2} \right). \end{aligned}$$

Then, remarking that

$$\begin{aligned} p_i (p_i + p_{i-1} + p_{i+1} + p_i) &= p_i^2 + \frac{1}{2} p_i^2 + p_i p_{i+1} + \frac{1}{2} p_{i+1}^2 + \frac{1}{2} p_i^2 + p_i p_{i-1} + \frac{1}{2} p_{i-1}^2 - \frac{1}{2} p_{i+1}^2 - \frac{1}{2} p_{i-1}^2 \\ &= p_i^2 - \frac{1}{2} p_{i+1}^2 - \frac{1}{2} p_{i-1}^2 + (p_i + p_{i+1})^2 + (p_i + p_{i-1})^2, \end{aligned}$$

observing that $\sum_{i=1}^{N_x} (p_i^2 - \frac{1}{2} p_{i+1}^2 - \frac{1}{2} p_{i-1}^2) = 0$, and by changing the indexes in the third sum on the left hand side (using the Dirichlet boundary conditions), we get

$$\begin{aligned} \frac{\varepsilon}{2(\gamma - 1)} \sum_{i=1}^{N_x} \Delta x ((p_i + p_{i+1})^2 + (p_i + p_{i-1})^2) - \frac{\Delta t^2}{2 \Delta x} \left[\sum_{i=1}^{N_x} \tilde{h}_{i+1/2} \left(\frac{p_{i+2} + p_{i+1}}{2} - \frac{p_i + p_{i-1}}{2} \right) \left(\frac{p_i + p_{i-1}}{2} + \frac{p_{i+1} + p_i}{2} \right) \right. \\ \left. - \sum_{i=1}^{N_x} \tilde{h}_{i+1/2} \left(\frac{p_{i+2} + p_{i+1}}{2} - \frac{p_i + p_{i-1}}{2} \right) \left(\frac{p_{i+1} + p_i}{2} + \frac{p_{i+2} + p_{i+1}}{2} \right) \right] = \sum_{i=1}^{N_x} \varepsilon \Delta x b_i \left(\frac{p_i + p_{i-1}}{2} + \frac{p_{i+1} + p_i}{2} \right). \end{aligned}$$

Finally, relying on the Cauchy-Schwarz and Young inequalities and canceling the equal terms in the second and in the third sum, we obtain

$$\begin{aligned} \frac{\Delta x \varepsilon}{2(\gamma - 1)} \sum_{i=1}^{N_x} (p_i + p_{i+1})^2 + (p_i + p_{i-1})^2 - \frac{\Delta t^2}{2 \Delta x} \left[\sum_{i=1}^{N_x} \tilde{h}_{i+1/2} \left(\frac{p_{i+2} + p_{i+1}}{2} - \frac{p_i + p_{i-1}}{2} \right) \frac{p_i + p_{i-1}}{2} \right. \\ \left. - \sum_{i=1}^{N_x} \tilde{h}_{i+1/2} \left(\frac{p_{i+2} + p_{i+1}}{2} - \frac{p_i + p_{i-1}}{2} \right) \frac{p_{i+2} + p_{i+1}}{2} \right] \leq \frac{\Delta x \varepsilon (\gamma - 1)}{4} \sum_{i=1}^{N_x} |b_i|^2 + \frac{\Delta x \varepsilon}{4(\gamma - 1)} \sum_{i=1}^{N_x} ((p_i + p_{i-1})^2 + (p_{i+1} + p_i)^2). \end{aligned}$$

This gives

$$\frac{\Delta x \varepsilon}{2(\gamma - 1)} \sum_{i=1}^{N_x} (p_i + p_{i+1})^2 + \frac{\Delta t^2}{2 \Delta x} \sum_{i=1}^{N_x} \tilde{h}_{i+1/2} \left(\frac{p_{i+2} + p_{i+1}}{2} - \frac{p_i + p_{i-1}}{2} \right)^2 \leq \frac{\Delta x \varepsilon (\gamma - 1)}{4} \sum_{i=1}^{N_x} |b_i|^2. \quad (51)$$

Then, if $b_i = 0$, we obtain $p_i + p_{i+1} = 0$ and using the Dirichlet boundary condition, this gives $p_i = 0$ for all i and thus the well posedness of the system. \square

The system (49)-(50) can be written in compact matrix form as

$$\mathbf{T}^r \mathbf{p}^{n+1, r+1} = \varepsilon \Delta x \mathbf{b}^r, \quad (52)$$

with $\mathbf{p}^{n+1, r+1} = (p_1^{n+1, r+1}, \dots, p_{N_x}^{n+1, r+1})$ representing the vector of unknown pressures, and $\mathbf{b}^r = (b_1^r, \dots, b_{N_x}^r)$ the right hand side entries given by (50). The system matrix \mathbf{T}^r involves only positive quantities, namely pressure, enthalpy,

time step and mesh size and has a pentadiagonal form. In particular, if $\varepsilon = 0$, system (49) admits a constant in space pressure field which precise value depends on the chosen boundary conditions. This implies that the enthalpy is also constant in space. These two conditions finally give the incompressibility condition, i.e. $\partial_x u = 0$, and, the asymptotic consistency with the incompressible limit also at the fully discrete level. The system matrix \mathbf{T}^r is not proved to be symmetric, thus the pressure system is solved relying on the GMRES algorithm [97].

The initialization of all quantities defined at the first Picard iteration ($r = 1$) are taken as

$$p_i^{n+1,1} = (\gamma - 1) E_i^n - k_i^n, \quad q_i^{n+1,1} = F^e [q_i^n], \quad k_i^{n+1,1} = F^e [(\rho E)_i^n] - e_i^n. \quad (53)$$

Thus, the iterative Picard algorithm and the time evolution of the solution work in the following way. Equation (52) provides a new pressure $p_i^{n+1,r+1}$, then the momentum equation (44) provides a new value for $q_i^{n+1,r+1}$

$$q_i^{n+1,r+1} = F^e [(\rho u)_i^n] - \frac{1}{2} \frac{\Delta t}{\Delta x} (p_{i+1}^{n+1,r+1} - p_{i-1}^{n+1,r+1}), \quad (54)$$

from which we can update the enthalpy (46) and the kinetic energy (48). Typically, from our numerical experiments, a total number of $R_p = 2$ iterations are sufficient to obtain convergence of the nonlinear system (49)-(50). Therefore, after R_p Picard iterations, we set:

$$p_i^{n+1} = p_i^{n+1,R_p}, \quad q_i^{n+1} = q_i^{n+1,R_p}. \quad (55)$$

Once the solution at time index $(n + 1)$ in terms of pressure, enthalpy and momentum is found, then the total energy is updated thanks to the conservative equation (45). This concludes one time step of the proposed scheme.

Remark 1. *In cases of high Mach number flows, additional stability can be required. If this is the case, the centered discretization (44) employed for the pressure field can be supplemented with a numerical dissipation as done for the explicit operator $F[q_i^n]$ in (42), thus leading to a modified scheme for the momentum equation. This reads*

$$\bar{q}_i^{n+1} = q_i^{n+1} - \frac{1}{2} s_{i+1/2} (p_{i+1}^{n+1} - p_i^{n+1}) - \frac{1}{2} s_{i-1/2} (p_i^{n+1} - p_{i-1}^{n+1}), \quad (56)$$

where

$$s_{i+1/2} = (|u_{i+1/2}^n| + \max(c_{i+1}^n, c_i^n)) \frac{1}{\gamma - 1}, \quad s_{i-1/2} = (|u_{i-1/2}^n| + \max(c_i^n, c_{i-1}^n)) \frac{1}{\gamma - 1}. \quad (57)$$

The choice of the numerical dissipation in (57) is different from the one proposed in [90] and it makes use of a Rusanov-type scheme that can be very general and allows for applications to different systems of governing equations, thus not being limited only to the Euler equations [96].

3.3. Exact preservation of pressure and velocity across a contact discontinuity

In this part, we discuss the ability of our novel scheme to preserve a class of exact solutions of the compressible Euler equation. In particular, we study the capability of the fully discrete scheme to maintain a pressure equilibrium across a contact wave. This condition has been studied in [5], for instance, and it consists in verifying that a constant pressure and a constant velocity field are preserved through a discontinuity in the fluid density during the time evolution of the solution. To that aim, let us consider the following initial conditions for all $i = 1, \dots, N_x$:

$$\rho_i(t = 0) = \rho_i^0 = \begin{cases} \rho_L & x \leq x_D \\ \rho_R & x > x_D \end{cases}, \quad u_i(t = 0) = u_i^0 = u_0, \quad p_i(t = 0) = p_i^0 = p_0, \quad i = 1, \dots, N_x, \quad (58)$$

with x_D representing a given position inside the domain and $\rho_L \neq \rho_R$ being non-negative real numbers. In this case, one has for the first time step

$$F^e [\rho_i^0] = \rho_i^0 - \frac{\Delta t}{\Delta x} (f_{i+1/2}^{\rho,0} - f_{i-1/2}^{\rho,0}) = \rho_i^1, \quad (59)$$

$$F^e [q_i^0] = q_i^0 - \frac{\Delta t}{\Delta x} (f_{i+1/2}^{q,0} - f_{i-1/2}^{q,0}) = \rho_i^0 u_0 - \frac{\Delta t}{\Delta x} u_0 (f_{i+1/2}^{\rho,0} - f_{i-1/2}^{\rho,0}) = \rho_i^1 u_0, \quad (60)$$

$$F^e [E_i^0] = \underbrace{e_i^0 + k_i^0}_{=E_i^0} - \frac{\Delta t}{\Delta x} \frac{u_0^2}{2} (f_{i+1/2}^{\rho,0} - f_{i-1/2}^{\rho,0}) = e_i^0 + \rho_i^0 \frac{u_0^2}{2} - \frac{\Delta t}{\Delta x} \frac{u_0^2}{2} (f_{i+1/2}^{\rho,0} - f_{i-1/2}^{\rho,0}) = \frac{p_0}{\gamma - 1} + \rho_i^1 \frac{u_0^2}{2}. \quad (61)$$

Now, thanks to (59)-(61), one can observe that the equations (44) and (45) are solved by $p_i^{n+1} = p_0$ and $u_i^{n+1} = u_0$ thus guaranteeing the preservation of the exact solution. Indeed, under the above hypothesis we get from (44)

$$q_i^1 = F^e [q_i^0] - \frac{\Delta t}{2 \varepsilon \Delta x} (p_{i+1}^1 - p_{i-1}^1) = \rho_i^1 u_0, \quad i = 1, \dots, N_x, \quad (62)$$

and then by dividing the momentum q_i^1 at time level 1 by the density ρ_i^1 from (60) one obtains

$$u_i^1 = \frac{q_i^1}{\rho_i^1} = \frac{\rho_i^1 u_0}{\rho_i^1} = u_0, \quad i = 1, \dots, N_x. \quad (63)$$

Substituting this result into the energy equation (45) leads to

$$E_i^1 = \frac{p_0}{\gamma - 1} + \rho_i^1 \frac{u_0^2}{2} - \frac{\Delta t}{2 \Delta x} \frac{\gamma u_0}{\gamma - 1} (p_{i+1}^1 - p_{i-1}^1) = \frac{p_0}{\gamma - 1} + \rho_i^1 \frac{u_0^2}{2} = \frac{p_i^1}{\gamma - 1} + \rho_i^1 \frac{u_0^2}{2}, \quad (64)$$

which proves that $p_i^1 = p_0$ and $u_i^1 = u_0$ with $\rho_i^1 = \rho_i^0 - \frac{\Delta t}{\Delta x} (f_{i+1/2}^{\rho,0} - f_{i-1/2}^{\rho,0})$ are a solution of the proposed numerical scheme corresponding to a shift of the density profile at speed u_0 . Finally, being the structure of the solution preserved from time index $n = 0$ to $n = 1$, then the solution remains preserved for all $n \geq 1$. Numerical evidences of this property are reported in Section 5.2, where a moving contact wave at low Mach number is shifted in time exactly preserving a constant pressure and velocity background field.

3.4. Multi-dimensions extension

In this part, we discuss the multi-dimensional extension of the previously introduced numerical scheme. The computational domain $\Omega = [x_{\min}, x_{\max}] \times [y_{\min}, y_{\max}] \times [z_{\min}, z_{\max}]$ is paved with $N_x \times N_y \times N_z$ uniform cells of size $\Delta x \times \Delta y \times \Delta z$. A cell is labeled by three indices i, j, k , namely one for each direction, while a face is referred to as $i+1/2, j, k$, $i, j+1/2, k$ or $i, j, k+1/2$ and, any unit normal vector to a face is denoted by \mathbf{n} . The cell center is located at point $\mathbf{x}_{i,j,k} = (x_i, y_j, z_k)$ and a face center is at point $\mathbf{x}_{i+1/2,j,k} = \left(\frac{1}{2}(x_i + x_{i+1}), y_j, z_k \right)$. By definition we have the normal component of the velocity with respect to a face given by

$$\mathbf{u} \cdot \mathbf{e}_x = u, \quad \mathbf{u} \cdot \mathbf{e}_y = v, \quad \mathbf{u} \cdot \mathbf{e}_z = w, \quad (65)$$

where \mathbf{e} are the canonical unit vectors, that is $\mathbf{e}_x = (1, 0, 0)$, $\mathbf{e}_y = (0, 1, 0)$, $\mathbf{e}_z = (0, 0, 1)$. The numerical method in 3D follows the same path as its one dimensional version. Any generic cell-centered quantity $m_{i,j,k}$ is advected by a combination of an explicit and implicit flux. The explicit part reads

$$F^e [m_{i,j,k}^n] = m_{i,j,k}^n - \frac{\Delta t}{\Delta x} (f_{i+1/2,j,k}^m - f_{i-1/2,j,k}^m) - \frac{\Delta t}{\Delta y} (g_{i,j+1/2,k}^m - g_{i,j-1/2,k}^m) - \frac{\Delta t}{\Delta z} (\omega_{i,j,k+1/2}^m - \omega_{i,j,k-1/2}^m), \quad (66)$$

where the numerical fluxes are given by

$$\begin{aligned} f_{i+1/2,j,k}^m &= \frac{1}{2} (f(m_{i+1,j,k}^n) + f(m_{i,j,k}^n)) - \frac{1}{2} |u_{i+1/2,j,k}^{\max}| (m_{i+1,j,k}^n - m_{i,j,k}^n), \\ g_{i,j+1/2,k}^m &= \frac{1}{2} (g(m_{i,j+1,k}^n) + g(m_{i,j,k}^n)) - \frac{1}{2} |v_{i,j+1/2,k}^{\max}| (m_{i,j+1,k}^n - m_{i,j,k}^n), \\ \omega_{i,j,k+1/2}^m &= \frac{1}{2} (\omega(m_{i,j,k+1}^n) + \omega(m_{i,j,k}^n)) - \frac{1}{2} |w_{i,j,k+1/2}^{\max}| (m_{i,j,k+1}^n - m_{i,j,k}^n), \end{aligned}$$

with the speeds at the interfaces

$$|u_{i+1/2,j,k}^{\max}| = \max(|u_{i+1,j,k}^n|, |u_{i,j,k}^n|), \quad |v_{i,j+1/2,k}^{\max}| = \max(|v_{i,j+1,k}^n|, |v_{i,j,k}^n|), \quad |w_{i,j,k+1/2}^{\max}| = \max(|w_{i,j,k+1}^n|, |w_{i,j,k}^n|).$$

The density $\rho_{i,j,k}^n$ is then computed as

$$\rho_{i,j,k}^{n+1} = F^e [\rho_{i,j,k}^n], \quad (67)$$

while the three components of the momentum equation verify

$$(\rho u)_{i,j,k}^{n+1} = F^e[(\rho u)_{i,j,k}^n] - \frac{1}{2} \frac{\Delta t}{\varepsilon \Delta x} (p_{i+1,j,k}^{n+1} - p_{i-1,j,k}^{n+1}), \quad (68)$$

$$(\rho v)_{i,j,k}^{n+1} = F^e[(\rho v)_{i,j,k}^n] - \frac{1}{2} \frac{\Delta t}{\varepsilon \Delta y} (p_{i,j+1,k}^{n+1} - p_{i,j-1,k}^{n+1}), \quad (69)$$

$$(\rho w)_{i,j,k}^{n+1} = F^e[(\rho w)_{i,j,k}^n] - \frac{1}{2} \frac{\Delta t}{\varepsilon \Delta z} (p_{i,j,k+1}^{n+1} - p_{i,j,k-1}^{n+1}), \quad (70)$$

where one notices that, as for the one dimensional case, the pressure is taken implicitly using a Rusanov flux without adding any numerical viscosity. The energy equation is expressed by

$$\begin{aligned} E_{i,j,k}^{n+1} &= F^e \left[E_{i,j,k}^n \right] \\ &- \frac{1}{2} \frac{\Delta t}{\Delta x} \left(\tilde{h}_{i+1/2,j,k}^{n+1} \left((\rho u)_{i+1,j,k}^{n+1} + (\rho u)_{i,j,k}^{n+1} \right) - \tilde{h}_{i-1/2,j,k}^{n+1} \left((\rho u)_{i,j,k}^{n+1} + (\rho u)_{i-1,j,k}^{n+1} \right) \right) \\ &- \frac{1}{2} \frac{\Delta t}{\Delta y} \left(\tilde{h}_{i,j+1/2,k}^{n+1} \left((\rho v)_{i,j+1,k}^{n+1} + (\rho v)_{i,j,k}^{n+1} \right) - \tilde{h}_{i,j-1/2,k}^{n+1} \left((\rho v)_{i,j,k}^{n+1} + (\rho v)_{i,j-1,k}^{n+1} \right) \right) \\ &- \frac{1}{2} \frac{\Delta t}{\Delta z} \left(\tilde{h}_{i,j,k+1/2}^{n+1} \left((\rho w)_{i,j,k+1}^{n+1} + (\rho w)_{i,j,k}^{n+1} \right) - \tilde{h}_{i,j,k-1/2}^{n+1} \left((\rho w)_{i,j,k}^{n+1} + (\rho w)_{i,j,k-1}^{n+1} \right) \right). \end{aligned} \quad (71)$$

Then, by writing the total energy $E_{i,j,k}^{n+1}$ as the sum of internal $e_{i,j,k}^{n+1}$ and kinetic energy $k_{i,j,k}^{n+1}$, given by

$$e_{i,j,k}^{n+1} = \frac{p_{i,j,k}^{n+1}}{\gamma - 1}, \quad k_{i,j,k}^{n+1} = \frac{1}{2} \rho_{i,j,k}^{n+1} \left[\left(\frac{(\rho u)_{i,j,k}^{n+1}}{\rho_{i,j,k}^{n+1}} \right)^2 + \left(\frac{(\rho v)_{i,j,k}^{n+1}}{\rho_{i,j,k}^{n+1}} \right)^2 + \left(\frac{(\rho w)_{i,j,k}^{n+1}}{\rho_{i,j,k}^{n+1}} \right)^2 \right],$$

one recasts the energy equation by substituting the momentum equations (68)-(69)-(70) into (71) yielding

$$\begin{aligned} \frac{p_{i,j,k}^{n+1}}{\gamma - 1} &= F^e \left[E_{i,j,k}^n \right] - k_{i,j,k}^{n+1} \\ &- \frac{1}{2} \frac{\Delta t}{\Delta x} \tilde{h}_{i+1/2,j,k}^{n+1} \left[F^e \left[(\rho u)_{i+1,j,k}^n \right] + F^e \left[(\rho u)_{i,j,k}^n \right] - \frac{1}{2} \frac{\Delta t}{\varepsilon \Delta x} (p_{i+2,j,k}^{n+1} - p_{i,j,k}^{n+1} + p_{i+1,j,k}^{n+1} - p_{i-1,j,k}^{n+1}) \right] \\ &+ \frac{1}{2} \frac{\Delta t}{\Delta x} \tilde{h}_{i-1/2,j,k}^{n+1} \left[F^e \left[(\rho u)_{i,j,k}^n \right] + F^e \left[(\rho u)_{i-1,j,k}^n \right] - \frac{1}{2} \frac{\Delta t}{\varepsilon \Delta x} (p_{i+1,j,k}^{n+1} - p_{i-1,j,k}^{n+1} + p_{i,j,k}^{n+1} - p_{i-2,j,k}^{n+1}) \right] \\ &- \frac{1}{2} \frac{\Delta t}{\Delta y} \tilde{h}_{i,j+1/2,k}^{n+1} \left[F^e \left[(\rho v)_{i,j+1,k}^n \right] + F^e \left[(\rho v)_{i,j,k}^n \right] - \frac{1}{2} \frac{\Delta t}{\varepsilon \Delta y} (p_{i,j+2,k}^{n+1} - p_{i,j,k}^{n+1} + p_{i,j+1,k}^{n+1} - p_{i,j-1,k}^{n+1}) \right] \\ &+ \frac{1}{2} \frac{\Delta t}{\Delta y} \tilde{h}_{i,j-1/2,k}^{n+1} \left[F^e \left[(\rho v)_{i,j,k}^n \right] + F^e \left[(\rho v)_{i,j-1,k}^n \right] - \frac{1}{2} \frac{\Delta t}{\varepsilon \Delta y} (p_{i,j+1,k}^{n+1} - p_{i,j-1,k}^{n+1} + p_{i,j,k}^{n+1} - p_{i,j-2,k}^{n+1}) \right] \\ &- \frac{1}{2} \frac{\Delta t}{\Delta z} \tilde{h}_{i,j,k+1/2}^{n+1} \left[F^e \left[(\rho w)_{i,j,k+1}^n \right] + F^e \left[(\rho w)_{i,j,k}^n \right] - \frac{1}{2} \frac{\Delta t}{\varepsilon \Delta z} (p_{i,j,k+2}^{n+1} - p_{i,j,k}^{n+1} + p_{i,j,k+1}^{n+1} - p_{i,j,k-1}^{n+1}) \right] \\ &+ \frac{1}{2} \frac{\Delta t}{\Delta z} \tilde{h}_{i,j,k-1/2}^{n+1} \left[F^e \left[(\rho w)_{i,j,k}^n \right] + F^e \left[(\rho w)_{i,j,k-1}^n \right] - \frac{1}{2} \frac{\Delta t}{\varepsilon \Delta z} (p_{i,j,k+1}^{n+1} - p_{i,j,k-1}^{n+1} + p_{i,j,k}^{n+1} - p_{i,j,k-2}^{n+1}) \right]. \end{aligned} \quad (72)$$

As for the one dimensional case, in the above equation, the variables which are already known at time t^{n+1} are the density ρ_i^{n+1} , because the continuity equation is solved explicitly, and the explicit operators $F[(\rho u)_{i,j,k}^n]$, $F[(\rho v)_{i,j,k}^n]$ and $F[(\rho w)_{i,j,k}^n]$. The resulting system of equations is strongly nonlinear in the pressure unknown. To solve this system we rely again on a Picard algorithm to converge to the solution at time t^{n+1} by selecting only the pressure terms as

implicit in the iterative method. Following this strategy, the resulting nonlinear system is given by

$$\begin{aligned}
& -\frac{1}{4} \frac{\Delta t^2}{\Delta x^2} \tilde{h}_{i+1/2,j,k}^{n+1,r} p_{i+2,j,k}^{n+1,r+1} - \frac{1}{4} \frac{\Delta t^2}{\Delta y^2} \tilde{h}_{i,j+1/2,k}^{n+1,r} p_{i,j+2,k}^{n+1,r+1} - \frac{1}{4} \frac{\Delta t^2}{\Delta z^2} \tilde{h}_{i,j,k+1/2}^{n+1,r} p_{i,j,k+2}^{n+1,r+1} \\
& - \frac{1}{4} \frac{\Delta t^2}{\Delta x^2} (\tilde{h}_{i+1/2,j,k}^{n+1,r} - \tilde{h}_{i-1/2,j,k}^{n+1,r}) p_{i+1,j,k}^{n+1,r+1} - \frac{1}{4} \frac{\Delta t^2}{\Delta y^2} (\tilde{h}_{i,j+1/2,k}^{n+1,r} - \tilde{h}_{i,j-1/2,k}^{n+1,r}) p_{i,j+1,k}^{n+1,r+1} - \frac{1}{4} \frac{\Delta t^2}{\Delta z^2} (\tilde{h}_{i,j,k+1/2}^{n+1,r} - \tilde{h}_{i,j,k-1/2}^{n+1,r}) p_{i,j,k+1}^{n+1,r+1} \\
& + \left(\frac{\varepsilon}{\gamma-1} + \frac{1}{4} \frac{\Delta t^2}{\Delta x^2} (\tilde{h}_{i+1/2,j,k}^{n+1,r} + \tilde{h}_{i-1/2,j,k}^{n+1,r}) + \frac{1}{4} \frac{\Delta t^2}{\Delta y^2} (\tilde{h}_{i,j+1/2,k}^{n+1,r} + \tilde{h}_{i,j-1/2,k}^{n+1,r}) + \frac{1}{4} \frac{\Delta t^2}{\Delta z^2} (\tilde{h}_{i,j,k+1/2}^{n+1,r} + \tilde{h}_{i,j,k-1/2}^{n+1,r}) \right) p_{i,j,k}^{n+1,r+1} \\
& + \frac{1}{4} \frac{\Delta t^2}{\Delta x^2} (\tilde{h}_{i+1/2,j,k}^{n+1,r} - \tilde{h}_{i-1/2,j,k}^{n+1,r}) p_{i-1,j,k}^{n+1,r+1} + \frac{1}{4} \frac{\Delta t^2}{\Delta y^2} (\tilde{h}_{i,j+1/2,k}^{n+1,r} - \tilde{h}_{i,j-1/2,k}^{n+1,r}) p_{i,j-1,k}^{n+1,r+1} + \frac{1}{4} \frac{\Delta t^2}{\Delta z^2} (\tilde{h}_{i,j,k+1/2}^{n+1,r} - \tilde{h}_{i,j,k-1/2}^{n+1,r}) p_{i,j,k-1}^{n+1,r+1} \\
& - \frac{1}{4} \frac{\Delta t^2}{\Delta x^2} \tilde{h}_{i-1/2,j,k}^{n+1,r} p_{i-2,j,k}^{n+1,r+1} - \frac{1}{4} \frac{\Delta t^2}{\Delta y^2} \tilde{h}_{i,j-1/2,k}^{n+1,r} p_{i,j-2,k}^{n+1,r+1} - \frac{1}{4} \frac{\Delta t^2}{\Delta z^2} \tilde{h}_{i,j,k-1/2}^{n+1,r} p_{i,j,k-2}^{n+1,r+1} = \varepsilon b_i^r, \tag{73}
\end{aligned}$$

where the right hand side reads

$$\begin{aligned}
b_{i,j,k}^r &= (F[E_i^n] - k_i^{n+1,r}) \\
& - \frac{1}{2} \frac{\Delta t}{\Delta x} (\tilde{h}_{i+1/2,j,k}^{n+1,r} F^e[(\rho u)_{i+1,j,k}^n] + (\tilde{h}_{i+1/2,j,k}^{n+1,r} - \tilde{h}_{i-1/2,j,k}^{n+1,r}) F^e[(\rho u)_{i,j,k}^n] - \tilde{h}_{i-1/2,j,k}^{n+1,r} F^e[(\rho u)_{i-1,j,k}^n]), \\
& - \frac{1}{2} \frac{\Delta t}{\Delta y} (\tilde{h}_{i,j+1/2,k}^{n+1,r} F^e[(\rho v)_{i,j+1,k}^n] + (\tilde{h}_{i,j+1/2,k}^{n+1,r} - \tilde{h}_{i,j-1/2,k}^{n+1,r}) F^e[(\rho v)_{i,j,k}^n] - \tilde{h}_{i,j-1/2,k}^{n+1,r} F^e[(\rho v)_{i,j-1,k}^n]), \\
& - \frac{1}{2} \frac{\Delta t}{\Delta z} (\tilde{h}_{i,j,k+1/2}^{n+1,r} F^e[(\rho w)_{i,j,k+1}^n] + (\tilde{h}_{i,j,k+1/2}^{n+1,r} - \tilde{h}_{i,j,k-1/2}^{n+1,r}) F^e[(\rho w)_{i,j,k}^n] - \tilde{h}_{i,j,k-1/2}^{n+1,r} F^e[(\rho w)_{i,j,k-1}^n]). \tag{74}
\end{aligned}$$

As for the one dimensional case, the iterative system involves only positive quantities, namely pressure, enthalpy, time step and mesh size. In particular, if $\varepsilon = 0$, system (73) admits a constant in space pressure field as solution which value depends on the imposed boundary conditions. This implies that the enthalpy is also constant in space. These two conditions finally give the incompressibility condition, i.e. $\nabla_x \cdot \mathbf{u} = 0$, and the asymptotic consistency with the incompressible limit at the fully discrete level. The initialization of all quantities defined at the first step of the Picard iterative solver ($r = 1$) directly follows what already presented for the one dimensional case, given by (53). Thus, the iterative Picard algorithm works in the following way: equation (73) provides a new pressure $p_{i,j,k}^{n+1,r+1}$, then the momentum equations (68)-(70) give a new value for $(\rho u)_{i,j,k}^{n+1,r+1}$, $(\rho v)_{i,j,k}^{n+1,r+1}$, $(\rho w)_{i,j,k}^{n+1,r+1}$ from which we can update the enthalpy and the kinetic energy. Once the solution at time index $(n+1)$ in terms of pressure, enthalpy and momentum is found, then the total energy is updated thanks to equation (72). This concludes one time step of the proposed scheme.

Thanks to the semi-implicit discretization finally, the CFL condition is only dictated by the fluid velocity \mathbf{u} , and not by the sound speed c , hence

$$\Delta t \leq \text{CFL} \frac{\min(\Delta x, \Delta y, \Delta z)}{\max_{i,j,k} |2 \mathbf{u}_{i,j,k}^n|}. \tag{75}$$

4. A second order in time and space scheme for the Euler equations in all Mach regimes

In this part we discuss how the first-order accurate scheme proposed in the previous Section can be extended to a scheme with second order accuracy in time and space. We first discuss the time marching approach through a semi-discretization of the Euler equations (13) and we study its properties. In a second part, we give the details of the discretization of the space variables.

4.1. A second-order in time scheme

The second-order in time extension is based on an Implicit-Explicit (IMEX) Runge-Kutta approach [3, 71, 72, 31, 8, 32, 9, 6]. In this work, to improve the time accuracy of our first order method, we use the second-order Ascher, Ruuth and Spiteri [3] tableaux denoted in the following by ARS(2,2,2), where the arguments in the acronym

means that the explicit and the implicit schemes have both two stages (first and second number) and are second order accurate (third number), the same holds true for the global scheme. We recall that implicit methods of order higher than one for hyperbolic problems cannot be TVD [40] and this situation does not change when IMEX methods are employed [29]. In this work, as explained in the Introduction, we do not focus on this limit imposed on the high order schemes and we propose a second order scheme which, in principle, may not assure L^∞ stability. Possible remedies are postponed to future investigations. The Butcher tableaux of the ARS(2,2,2) discretization are detailed in Table 1, where $\beta = 1 - \sqrt{2}/2$ and $\alpha = 1 - 1/(2\beta)$. Remarking that $\alpha = \beta - 1$ and $1 - \alpha = 2 - \beta$, the following scheme for the

Explicit	0	0	0	0	Implicit	0	0	0	0
	β	β	0	0		β	0	β	0
	1	α	$1 - \alpha$	0		1	0	$1 - \beta$	β
		α	$1 - \alpha$	0		0	$1 - \beta$	β	

Table 1: Butcher tableaux for the ARS(2,2,2) time discretization. Left panel: explicit tableau. Right panel: implicit tableau. $\beta = 1 - \frac{\sqrt{2}}{2}$, $\alpha = \beta - 1$.

Euler equations (14) is obtained:

$$\frac{\mathbf{Q}^{n,\star} - \mathbf{Q}^n}{\Delta t} + \beta \nabla_x \cdot \mathcal{L}_i(\mathbf{Q}^{n,\star}) + \beta \nabla_x \cdot \mathcal{L}_e(\mathbf{Q}^n) = 0, \quad (76)$$

$$\frac{\mathbf{Q}^{n+1} - \mathbf{Q}^n}{\Delta t} + (\beta - 1) \nabla_x \cdot \mathcal{L}_e(\mathbf{Q}^n) + (2 - \beta) \nabla_x \cdot \mathcal{L}_e(\mathbf{Q}^{n,\star}) + (1 - \beta) \nabla_x \cdot \mathcal{L}_i(\mathbf{Q}^{n,\star}) + \beta \nabla_x \cdot \mathcal{L}_i(\mathbf{Q}^{n+1}) = 0. \quad (77)$$

The partition between $\mathcal{L}_i = (\mathbf{F}^i, \mathbf{G}^i, \mathbf{W}^i)$, the implicit operator, and, $\mathcal{L}_e = (\mathbf{F}^e, \mathbf{G}^e, \mathbf{W}^e)$, the explicit one is the same of the first order case. In the following, we furnish a formal proof that the scheme (76)-(77) provides a consistent and second order accurate discretization of the limit model (19)-(21) when $\varepsilon \rightarrow 0$.

Lemma 3. *We denote by $(\rho_\varepsilon^n(\mathbf{x}), \mathbf{u}_\varepsilon^n(\mathbf{x}), E_\varepsilon^n(\mathbf{x}))_{n \geq 1}$ the solution of the semi discrete numerical scheme (76)-(77) with the boundary condition $\mathbf{u}^n(\mathbf{x}) \cdot \mathbf{v}(\mathbf{x}) = 0$, $\forall \mathbf{x} \in \partial\Omega, \forall n \geq 0$. Then, if the initial data are well prepared, the semi discrete numerical scheme is asymptotically consistent and accurate, i.e. in the limit $\varepsilon \rightarrow 0$ it becomes a consistent and second order in time discretization of system (19)-(21) from $n \geq 1$. Moreover, from $n \geq 2$, the scheme becomes a consistent and second order time discretization of the limit system (19)-(21) independently on the initial data.*

Proof. We assume that, for all $n \geq 1$, all quantities $\rho_\varepsilon^n(\mathbf{x})$, $\mathbf{u}_\varepsilon^n(\mathbf{x})$, $p_\varepsilon^n(\mathbf{x})$ and $E_\varepsilon^n(\mathbf{x})$, solutions of (76)-(77), converge respectively towards some values $\rho^n(\mathbf{x})$, $\mathbf{u}^n(\mathbf{x})$, $p^n(\mathbf{x})$ and $E^n(\mathbf{x})$ when ε tends to 0. Under this hypothesis, multiplying by ε the momentum equations of (76) and letting ε tend to zero, gives $\nabla_x p^{n,\star}(\mathbf{x}) = 0$, i.e. the pressure at time level (n, \star) is constant. From the equation of state (28), we have

$$E_\varepsilon^{n,\star}(\mathbf{x}) = e_\varepsilon^{n,\star}(\mathbf{x}) + \frac{\varepsilon}{2} \rho_\varepsilon^{n,\star}(\mathbf{x}) |\mathbf{u}_\varepsilon^{n,\star}(\mathbf{x})|^2 = \frac{p_\varepsilon^{n,\star}(\mathbf{x})}{(\gamma - 1)} + \frac{\varepsilon}{2} \rho_\varepsilon^{n,\star}(\mathbf{x}) |\mathbf{u}_\varepsilon^{n,\star}(\mathbf{x})|^2,$$

which gives $E^{n,\star}(\mathbf{x}) = p^{n,\star}(\mathbf{x})/(\gamma - 1)$ in the limit $\varepsilon \rightarrow 0$. Since the explicit part of the energy equation depends on the kinetic term which goes to zero in the limit, for all $\mathbf{x} \in \Omega$ the limit yields

$$E^{n,\star}(\mathbf{x}) = E^n(\mathbf{x}) - \beta \Delta t (e^{n,\star}(\mathbf{x}) + p^{n,\star}(\mathbf{x})) \nabla_x \cdot \mathbf{u}^{n,\star}(\mathbf{x}). \quad (78)$$

Thus, integrating this equation on the domain, using the Green formula and the boundary condition gives

$$\begin{aligned} |\Omega| E^{n,\star}(\mathbf{x}) &= \int_\Omega E^n(\mathbf{x}) d\mathbf{x} - \beta \Delta t (e^n(\mathbf{x}) + p^n(\mathbf{x})) \int_{\partial\Omega} \mathbf{u}^n(\mathbf{x}) \cdot \mathbf{v}(\mathbf{x}) d\sigma(\mathbf{x}) + \\ &\quad - \beta \Delta t (e^{n,\star}(\mathbf{x}) + p^{n,\star}(\mathbf{x})) \int_{\partial\Omega} \mathbf{u}^{n,\star}(\mathbf{x}) \cdot \mathbf{v}(\mathbf{x}) d\sigma(\mathbf{x}) = \int_\Omega E^n(\mathbf{x}) d\mathbf{x}, \end{aligned} \quad (79)$$

which means that both pressure and energy are projected over constant states (\tilde{p}^0, E^0) at the intermediate stage level (n, \star) if initial data are well prepared, i.e. $E^0(\mathbf{x}) = E^0$. Using again (78), gives under the hypothesis of consistent

initial data $\nabla_x \cdot \mathbf{u}^{n,\star}(\mathbf{x}) = 0$, and

$$\frac{(\rho^{n,\star} \mathbf{u}^{n,\star})(\mathbf{x}) - (\rho^n \mathbf{u}^n)(\mathbf{x})}{\Delta t} + \beta \nabla_x \cdot \left((\rho^n \mathbf{u}^n)(\mathbf{x}) \otimes \mathbf{u}^n(\mathbf{x}) \right) + \beta \nabla_x p_1^{n,\star}(\mathbf{x}) = 0,$$

where we have assumed that the following limit for $\varepsilon \rightarrow 0$ exists

$$\lim_{\varepsilon \rightarrow 0} \frac{1}{\varepsilon} (p_\varepsilon^{n,\star} - p^{n,\star}) = \lim_{\varepsilon \rightarrow 0} \frac{1}{\varepsilon} \left(p_\varepsilon^{n,\star} - \frac{1}{|\Omega|} \int_{\Omega} \bar{p}^0(\mathbf{x}) d\mathbf{x} \right) = p_1^{n,\star}.$$

Otherwise, if initial data are not well prepared, the incompressibility condition is not in general verified. This means that for the first time step, the velocity field obeys to

$$\nabla_x \cdot \mathbf{u}^{0,\star} = - \frac{E^{0,\star} - E^0(\mathbf{x})}{\Delta t (e^{0,\star} + p^{0,\star})}. \quad (80)$$

We discuss now the second step of the Runge-Kutta method. We have again that multiplying by ε the momentum equations of (77) and letting ε tend to zero, gives $\nabla_x p^{n+1}(\mathbf{x}) = 0$, i.e. the pressure at time level $(n+1)$ is constant. The energy equation yields

$$E^{n+1}(\mathbf{x}) = E^n(\mathbf{x}) - (1 - \beta) \Delta t (e^{n,\star}(\mathbf{x}) + p^{n,\star}(\mathbf{x})) \nabla_x \cdot \mathbf{u}^{n,\star}(\mathbf{x}) - \beta \Delta t (e^{n+1}(\mathbf{x}) + p^{n+1}(\mathbf{x})) \nabla_x \cdot \mathbf{u}^{n+1}(\mathbf{x}). \quad (81)$$

Now, by integration in space the above equation and using the boundary conditions, one ends with

$$|\Omega| E^{n+1}(\mathbf{x}) = \int_{\Omega} E^n(\mathbf{x}) d\mathbf{x}, \quad (82)$$

which means that the energy is projected over a constant state in the limit $\varepsilon \rightarrow 0$. On the other hand, the pressure field is given, in the hypothesis of existence of the limit $\lim_{\varepsilon \rightarrow 0} \frac{1}{\varepsilon} (p_\varepsilon^{n+1} - p^{n+1})$, by

$$\begin{aligned} & \frac{(\rho^{n+1} \mathbf{u}^{n+1})(\mathbf{x}) - (\rho^n \mathbf{u}^n)(\mathbf{x})}{\Delta t} + (\beta - 1) \nabla_x \cdot \left((\rho^n \mathbf{u}^n)(\mathbf{x}) \otimes \mathbf{u}^n(\mathbf{x}) \right) + \\ & + (2 - \beta) \nabla_x \cdot \left((\rho^{n,\star} \mathbf{u}^{n,\star})(\mathbf{x}) \otimes \mathbf{u}^{n,\star}(\mathbf{x}) \right) + (1 - \beta) \nabla_x p_1^{n,\star}(\mathbf{x}) + \beta \nabla_x p_1^{n+1}(\mathbf{x}) = 0. \end{aligned} \quad (83)$$

Finally, we have that the incompressibility constraint is satisfied if initial data are well prepared

$$\nabla_x \cdot \mathbf{u}^{n+1} = 0, \quad \forall n \geq 1, \quad (84)$$

while this holds true only for $n \geq 2$ for non well prepared initial data because

$$\nabla_x \cdot \mathbf{u}^{n+1} = - \frac{E^{n+1} - E^n}{\Delta t (e^{n+1} + p^{n+1})} = 0, \quad n \geq 2, \quad (85)$$

and

$$\nabla_x \cdot \mathbf{u}^1 = - \frac{E^1 - E^0(\mathbf{x})}{\Delta t (e^1 + p^1)} \neq 0, \quad (86)$$

for the first time step, since energy is not necessarily constant at time $t = 0$. \square

4.2. Second-order in space

In order to extend the space accuracy to a second order, we use a piecewise linear reconstruction of the conserved variables supplemented with classical minmod limiter to avoid the development of spurious oscillations. For a cell $c := (i, j, k)$, the linear reconstruction polynomial $\mathbf{w}_h(x, y, z)$ writes

$$\mathbf{w}_{h,c}(x, y, z) = \mathbf{Q}_c + (x - x_c) \boldsymbol{\alpha}_{x,c} + (y - y_c) \boldsymbol{\alpha}_{y,c} + (z - z_c) \boldsymbol{\alpha}_{z,c}, \quad (87)$$

where $\mathbf{Q}_c = (\rho_c, q_c, E_c)$ is the vector of conserved mean values, $\mathbf{x}_c := (x_c, y_c, z_c)$ denotes the position of the centroid of cell c and $\mathbf{\Sigma}_c := (\alpha_{x,c}, \alpha_{y,c}, \alpha_{z,c})$ is the vector of slopes in x , y and z direction. Due to the fact that we use a three-dimensional Cartesian grid, the reconstruction is efficiently carried out dimension-by-dimension. Thus, let us consider the x direction with cell c and its direct Neumann neighbors $L := (i - 1, j, k)$ and $R := (i + 1, j, k)$, which are nothing but the so-called reconstruction stencil $\mathcal{S}_c := (L, c, R)$. The unknown reconstruction coefficient $\alpha_{x,c}$ is computed by applying a classical minmod limiter to the left $\alpha_{x,L}$ and right $\alpha_{x,R}$ slopes, that is

$$\alpha_{x,c} = \begin{cases} 0 & \text{if } \alpha_{x,L}\alpha_{x,R} \leq 0 \\ \alpha_{x,L} & \text{if } |\alpha_{x,L}| \leq |\alpha_{x,R}| \\ \alpha_{x,R} & \text{if } |\alpha_{x,L}| > |\alpha_{x,R}| \end{cases}, \quad (88)$$

where

$$\alpha_{x,L} = \frac{\mathbf{Q}_c - \mathbf{Q}_L}{\Delta x}, \quad \alpha_{x,R} = \frac{\mathbf{Q}_c - \mathbf{Q}_R}{\Delta x}. \quad (89)$$

Notice that the slopes are evaluated in (89) for each conserved variable of the Euler system contained in \mathbf{Q}_c . The same procedure applies for y and z direction independently, hence obtaining explicit values for all the reconstruction coefficients $\mathbf{\Sigma}_c \in \mathbb{R}$ in (87). The reconstruction is carried out only for all terms evaluated at time t^n because it is needed to feed the numerical fluxes of the explicit operators detailed in Section 3.2. Notice that all implicit terms are discretized with a centered second-order operator, which does not need any reconstruction procedure.

5. Numerical experiments

In this Section, we perform some numerical test cases which aim at demonstrating the accuracy and the robustness of our numerical method. A wide range of Mach numbers are considered, enhancing the capability of the pressure based implicit-explicit (P-IMEX) schemes to deal with strong shocks as well as smooth solutions. Specifically, we consider an isentropic vortex flow for studying the convergence, a set of Riemann problems with shock waves, including one Riemann problem that gives evidences about the asymptotic preserving property detailed in Sections 3.3 and 4.1. Finally, the Gresho vortex test problem is run at different Mach numbers and an explosion problem is performed using approximately 6 million cells. Comparison with a fully explicit method is also proposed. If not specified, in all the simulations shown hereafter, the CFL number is set to $CFL = 0.95$ in (40) and a total number of $R_p = 2$ has been used for the iterative solver (49). Finally, the rescaled parameter is set to $\epsilon = 1$ if not stated otherwise.

5.1. Isentropic vortex

The isentropic vortex problem was initially introduced in a two-dimensional setting [49] to test the accuracy of numerical methods since the analytical solution is known and regular. On the computational domain $\Omega = [-5; 5]^3$, the flow is characterized by $\rho_\infty = 1.0$, $u_\infty = 1.0$, $v_\infty = 1.0$, $w_\infty = 0.0$, $p_\infty = 1.0$, with a normalized temperature $T_\infty^* = 1.0$. A vortex centred on the z -axis line at $(x_c, y_c) = (0, 0)$ supplements the background gas at time $t = 0$ with the following conditions:

$$\rho = \rho_\infty + \delta\rho, \quad u = u_\infty + \delta u, \quad v = v_\infty + \delta v, \quad w = w_\infty, \quad T^* = T_\infty^* + \delta T^*, \quad (90)$$

where the perturbations for density and pressure read

$$\delta\rho = (1 + \delta T^*)^{\frac{1}{\gamma-1}} - 1, \quad \delta p = (1 + \delta T^*)^{\frac{\gamma}{\gamma-1}} - 1, \quad (91)$$

with the temperature fluctuation $\delta T^* = -\frac{(\gamma-1)\epsilon^2}{8\gamma\pi^2} e^{1-r^2}$. According to [49], the vortex strength is $\epsilon = 5$ and the adiabatic index is set to $\gamma = 1.4$, while the velocity field is affected by the following perturbations:

$$\begin{pmatrix} \delta u \\ \delta v \end{pmatrix} = \frac{\epsilon}{2\pi} e^{-\frac{1-r^2}{2}} \begin{pmatrix} -y' \\ x' \end{pmatrix}, \quad (92)$$

with $r = \sqrt{x'^2 + y'^2}$ and $x' = x - x_c$, $y' = y - y_c$. Periodic boundary conditions are prescribed and the final time is $t_{\text{final}} = 1$. The simulation has been run on a sequence of four successively refined Cartesian meshes on the xy plane

with characteristic mesh size $h(\Omega) = L_{char}/N_x$ with $L_{char} = 10$. The exact solution is the time-shifted initial condition at speed (u_∞, v_∞) , thus we compute the errors in L^2 norm for the density, velocity and pressure variables, and the numerical order of convergence can be deduced for both space and time discretization, simultaneously. Tables 2 and 3 report the data obtained for the first order and second order P-IMEX schemes, respectively. Comparison against the corresponding explicit algorithms are shown as well. The explicit methods are based on first order Godunov finite volume schemes and on MUSCL TVD algorithms for the second order case. The formal order of accuracy is achieved and the P-IMEX schemes are proven to be more accurate for order $O(2)$.

Table 2: Numerical convergence results for the compressible Euler equations using the first order P-IMEX scheme and comparison with the corresponding explicit method. The errors are measured in L_2 norm and refer to the variables ρ (density), u (horizontal velocity) and p (pressure) at time $t = 1$.

P-IMEX O1						
$h(\Omega)$	ρ_{L_2}	$O(\rho)$	u_{L_2}	$O(u)$	p_{L_2}	$O(p)$
5.00E-01	3.949E-01	-	1.149E-00	-	8.043E-01	-
3.82E-01	2.847E-01	1.2	8.140E-01	1.3	5.884E-01	1.2
3.15E-01	2.230E-01	1.3	6.308E-01	1.3	4.638E-01	1.2
2.71E-01	1.834E-01	1.3	5.152E-01	1.4	3.827E-01	1.3

EXPLICIT O1						
$h(\Omega)$	ρ_{L_2}	$O(\rho)$	u_{L_2}	$O(u)$	p_{L_2}	$O(p)$
5.00E-01	9.132E-01	-	2.139E-00	-	1.393E-01	-
3.82E-01	6.983E-01	1.0	1.601E-01	1.1	1.082E-01	0.9
3.15E-01	5.634E-01	1.1	1.277E-01	1.2	8.801E-01	1.1
2.71E-01	4.717E-01	1.2	1.062E-01	1.2	7.403E-01	1.2

Table 3: Numerical convergence results for the compressible Euler equations using the second order P-IMEX scheme and comparison with the corresponding explicit method (MUSCL-TVD discretization). The errors are measured in L_2 norm and refer to the variables ρ (density), u (horizontal velocity) and p (pressure) at time $t = 1$.

P-IMEX O2						
$h(\Omega)$	ρ_{L_2}	$O(\rho)$	u_{L_2}	$O(u)$	p_{L_2}	$O(p)$
5.00E-01	2.149E-01	-	4.565E-01	-	2.164E-01	-
3.82E-01	1.427E-01	2.0	2.559E-01	2.1	1.166E-01	2.3
3.15E-01	9.161E-02	2.3	1.673E-01	2.2	7.167E-02	2.5
2.71E-01	6.578E-02	2.2	1.184E-01	2.3	4.956E-02	2.5

EXPLICIT O2						
$h(\Omega)$	ρ_{L_2}	$O(\rho)$	u_{L_2}	$O(u)$	p_{L_2}	$O(p)$
5.00E-01	2.873E-01	-	6.620E-00	-	4.220E-01	-
3.82E-01	1.673E-01	2.0	3.966E-01	1.9	2.126E-01	2.5
3.15E-01	1.108E-01	2.1	2.621E-01	2.2	1.303E-01	2.6
2.71E-01	8.121E-02	2.1	1.848E-01	2.4	8.840E-02	2.6

5.2. Riemann problems in 1D

In this Section, we want to test the proposed numerical method against some classical Riemann problems whose initial states are summarized in Table 4. For all the tests we use a 3D grid made of $N_x = 200$, $N_y = 20$ and $N_z = 5$ cells on the computational domain $\Omega = [0; 1] \times [0; 0.1] \times [0; 0.1]$. Periodic boundaries are set in y - and z - direction, while Dirichlet conditions based on the initial state are imposed at $x = 0$ and $x = 1$. The Riemann problem RP1 and RP2, known as Sod and Lax problems, represent two benchmarks in gas dynamics [89]. They both are constituted by a

right-moving shock wave, a left-moving rarefaction fan and an intermediate contact discontinuity. The Mach number ranges from 0 to 1 for Sod problem and 0.17 to 0.87 for Lax. The results given by the P-IMEX schemes are depicted in Figures 2 and 3, respectively, where both first and second order simulations are considered.

Table 4: Initialization of 1d test cases.

Name	RP#	t_{final}	x_L	x_R	x_d	ρ_L	u_L	p_L	ρ_R	u_R	p_R
Contact	0	0.5	0	1	0.25	1000	0	10^5	0.01	0	10^5
Sod	1	0.2	0	1	0.5	1	0	1	0.125	0	0.1
Lax	2	0.14	0	1	0.5	0.445	1.698	3.528	0.5	0	0.571

The reference solution is well captured and the second-order P-IMEX scheme is more accurate, as expected. It is particularly visible for Lax test, where, for the first order scheme, the left rarefaction interacts with the boundary condition because the dissipation is too important (and the number of cells is not sufficiently high). The dissipation is reduced when using a second-order scheme, therefore such spurious interaction does not occur. Moreover the spreading of the shock wave and contact discontinuity is less pronounced with the second order version.

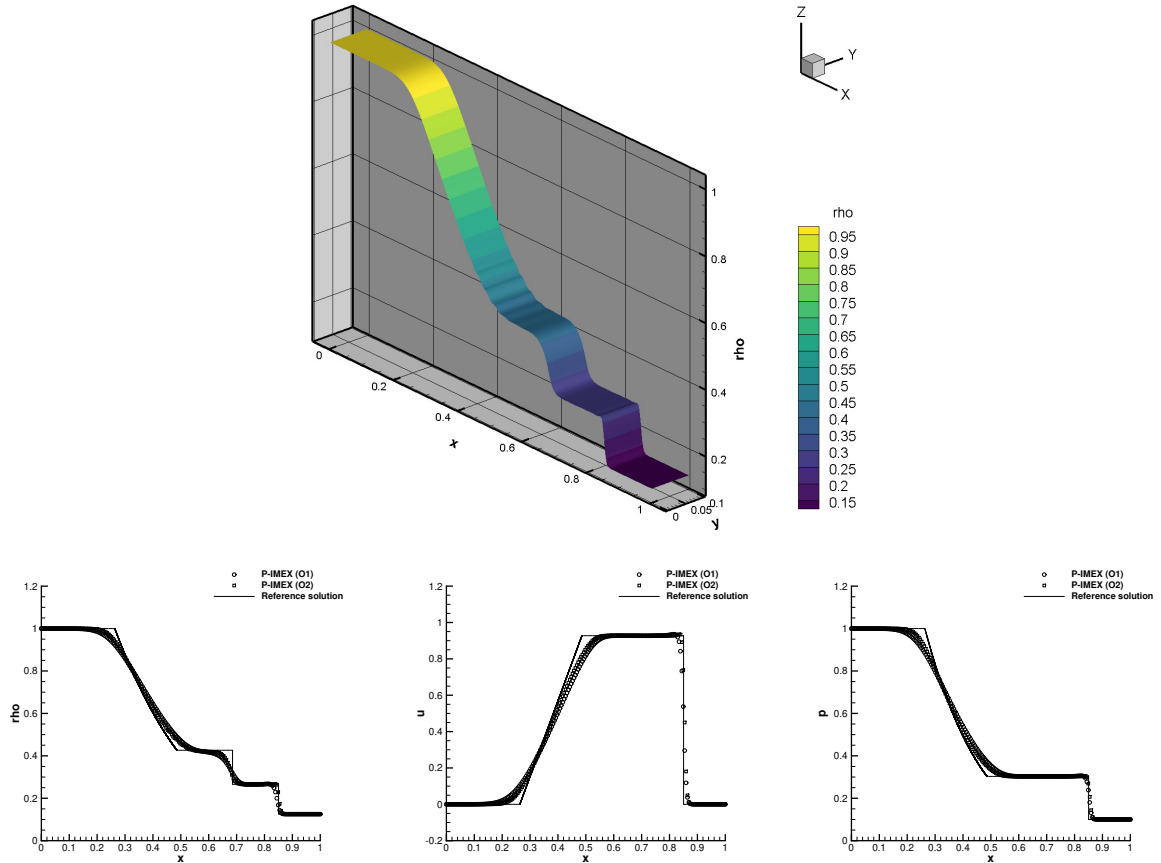


Figure 2: RP1 test case — Sod problem — 1st and 2nd order IMEX schemes — Top: 3d view of the density variable — Bottom (left to right): density, velocity and pressure (symbols) vs the exact solution (straight line).

Now, we consider the Riemann problem RP0 described in Table 4 where the initial density is a step-like function of five orders of magnitude, $\rho_L - \rho_R = 10^5$, $p_L = p_R = 10^5$, $u_L = u_R = u = 1$. The Mach number ranges from

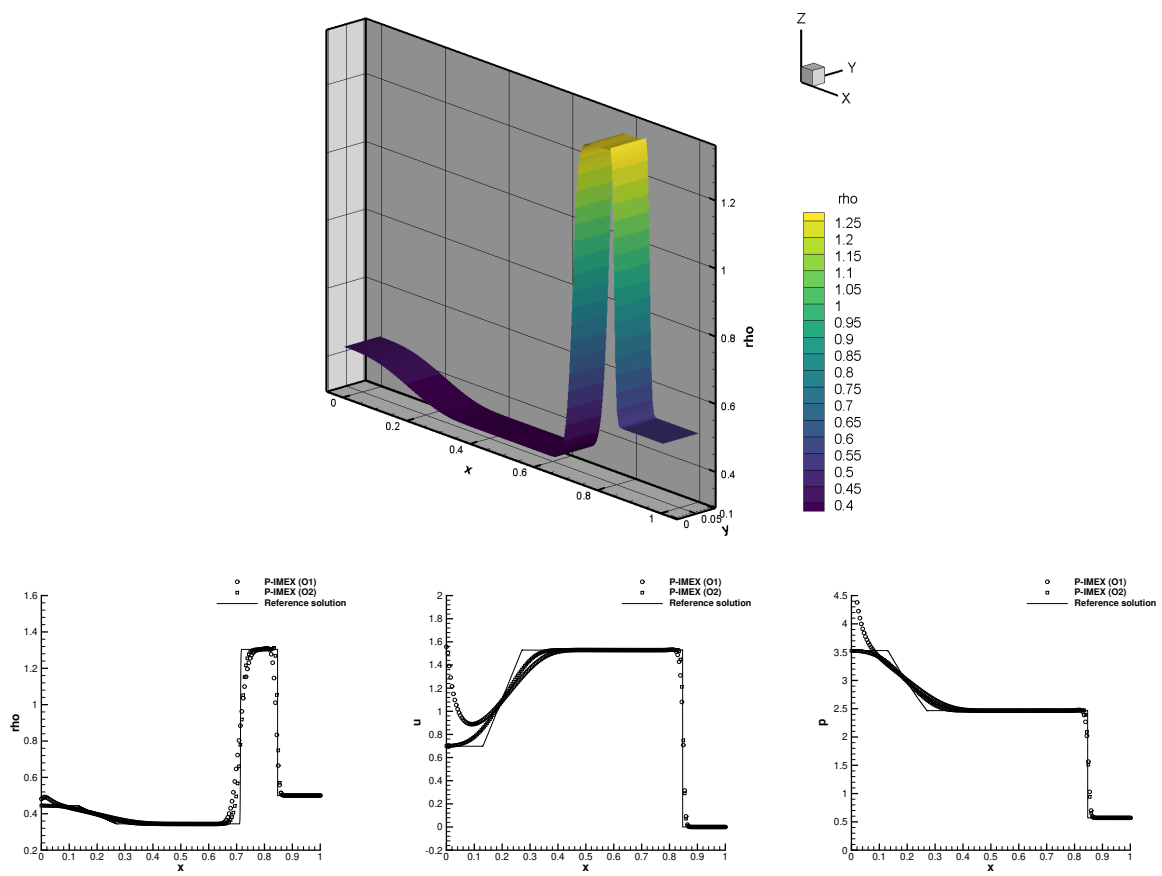


Figure 3: RP2 test case — Lax problem — 1st and 2nd order IMEX schemes — Top: 3d view of the density variable — Bottom (left to right): density, velocity and pressure (symbols) vs the exact solution (straight line).

approximately $2.7 \cdot 10^{-4}$ to $8.5 \cdot 10^{-2}$. In Figure 4 we present the results obtained with the first and second order P-IMEX schemes (density, velocity and pressure with symbols) versus the exact solution (straight line) on the bottom panels. On the top panel we present an extruded three-dimensional view where the color and azimuth scale referred to the density variable. As expected both schemes maintain a constant velocity and pressure, while the initial contact discontinuity is numerically dissipated. The second-order IMEX scheme produces a sharper contact discontinuity.

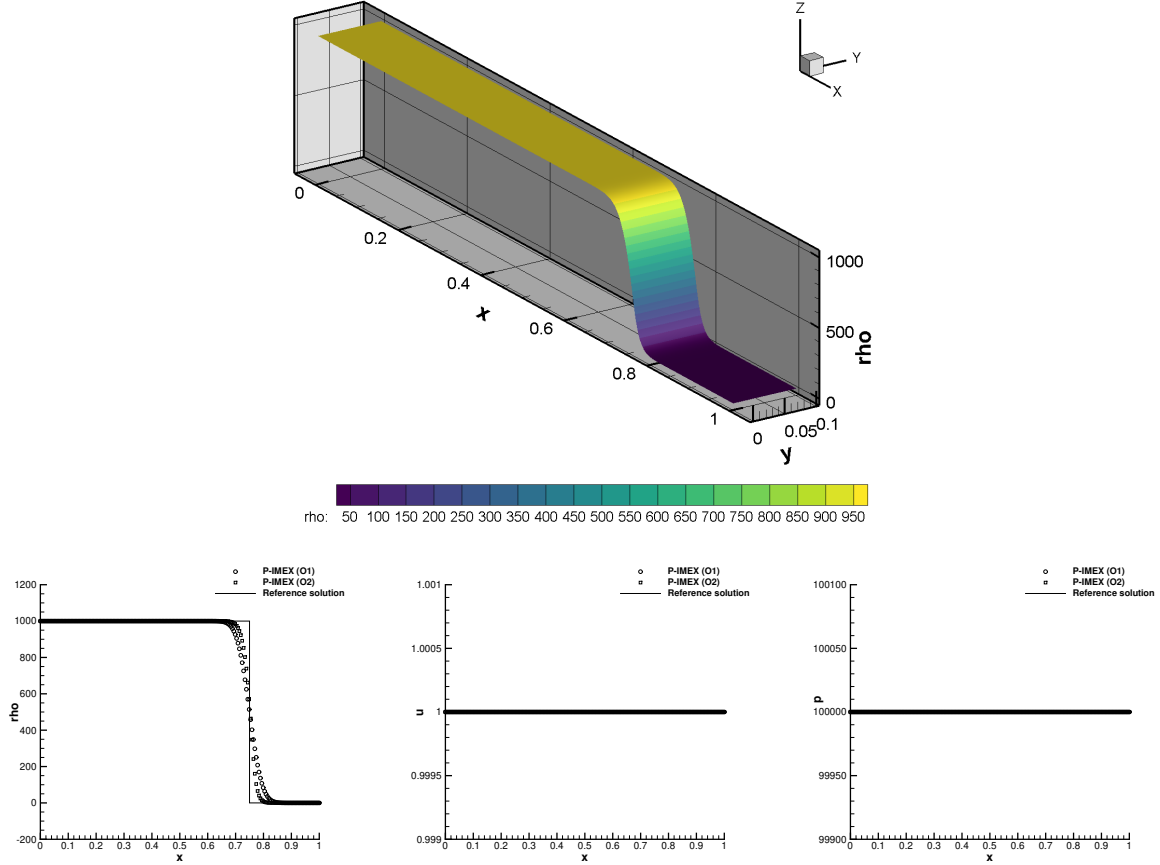


Figure 4: RP0 test case — Moving contact discontinuity — 1st and 2nd order IMEX schemes — Top: 3d view of the density variable — Bottom (left to right): density, velocity and pressure (symbols) vs the exact solution (straight line).

5.3. Shock tube problem at different Mach numbers

We test the behavior of the novel P-IMEX schemes for different values of the Mach number. To this aim, let us consider two regimes, namely $\epsilon = 10^{-2}$ and $\epsilon = 10^{-4}$. The initial condition for the shock tube problem is simply given by

$$(\rho, u, v, w, p) = \begin{cases} (1, 0, 0, 0, 1 + \epsilon) & x < x_s \\ (1, 0, 0, 0, 1) & x \geq x_s \end{cases}, \quad (93)$$

with $x_s = 0.5$. The computational domain is again the box $\Omega = [0; 1] \times [0; 0.1] \times [0; 0.1]$, with Dirichlet boundary conditions in x -direction and periodic boundaries imposed elsewhere. We always use $N_y = N_z = 5$ control volumes, while a total number of $N_x = 500$ and $N_x = 1000$ cells have been employed for $\epsilon = 10^{-2}$ and $\epsilon = 10^{-4}$, respectively. The shock tube problems are run with the following schemes: i) second order explicit (MUSCL-TVD), ii) second order P-IMEX and iii) first order in time and second order in space P-IMEX. A comparison for pressure and momentum distribution is depicted in Figure 5, where one can clearly notice the non-monotonicity of the second order implicit

P-IMEX scheme as expected. On the contrary, the first order P-IMEX method provides monotone results but highly diffusive as also expected. The total number of time steps as well as the computational time needed for carrying out the simulations are reported in Table 5. The lower is the Mach number, the more efficient becomes the P-IMEX scheme which only performs one time step to finish the computation. The linear system for the unknown pressure is solved using a matrix-free GMRES algorithm [97] without any preconditioner.

From this test we can conclude that the space limiter is not able to remove overshooting and undershooting of the numerical solution in the case of an implicit time discretization as figured out in [40]. Thus, the violation of the explicit CFL stability condition leads to spurious oscillations which remain bounded but do not disappear. A limiter for this part of the numerical method would be necessary and it will be addressed in the future.

Table 5: Computational time and number of time steps needed for the simulation of the shock tube problems with different schemes.

$\epsilon = 10^{-2}$			
Scheme	Order in time	CPU time [s]	time steps
Explicit	2	1.8125	144
Implicit	2	1.8281	1
Implicit	1	1.6406	1

$\epsilon = 10^{-4}$			
Scheme	Order in time	CPU time [s]	time steps
Explicit	2	5.9062	261
Implicit	2	3.3437	1
Implicit	1	2.9218	1

5.4. Gresho vortex

In this Section we solve the so-called Gresho vortex problem [17] that is a known stationary solution of the Euler equation. The solution, written in polar coordinates, reads

$$\begin{aligned}
 u_\theta(r) &= \begin{cases} 5r & 0 \leq r < 0.2 \\ 2 - 5r & 0.2 \leq r < 0.4 \\ 0 & r \geq 0.4 \end{cases} \\
 p(r) &= \begin{cases} p_0 + \frac{25}{2}r^2 & 0 \leq r < 0.2 \\ p_0 + \frac{25}{2}r^2 + 4[1 - 5r - \ln(0.2) + \ln(r)] & 0.2 \leq r < 0.4 \\ p_0 - 2 + 4\ln(2) & r \geq 0.4 \end{cases} \\
 \rho(r) &= 1
 \end{aligned} \tag{94}$$

where $u_\theta(r)$ is the angular velocity, $r = \sqrt{x^2 + y^2}$ is the radius and $p_0 = \rho/\gamma M^2$ is expressed in terms of the Mach number. The velocity field can be obtained from u_θ with a simple rotation, in particular $(u, v) = u_\theta/r \cdot (-y, x)$. In order to test the proposed algorithm against different magnitudes of the Mach number, we will use $M = 0.1, 0.01, 0.001$. The computational domain is $\Omega = [-0.5; 0.5] \times [-0.5; 0.5] \times [-0.05; 0.05]$, which is discretized with $N_x = N_y = 80$ and $N_z = 5$ control volumes. We impose the exact solution (94) as initial and boundary condition. We perform the simulation up to the final time $t_{final} = 0.4\pi$ using $CFL = 0.15$, in order to mimic the same setting adopted in [4]. The comparison between exact and numerical results obtained with the novel second order P-IMEX scheme for several Mach numbers is depicted in Figure 6 for $N_x = N_y = 80$. A very good agreement can be observed at all the values of the Mach numbers. The resulting dissipation of the total energy K normalized with the initial kinetic energy K_0 is reported in Figure 7 for all the Mach numbers and for mesh size $h_1 = 1/40$ and $h_2 = 1/80$. It is clear from Figure 7 that the obtained values of the kinetic energy dissipation do not depend on the chosen Mach number. The resulting profile is also in good agreement with others obtained in literature using an implicit shock capturing scheme, see [4], showing a slightly less dissipative behavior for our P-IMEX schemes.

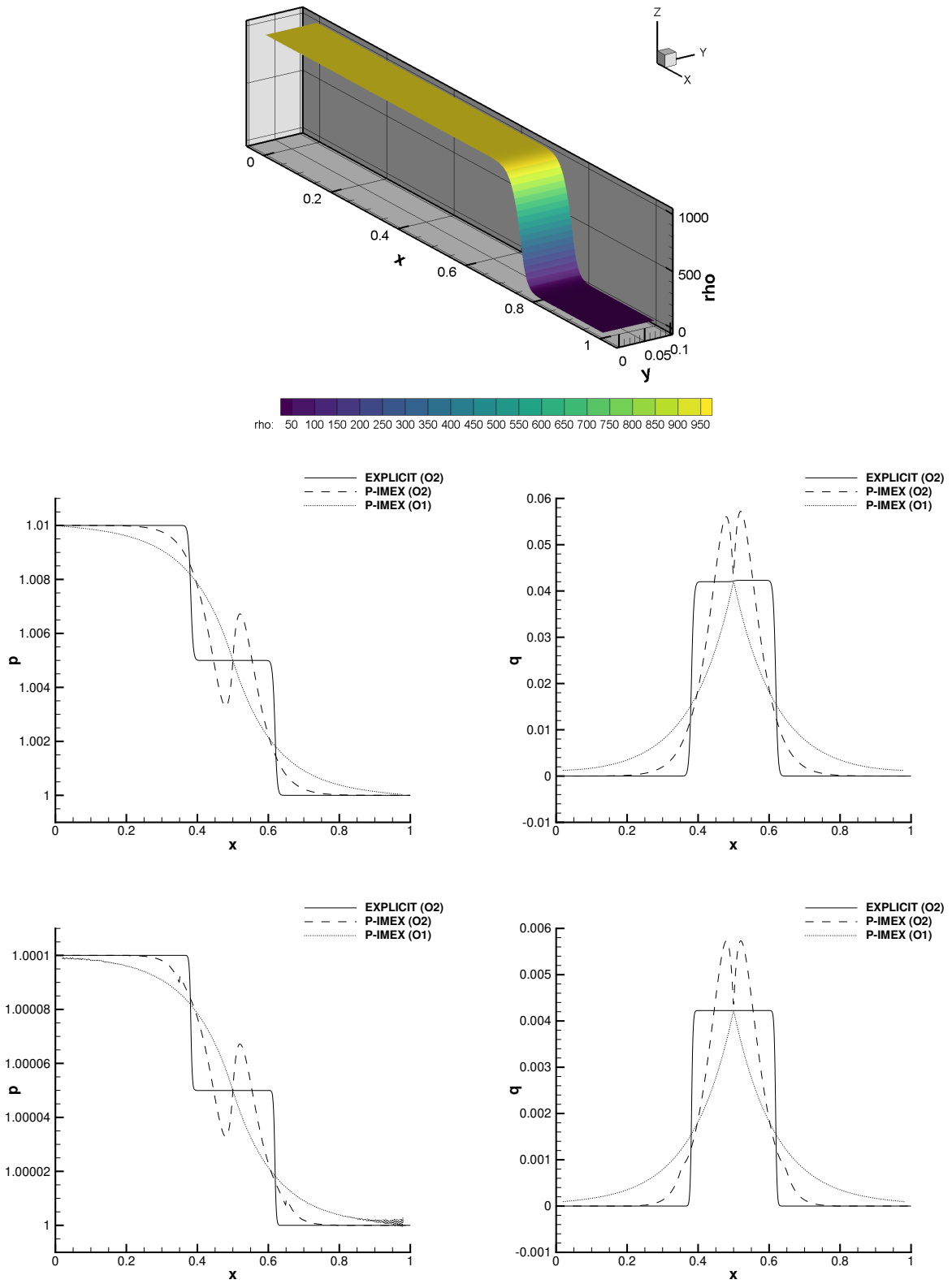


Figure 5: Shock tube problem with $\epsilon = 10^{-2}$ at time $t_{final} = 0.01$ (top row) and with $\epsilon = 10^{-4}$ at time $t_{final} = 0.001$ (bottom row). Pressure (left column) and momentum (right column) distribution are shown for the following schemes: second order explicit (solid line), second order P-IMEX (dashed line) and first order in time P-IMEX (dotted line).

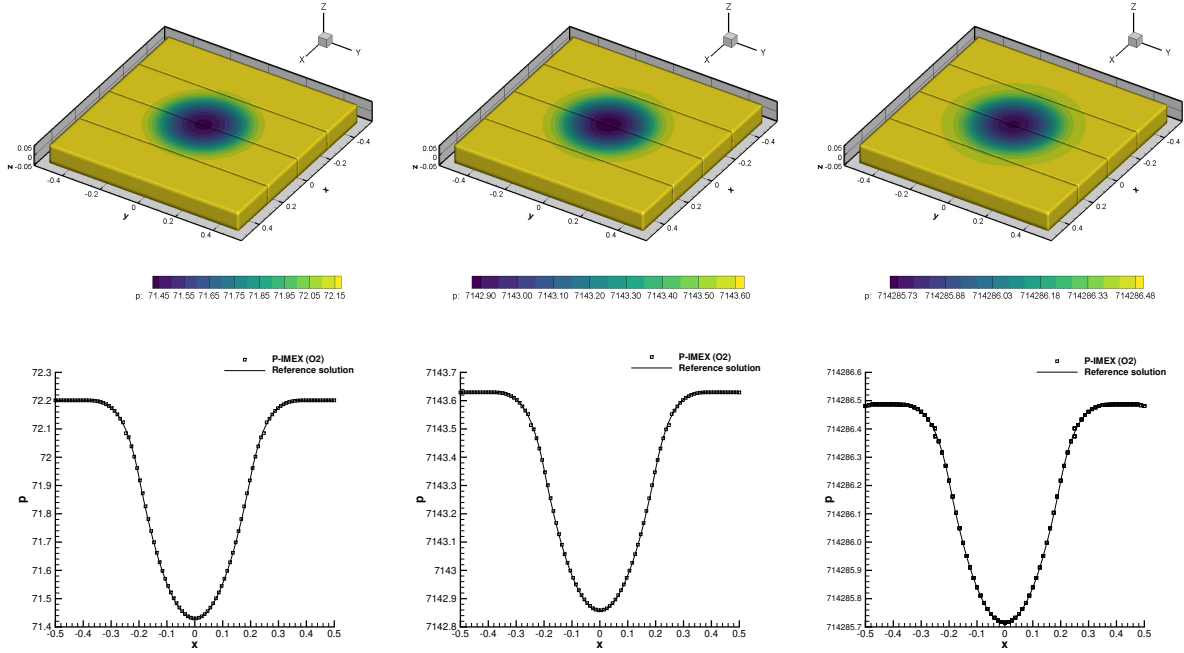


Figure 6: Gresho vortex problem at the final time $t_{final} = 0.4\pi$ with different Mach number regimes: $M = 0.1$ (left), $M = 0.01$ (middle), $M = 0.001$ (right). Top row: density contours. Bottom row: comparison between numerical and exact pressure distribution.

5.5. 3D Explosion problem

Finally, the proposed second order P-IMEX scheme is checked against the well known three-dimensional explosion problem [89]. This problem can be seen as the three-dimensional extension of the Sod problem already described in Section 5.2. The initial condition is given by

$$(\rho, u, v, w, p) = \begin{cases} (1, 0, 0, 0, 1) & r < R_s \\ (0.125, 0, 0, 0, 0.1) & r \geq R_s \end{cases}, \quad (95)$$

where $r = \sqrt{x^2 + y^2 + z^2}$ and $R_s = 0.5$ is the radius of the sphere. We consider $\Omega = [-1; 1]^3$ and Dirichlet boundary conditions everywhere. We use $N_x = N_y = N_z = 180$ elements so that the total number of elements is $N_E = 5'832'000$, as depicted in Figure 8. Furthermore $\gamma = 1.4$ and $t_{final} = 0.25$. We run the test in parallel with MPI using 64 AMD Opteron(TM) Processor 6272 CPU cores. The reference solution is obtained by considering the angular symmetry of the problem in order to reduce the 3D Euler system to a 1D system in radial direction with geometric source term, see [89]. The reference solution is actually computed employing a second order TVD scheme with Rusanov-type flux on 20000 radial points. The resulting numerical solution at the final time $t_{final} = 0.25$ is reported in Figure 9. In the first panel of 9 we plot some isolines for ρ in the semi-space $\{x < 0\}$ and show the radial symmetry of the obtained solution. This property is not trivial since the mesh is not radially aligned. Finally, we report the scatter plot of the main variables, considering all the 32'400 points in the slice $\{(x, y, z) \mid x = 0\}$. We observe a very good agreement with the reference solution also for this three-dimensional problem.

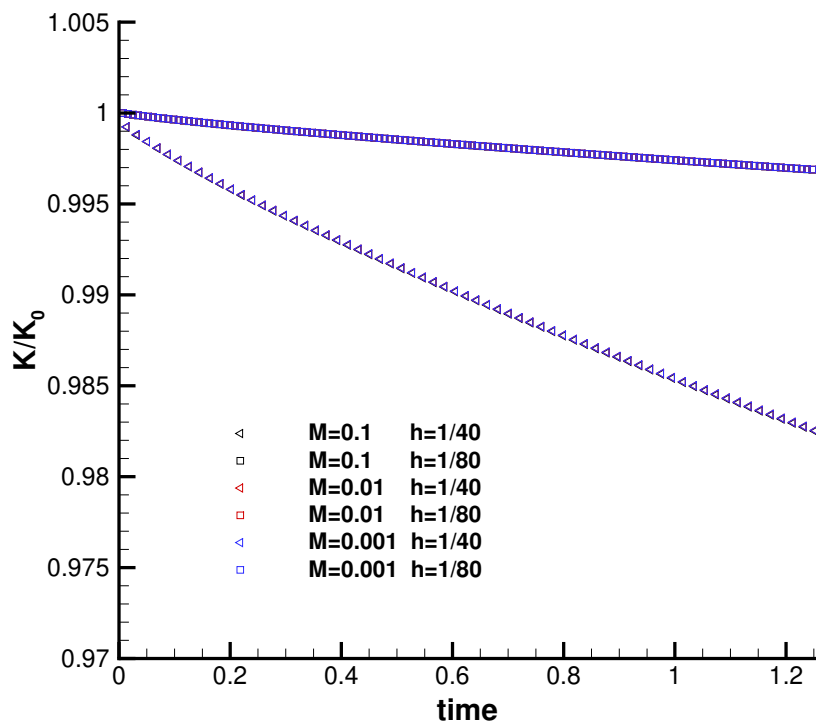


Figure 7: Evolution of the total kinetic energy K normalized with respect to the initial kinetic energy K_0 . Triangle scatters refer to the mesh size $N_x = N_y = 40$, while square scatters refer to the mesh size $N_x = N_y = 80$. Mach number 0.1 (black), 0.01 (red), 0.001 (blue) are considered.

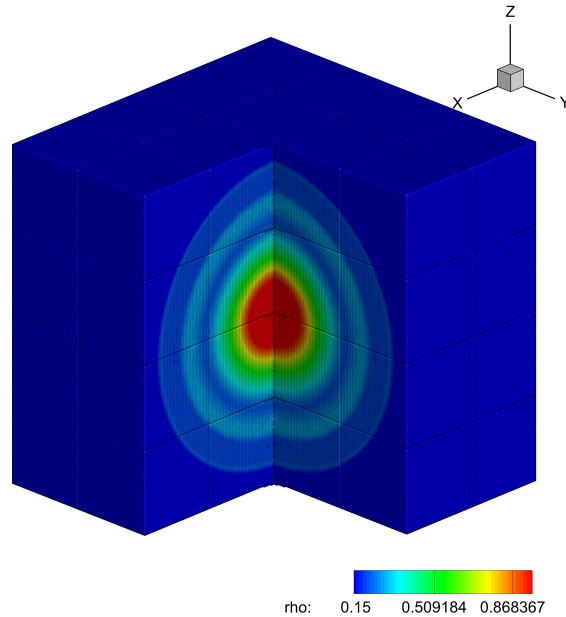


Figure 8: Computational mesh and density contours at time $t =$ for the 3D explosion problem. The computational domain is divided in 64 partitions corresponding to the MPI threads.

6. Conclusions and perspectives

In this article, a second order in space and time finite volume scheme for the solution of the compressible Euler equations in a wide range of Mach numbers has been discussed. The numerical method is based on the splitting of the fluxes into a non stiff and a stiff part. The first one corresponds to the advective contribution while the second one to the pressure part which may be responsible of creating very fast waves compared to the speed of the flow. These two different terms are then discretized, the first one explicitly while the second one implicitly. Second order of accuracy in space has been achieved relying on a piecewise linear reconstruction operator with classical minmod limiter for the explicit fluxes, while central finite difference discretization ensures the same accuracy for the implicit part as well. IMEX Runge-Kutta time stepping guarantees that the novel methods are fully second order accurate even in time.

For the proposed scheme, we showed that the asymptotic preserving property is satisfied and that the theoretical order of accuracy in the semi-discrete case for all values of the Mach number is maintained up to the incompressible limit. An analysis of the properties of the method has also been performed in the fully discretized case for the first order method.

The numerical results demonstrate that a wide spectrum of phenomena can be efficiently solved especially in the low Mach regime by our method, the stability depending only on the speed of the fluid flow and not on the sound speed. The method has been implemented on three-dimensional Cartesian grids and makes use of MPI parallelization. Demonstration of applicability is shown by a set of test problems that span from high to low Mach flows, including convergence studies.

In the future, we plan to extend the presented approach to the compressible Navier-Stokes equations in order to obtain an efficient solver for viscous fluids. Moreover, the development of higher order accurate methods which are consistent with the asymptotic incompressible limit of the Euler equations is the subject of future investigations. Finally, the study of L^∞ stable methods in the context of higher order time discretizations deserves to be analyzed in the next future.

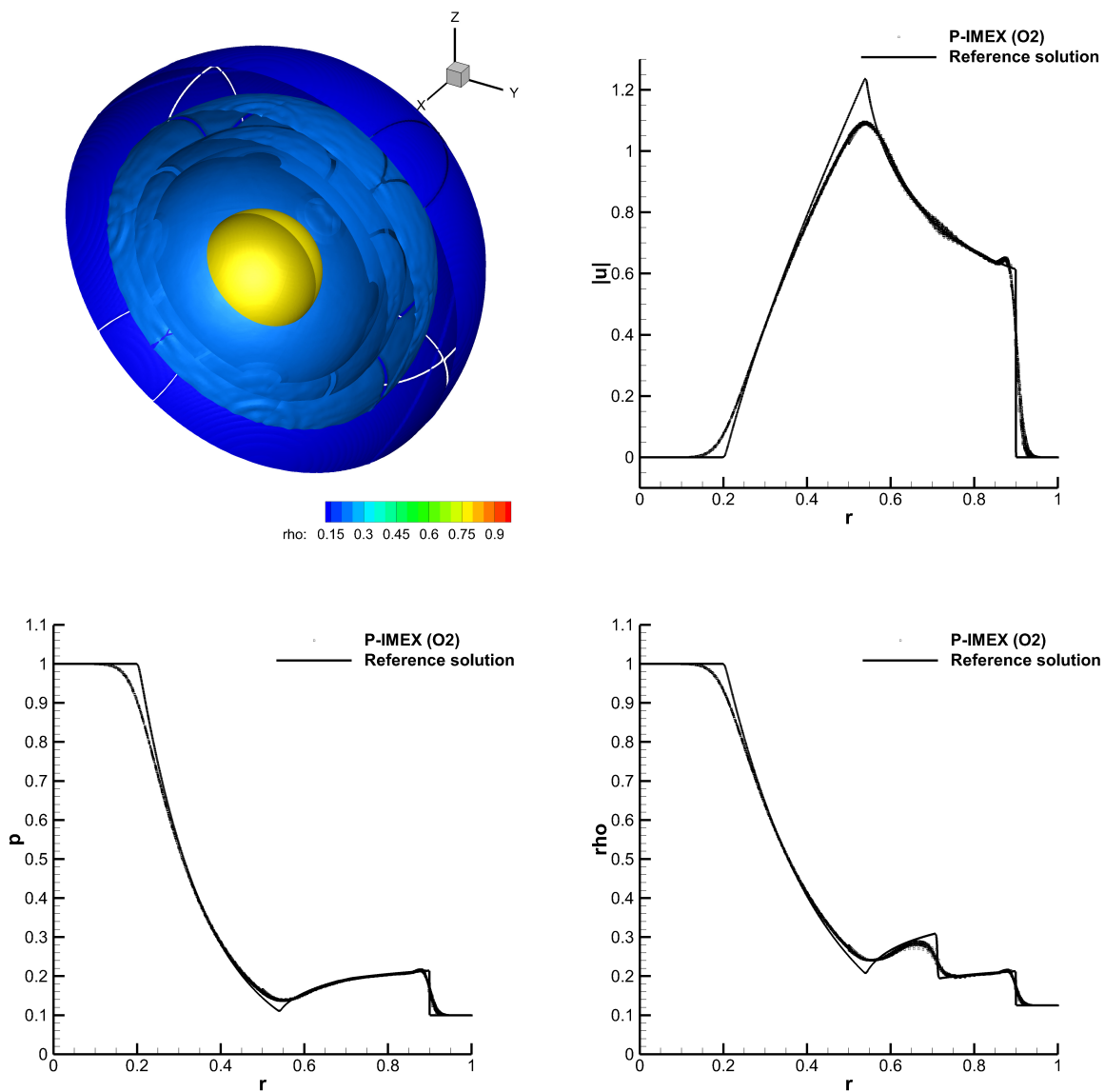


Figure 9: Scatter plot of the numerical solution for the velocity norm $|u| = \sqrt{u^2 + v^2 + w^2}$, pressure and density obtained with the new P-IMEX method at time $t_{final} = 0.25$ compared with the reference solution.

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